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Active Space Debris Removal by a Hybrid Propulsion Module

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Abstract

During the last 40 years, the mass of the artificial objects in orbit increased quite steadily at the rate of about 145 metric tons annually, leading to a total tally of approximately 7,000 metric tons. Now, most of the cross-sectional area and mass (97% in LEO) is concentrated in about 4,600 intact objects, i.e. abandoned spacecraft and rocket bodies, plus a further 1,000 operational spacecraft. Simulations and parametric analyses have shown that the most efficient and effective way to prevent the outbreak of a long-term exponential growth of the catalogued debris population would be to remove enough cross-sectional area and mass from densely populated orbits. In practice, according to the most recent NASA results, the active yearly removal of approximately 0.1% of the abandoned intact objects would be sufficient to stabilize the catalogued debris in low Earth orbit, together with the worldwide adoption of mitigation measures. The candidate targets for removal would have typical masses between 500 and 1,000 kg, in the case of spacecraft, and of more than 1,000 kg, in the case of rocket upper stages. Current data suggest that optimal active debris removal missions should be carried out in a few critical altitude-inclination bands. This paper deals with the feasibility study of a mission in which the debris is removed by using a hybrid propulsion module as propulsion unit. Specifically, the engine is transferred from a servicing platform to the debris target by a robotic arm so to perform a controlled disposal. Hybrid rocket technology for de-orbiting applications is considered a valuable option due to high specific impulse, intrinsic safety, thrust throttle ability, low environmental impact and reduced operating costs. Typically, in hybrid rockets a gaseous or liquid oxidizer is injected into the combustion chamber along the axial direction to burn a solid fuel. However, the use of tangential injection on a solid grain Pancake Geometry allows for more compact design of the propulsion unit. Only explorative tests were performed in the past on this rocket configuration, which appears to be suitable as de-orbiting system of new satellites as well as for direct application on large debris in the framework of a mission for debris removal. The paper describes some critical aspects of the mission with particular concern to the target selection, the hybrid propulsion module, the operations as well as the systems needed to rendezvous and dock with the target, and the disposal strategy.

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1. Introduction

Since the seminal work leaded by Donald Kessler in the 1970s on the artificial debris exponential growth [1], the publication of the position paper of the American Institute of Aeronautics and Astronautics (AIAA) in 1981 [2], the release of the report of the European Space Agency Space Debris Working Group in 1988 [3], the publication of the report of the Scientific and Technical Subcommittee of the United Nations Committee on the Peaceful Uses of Outer Space (UNCOPUOS) in 1999 [4], and the issuing of the position paper of the Space Debris Subcommittee of the International Academy of Astronautics (IAA) in 2001 [5], the international space community became progressively aware of the increasing relevance of the orbital debris problem, in order to guarantee the sustainable future use of the circumterrestrial space.

In addition to the voluntary adoption of some easy to implement mitigation measures by single space agencies, the need of addressing the problem on a global basis led to the creation of the Inter-Agency Space Debris Coordination Committee (IADC) [6]. The IADC is an international governmental forum for the worldwide coordination of activities related to the issues of man-made and natural debris in space. Its primary purposes are to exchange information on space debris research activities between member space agencies, to facilitate opportunities for cooperation in space debris research, to review the progress of ongoing cooperative activities, and to identify debris mitigation options. An important milestone was reached in 2002, when the IADC Steering Group approved the first version of a set of recommended space debris mitigation guidelines [7], intended to become a world standard for government and private entities active in space. In 2004 the IADC mitigation guidelines were basically incorporated in a code of conduct for space debris mitigation elaborated by the European Debris Mitigation Standards Working Group (EDMSWG) [8] and several standards discussed, or already approved, by the International Organization for Standardization (ISO) and by the UN International Telecommunication Union (ITU) tried to implement the IADC recommendations. Finally, in 2008, the United Nations endorsed a set of guidelines derived from the IADC ones [9]. It is therefore clear that the

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space debris issue is a global problem. Since the beginning the international cooperation to improve the knowledge on the subject and to adopt coordinated and cost-effective mitigation measures was of paramount importance. In the meantime, it became also clear that mitigation measures alone are probably not sufficient to avert the long-term onset of a debris exponential growth, possibly precluding the future use of certain orbital regimes (e.g. the sun-synchronous orbits), particularly popular today for many applications. For this reason, from a technical point of view, a clear distinction is made between orbital debris mitigation and remediation. Mitigation aims at reducing the generation of space debris through combined measures associated with the design, manufacture, operation, and disposal phases of a mission. Remediation, on the other hand, aims at managing the existing space debris population through debris removal, principally from the low Earth and geosynchronous protected regions [7, 10].

During the last 40 years, the mass of the artificial objects in orbit increased quite steadily at the rate of about 145 metric tons annually, leading to a total tally of approximately 7,000 metric tons [11]. Now, most of the cross-sectional area and mass is concentrated in about 4,600 intact objects, i.e. abandoned spacecraft and rocket bodies, plus a further 1,000 operational spacecraft. Simulations and parametric analyses have shown that the most efficient and effective way to prevent the outbreak of the Kessler syndrome, i.e. the long-term exponential growth of the cataloged debris population, would be to remove enough cross-sectional area and mass from densely populated orbits. In practice, the active yearly removal of approximately 0.1% of the abandoned intact objects would be sufficient to stabilize the cataloged debris in Low Earth Orbit (LEO) [12], together with the worldwide adoption of the mitigation measures recommended by the IADC and the United Nations [7, 9, 10]. The candidate targets for removal would have typical masses between 500 and 1,000 kg, in the case of spacecraft, and of more than 1,000 kg, in the case of rocket upper stages. All spacecraft in Earth orbit may experience hyper-velocity impacts from meteoroids and artificial orbital debris. Such impacts can occasionally result in the damage of critical systems, possibly leading to a mission loss. Several simulations and analyses, carried out since the 1970s, have come to the conclusion that this relatively manageable situation might dramatically worsen in a not so remote future, seriously jeopardizing the practical utilization of near-Earth space in selected altitude shells, already cluttered by abandoned intact objects and fragmentation debris [13]. Since the 1980s, and in particular during the last 20 years, the effort of the international space community was concentrated on the worldwide adoption of mitigation measures, able to reduce or prevent the production of new orbital debris. These measures include the passivation of satellites and upper stages at the end of their operational life to prevent accidental explosions, the choice of hardware and procedures to minimize the release of Mission Related Objects (MRO), the end-of-life removal of spacecraft from relatively crowded or important orbital regimes as the Geostationary Orbit (GEO), the orbits used by the telecommunications constellations (e.g. Iridium and Globalstar) in LEO, or the Medium Earth Orbits (MEO) used by the global navigation satellite systems (e.g. GPS and GLONASS), the limitation of the residual orbital lifetime of abandoned spacecraft and rocket bodies, and the prevention of accidental catastrophic collisions with conjunction assessments and, if needed, avoidance maneuvers.

During the last quarter of century, the progressive adoption of mitigation measures

was quite successful in putting under control the growth of cataloged orbital debris produced by on-orbit accidental fragmentations, but the recent Chinese anti-satellite test (2007), which destroyed the old Fengyun 1C spacecraft in the most crowded circumterrestrial region [14], and the catastrophic accidental collision among Iridium 33 and Cosmos 2251 (2009), basically in the same LEO critical orbit range [15], led to the production of a huge amount of new cataloged fragments, putting the mitigation clock back twenty years. Moreover, as pointed out by Donald Kessler in the 1970s, and later on confirmed by several teams of researchers around the world with progressively more detailed long-term simulations of the orbital debris evolution around the Earth, the artificial objects with sizes of 10 cm or more, i.e. those projectiles able to cause the catastrophic fragmentation of a typical spacecraft or rocket body at the average collision velocity in LEO of 10 km/s, might continue to grow, in certain altitude ranges, even if drastic measures, such as an immediate and complete halt of launches and on-orbit explosions, were enforced [16, 17, 18, 19, 20]. In fact, the fragments of collisional events among the objects already in space might drive the evolution of the environment over several decades, resulting in an exponential increase of the cataloged fragments able to cause further catastrophic collisions. A collisional cascading (Kessler syndrome) will finally follow, hampering any further space activity in certain altitude ranges [1, 21, 22]. For these reasons, it is being recognized that space debris mitigation alone might not be sufficient to guarantee the long-term utilization of some important orbital regimes. Therefore, some amount of remediation might be needed. Many authors have suggested and evaluated the use of electro-dynamic tethers applied to the objects to be disposed [23, 24, 25, 26], discussing also the potential benefits and risks of such a technology [27, 28, 29, 30, 31, 32]. Other de-orbiting or re-orbiting scenarios envisage the use of solid rocket motors [33], while ESA funded a study called ROGER to identify solutions to approach and capture non-operational satellites in GEO and tow them into a parking or graveyard orbit [34]. Melamed and Chobotov carried out a survey and evaluation of removal concepts using solar and magnetic sails, space tugs and several tether systems to compare their relative potential for mitigating the space debris crowding of the GEO region [35]. The conclusion was that the current practice of end-of-life re-orbiting is still the best. Comprehensive overviews of active debris removal proposals can be found in recent conference proceedings and study reports [36, 37, 38, 39].

Considering the LEO regions, in particular the altitude ranges more prone to the Kessler syndrome, the studies recently carried out indicated that, in addition to the generalized implementation of the IADC mitigation guidelines [7, 10], the most efficient and effective way to contrast the ignition of a debris chain reaction would be the removal of a few large mass intact objects per year from the 3-5 most crowded altitudes and inclination bands [12, 18, 19, 20, 34]. It was also shown that the removal of generic debris would lead to a growth reduction, but not to stabilization, because the reproduction of critical-size objects by collisions would more than balance the gain from removals [34].

Hybrid rocket technology for de-orbiting applications is considered a valuable option [40, 41] as discussed in detail in the section Propulsive Mission. The active debris removal by means of hybrid propulsion Modules (HPM) aims at achieving contact and control of large abandoned objects (typically spacecraft or launcher vehicle's upper stages), which have then to be removed thanks to a dedicated de-orbiting kit. The space platform in

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charge of this function may either be a large spacecraft for multiple targets or a smaller spacecraft for one single target.

2. Target Selection

Based on recent long-term simulation results [12, 18, 20, 42], a broad consensus exists among the space debris experts: the targets of active removal are large intact objects in crowded regions of space, since they are a potential source of numerous debris posing a collision risk. Generally the targets to be removed are ranked according to the following figure of merit: P_c (impact probability) $\times$ M (mass). However, the type of orbit and the estimated lifetime implicitly included in the estimation of the impact probability should also be considered in planning active debris removal missions. At the time of writing (19 July 2012), following 4,871 orbital launches, 3,638 intact payloads and 1,942 intact upper stages are in space, while 3,452 satellites and 3,603 rocket bodies have reentered in the atmosphere [43]. The current distribution of abandoned intact spacecraft and upper stages (Figures 1 and 2), together with the object ranking defined above [12], suggests that optimal active debris removal missions should be carried out in one of the following critical altitude (h) - inclination (i) bands:

(1) $h = 950 \pm 100 km, i = 82° \pm 1°$
(2) $h = 800 \pm 100 km, i = 99° \pm 1°$
(3) $h = 850 \pm 100 km, i = 71° \pm 1°$
(4) $h = 950 \pm 100 km, i = 65° \pm 1°$
(5) $h = 1000 \pm 100 km, i = 74° \pm 1°$
(6) $h = 750 \pm 100 km, i = 74° \pm 1°$
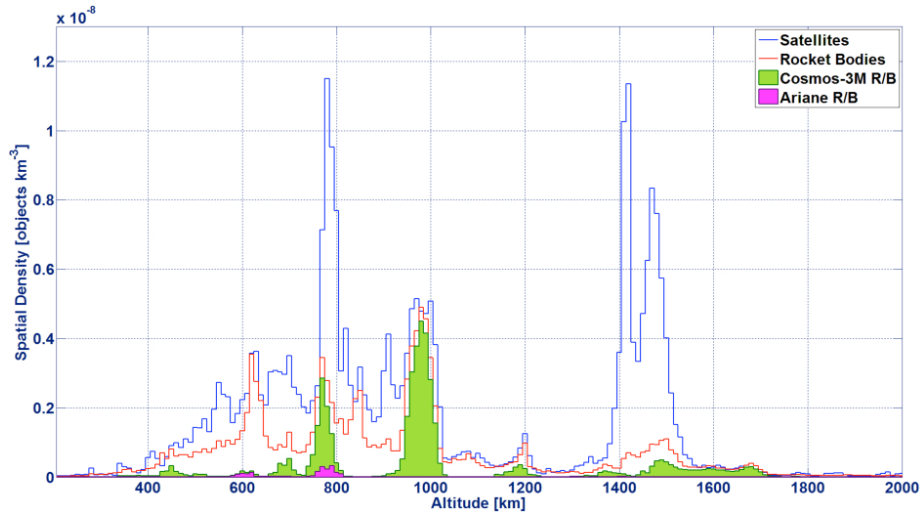(7) $h = 600 \pm 100 km, i = 82° \pm 1°$

Figure 1: Spatial density in LEO of intact satellites and rocket bodies (19 July 2012). The distribution of Cosmos-3M and Ariane upper stages is highlighted.

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The active debris removal from the first band might be very efficient, both for the high number of resident potential targets belonging just to four types (Cosmos-3M second stages, Vostok upper stages, Meteor and Parus satellites) and for the presence of a few objects in any 5° bin of right ascension of the ascending node (Ω), making possible, at least in principle, the removal of multiple targets with a single mission [42]. However, the long-term debris increase in other altitude regions cannot be suppressed by removing objects only from this altitude-inclination band [42].
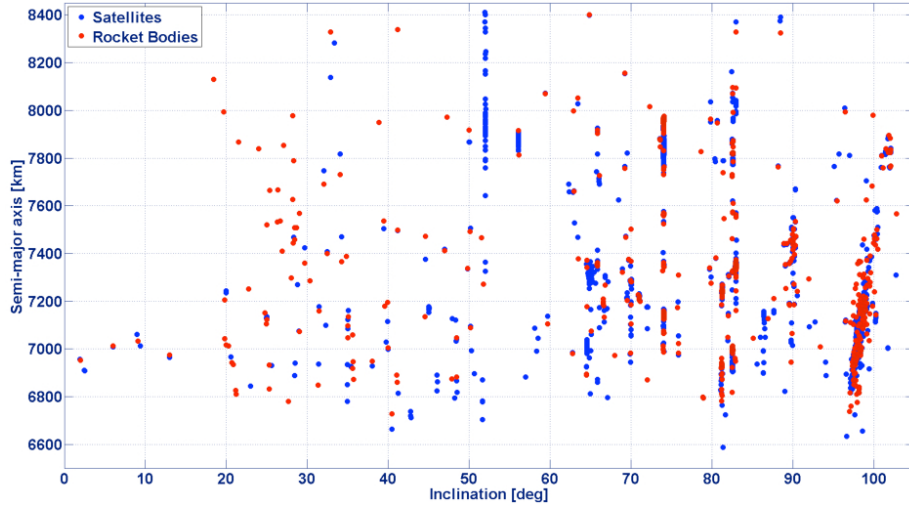

Figure 2: Distribution of intact satellites and rocket bodies fully residing in LEO as a function of semi-major axis and inclination (19 July 2012). The Earths equatorial radius is about 6378 km.

The removal of objects from the sun-synchronous regime (SSO), i.e. the second altitude-inclination band, is important, but less effective and should be anyway coupled with removal in the other critical regions as well [42]. Moreover, the population of intact objects present in SSO is quite heterogeneous and the presence of nearly coplanar targets with the same characteristics is not so frequent as in the first altitude-inclination band. The other altitude-inclination bands mainly include the following intact objects: Zenit-2 second stages and Tselina-2 satellites in the third; Cosmos-3M second stages, old US-A nuclear powered satellites with discarded nuclear reactor cores and DS-P1-M targets for anti-satellite weapon tests in the fourth; Cosmos-3M second stages and Tsiklon experimental navigation satellites in the fifth; Cosmos-3M second stages, Strela-2M and Tsiklon satellites in the sixth; and Tsiklon launcher upper stages and Tselina-D satellites in the seventh. In addition to the technological and economic aspects, the removal of space objects also presents subtle and not fully clarified legal facets. First of all, the launching states, as defined by the international law, retains jurisdiction over their objects in perpetuity, so any removal activity needs the approval of the object legal owner. Moreover, there is no clear liability definition and attribution for active debris removal attempts that go wrong, just to mention a not trivial problem which cannot be neglected. Because, at present, there are no appropriate Italian candidates for active removal, a feasibility study to apply the hybrid propulsion modules proposed by our team

should focus its attention on the following two options:

(1) European spacecraft and/or upper stages in SSO;
(2) LEO spacecraft and/or upper stages of a third cooperating party.

In general, spacecraft are more heterogeneous, fragile, complicated (in terms of shape, structure, appendages) and may pose confidentiality problems. Upper stages are easier and safer to grab, are less secretive and have simpler shapes, mass distributions, structures and rotational motions. Moreover, they belong to relatively few basic types, making possible many removal missions with basically the same docking and de-orbiting kit hardware. Taking into account the constraints just mentioned and the object removal effectiveness, three classes of potential targets have been considered so far:

(1) The ESAs ENVISAT satellite;
(2) The Ariane upper stages in LEO;
(3) The Russian Cosmos-3M upper stages.

In terms of $P_c \times$ M, ENVISAT, which suddenly failed in April 2012, is probably the worst unclassified object in space, with a mass of 8,050 kg and sizes of $25 \times 7 \times 10$ m. Placed into a sun-synchronous orbit of 766×768 km with an inclination of $97.5°$, it has now assumed a gravity gradient stabilization. On 21 January 2010, a CZ-2C rocket body (4 tons) missed the satellite by only 48 m, while, on 21 December 2010, an Iridium 33 fragment transited just 47 m away. In order to cause a catastrophic breakup of ENVISAT, a centered collision with a projectile with mass greater than 2.8 kg would be needed at the relative impact velocity of 15 km/s, quite common for objects in SSO, due to the actual debris distribution. The present probability of a catastrophic collisional breakup is therefore of the order of 0.06% per year, a not negligible value for a satellite with a residual lifetime of 100-150 years. Among the 122 Ariane upper stages in orbit, as of 19 July 2012, only 12 are entirely resident in LEO (Figure 3): 1 Ariane-1 H8 stage, with mass of 1,450 kg, diameter of 2.7 m and length of 10 m; 9 Ariane-4 H10 stages, with mass of 1,800 kg on average, diameter of 2.7 m and length of 12 m; and 2 Ariane-5 EPS stages, with mass of 3,600 kg on average, diameter of 5.4 m and length of 5 m. Except for the Ariane-4 H10 stage used to place into orbit the TOPEX/Poseidon satellite (i = 66°) in 1992, all the remaining 11 rocket bodies are in the SSO regime, with average altitudes in between 600 and 800 km and inclinations in the $98° - 99°$ range (Figure 3). Even though the Ariane upper stages in LEO are not, so far, a relevant component of the population of abandoned intact objects and present a limited growth potential, those in SSO, for the reasons previously mentioned, they might be good targets for Italian and/or European active removal demonstrative missions aiming at single objects.

Taking into account the LEO distribution of intact objects (Figure 1) and the collision risk ranking [12], a very attractive target for active removal is represented by the Russian Cosmos-3M second stages, with mass of 1,400 kg, diameter of 2.4 m and length of 6.5 m, of which 298 are in orbit as of 19 July 2012, mainly concentrated around two inclinations, $74°$ and $83°$ (Figure 4). In addition to their large number, they are significantly present in four critical altitude-inclination bands, i.e. the first (850-1,050 km, $i = 82°$), the fourth (850-1,050 km, i = 65°), the fifth (900-1,100 km, $i = 74°$) and the sixth one (650-850 km, $i = 74°$). The targeting of this upper stage presents quite evident advantages: among them, the same capture techniques and procedures might be used many times
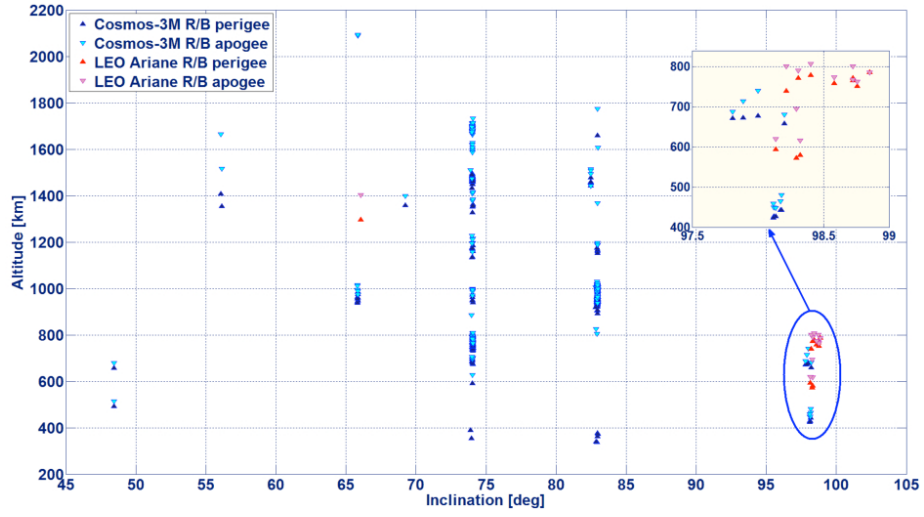
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Figure 3: Distribution of Ariane and Cosmos-3M rocket bodies (R/B) fully residing in LEO as a function of perigee/apogee altitude and inclination (19 July 2012).
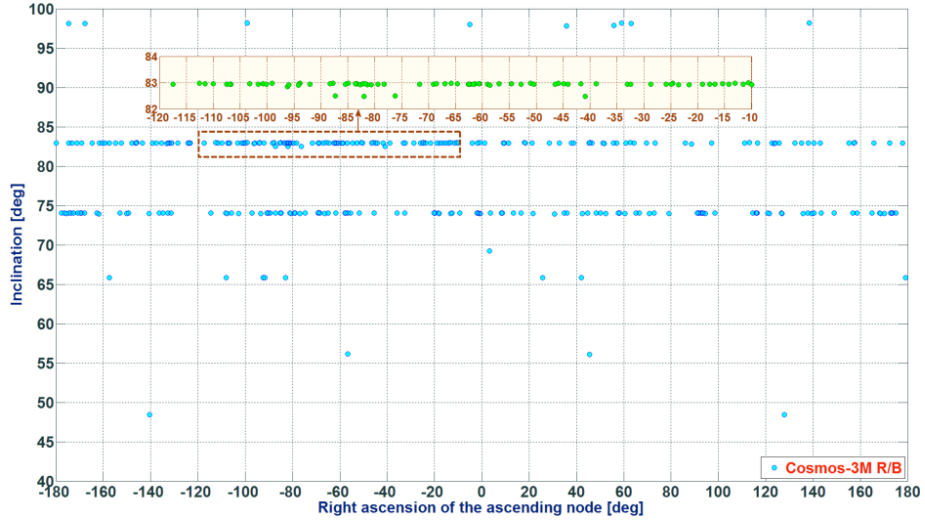

Figure 4: Dispersion of the orbital planes of the Cosmos-3M rocket bodies (R/B) in LEO (19 July 2012).

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over decades, it would be possible to operate in at least four separate altitude-inclination critical bands, the reentry risk assessment for de-orbiting (fragmentation analysis) should be carried out for only one object representative of the entire class, and the reduced set of de-orbiting kits needed might be tailored for small series production. In addition, multiple rendezvous might be possible within a single mission, because, for any given inclination, an average of about two stages would be present in each 5° bin of right ascension of the ascending node (RAAN), with more favorable concentrations around specific orbit planes (Figure 4). Last, but not least, the choice of the Cosmos-3M second stages as targets for active debris removal would offer the occasion for a broad cooperation with Russia, concerning both the rocket body itself (Omsk State Technical University) and the possible availability of launchers at low cost (Dnepr, Rokot) for the removal missions.

3. Mission Concept

In this section a preliminary concept for a removal demonstration mission is presented. Specific focus is on some critical issues affecting both mission performance and removal system design including:

- Mission profile;

- De-orbiting strategy;

- Debris target rendezvous;

- Capture and mating system;

- HPM sizing and de-orbiting capability.

3.1. Mission Profile

The removal mission under study relies on capturing a pre-selected large debris which is then de-orbited by exploiting the HPM. Once the removal platform has approached the target with a rendezvous maneuver, the debris capture is performed with a system that allows a safe and robust mating of removal platform and debris. The HPM is then detached from the removal platform with a robotic arm and installed on the debris in a convenient position with, in the case of nozzle, a specifically designed docking mechanism called corkscrew system [44]. Once the HPM has been safely docked, the capture system is deactivated and the target released so that the HPM can be remotely ignited for de-orbiting. The HPM is designed so to include all the avionics needed for debris attitude control, performed by RCS thrusters, during the de-orbiting.

3.2. Debris De-orbiting Strategy

Focusing the interest on large targets, the de-orbiting mission should be accomplished by steering the debris from its original orbit down to an altitude where it is supposed that a final impulse would direct it to a safe zone on Earth (typically, ocean regions). In particular for compact debris such as launcher upper stages, this strategy should present minimal risk since debris fragmentation due to interaction with the atmosphere should not change the overall trajectory and should keep the region of possible impact

with Earth sufficiently narrow. In this perspective, based on previous studies on de-orbiting strategies, a preliminary, non-optimized, trajectory analysis has been performed by assuming a multi-burn de-orbiting, an elliptical re-entry orbit with a perigee below 60 km and a flight path angle $< -1°$ at 120 km. To this end the debris is transferred from its initial orbit to a lower parking orbit, with altitude in the order of 350 km. After appropriate phasing on the parking orbit, a final impulse would decrease its perigee below 60 km altitude, from where the final descent to the ocean will be guided. Assuming, for simplicity, to adopt a two body model for the transfer and a Hohmann maneuver, to lower the orbit from 1,000 km to 350 km altitude the required velocity increment, ΔV, is about 350 m/s, while further 90 m/s are needed to lower the perigee below 60 km altitude. At this point, further 50 m/s should be considered for the final impulse to inject the debris into its final descent trajectory, Figure 5. This gives a preliminary estimate of total impulse of about 500 m/s, useful for the overall mission sizing.
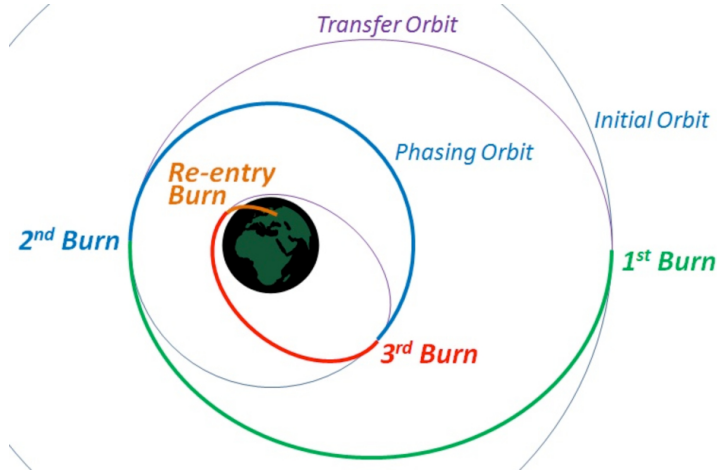

Figure 5: Debris disposal strategy scheme.

3.3. Rendezvous Strategy

It is assumed that the service platform is injected directly into the selected debris target orbit plane in a lower altitude parking orbit. By exploiting the different orbital periods, phasing with the target is achieved in order to start the rendezvous maneuver, which reduces to a few tens of kilometers the separation of the service platform from the target expected location. This last one may have an error up to 1-2 km due to uncertainty in ground tracking and available Two-Line Elements (TLE) data set, which, as well known, are updated at prefixed time intervals [45]. Before starting far/mid-range rendezvous, the actual position of the debris target shall be determined by using optical sensors (and IR sensors to guarantee continuous coverage also during eclipse conditions) on board the service platform, such as a far range camera or a star sensor. Specifically, at this stage the most important information coming from the far range sensor is the Line-Of-Sight (LOS) to the target, in order to correctly drive the approach maneuver. Indeed, in this phase, angle-only relative navigation can be performed, starting from a coarse a

priori relative orbit determination based on the knowledge of Debris TLE data and service platform absolute orbit from a GPS receiver. Optical systems also allow a preliminary positive identification of the target as the one to be removed. Technology for far/mid range rendezvous should not represent a critical issue for the mission. Indeed, relevant hardware and methodologies could be inherited from already flown space missions, like Orbital Express [46] and the more recent PRISMA [47], which demonstrated in flight autonomous rendezvous and docking starting from distances up to a few hundreds of kilometers.
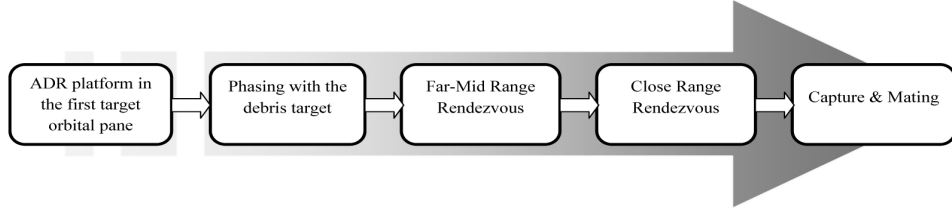

Figure 6: Mission phases up to target capture.

Based on the relative position information, the service platform can be maneuvered to gradually approach the target. Specifically, the far/mid-range rendezvous maneuver has to bring the service platform to a close proximity of the target to start close-range rendezvous and then target capture with the methodology and the system described in the following section. When the separation from the target reduces to about a few meters, close range rendezvous is started with the relative position and attitude (pose) of the target determined by using close range cameras and exploiting monocular or stereo-vision techniques [48]. Close-proximity relative navigation poses significant technology challenges, since pose determination techniques for non-cooperating targets shall be implemented. Techniques and algorithms capable of extracting natural features of the target with good invariance to lighting conditions (e.g. lines and edges), such as binarization, contour mapping, and edge detection, could be used to set up the synthetic information that will be used to determine the relative pose [49, 50, 51]. To this end, algorithms as target edge and 3D information matching relevant to monocular and binocular vision techniques, respectively, can be used. Also in this case, although to a smaller extent, hardware and methodologies could be inherited from already flown rendezvous technology demonstration missions [46, 47]. Before starting final approach for capture, a target fly around is performed for its final positive identification and inspection prior capture. This phase brings to the identification of the best points for capture, as well. Figure 6 sketches the several phases of the removal mission up to target capture. Instead Table 1 summarizes the several phases of the rendezvous maneuver.

3.4. Debris Capture and Mating

Other key aspects that have to be considered during the design process are the relative velocities of the two objects, the actual shape of the target and the local geometry of the debris surface. The proposed conceptual design is based on adhesive capture technologies and passive robotic joints. The traditional approach [52, 53, 54] to the capture problem suggests extremely accurate rendezvous and target inspection preliminary phases in order

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Phase	Operating	Sensors	Final range to target
Phasing	Target phasing Absolute navigation	GPS	~ 10 km
Far-Mid range rendezvous	Debris Target tracking and preliminary identification. Relative navigation.	Far-Mid range optical/IR cameras.	~ 10 m
Close-range rendezvous	Debris Target fly-around for identification and inspection. Close proximity navigation.	Close range optical/IR cameras.	~ 1 m

Table 1: Rendezvous phase summary.

to identify the debris angular rotation axis, to align the service satellite with it and to find a structural feature suitable for grasping. In this scenario an extremely precise attitude and orbital control of the chaser vehicle is necessary in addition to the robotic arm that holds the debris. The mentioned approach shows some weaknesses and is not necessarily the best for the considered scenario. The main issue is given by the need for suitable grasping points on the debris external surface. The connection points have to mechanically withstand to operation loads and need to be located in a convenient position with respect to the initial axis of rotation and center of mass of the object. As a matter of fact, the robotic arm has a limited work envelope therefore constraining the position of the grasping point. Furthermore, the position of the connection point determines the consequent manipulation forces and the requisites to the arm control system. The capability to position the capture system close to the target center of mass allows to reduce the inertial loads.
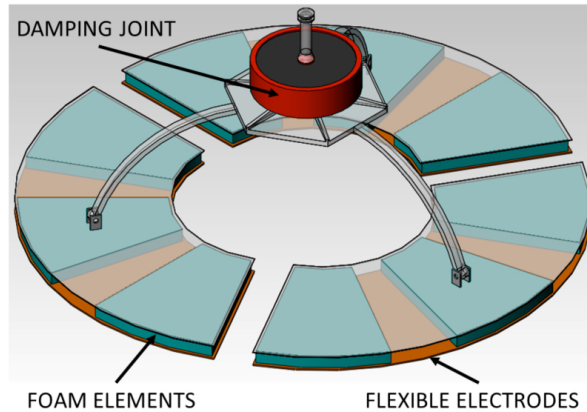

Figure 7: Capture mechanism conceptual drawing. The damping joint is placed at the interface with the robotic arm.

The proposed solution exploits adhesive capture technologies and, therefore, the core component of the system is the adhesion mechanism (see Figure 7 for a conceptual drawing). Electro-adhesion has been chosen to generate the required forces [55]. The working principle is based on electro-static attraction generated by an electric field. One advantage of this technology compared to other types of adhesives is the possibility to activate the system only when necessary with power consumption as a drawback. The forces are proportional to the available surface area and are influenced by the applied voltage, by the friction coefficient at the interface and by the quality of the contact. The local unknown irregularities on the object and the surface roughness can determine the presence of vacuum gaps reducing the effective contact surface. For this reason, the use of flexible electrodes mounted on a deformable material substrate is proposed. Polymeric foams can be employed to guarantee a better adaptability and adhesion between the interface surfaces. On top of the adaptable foam substrate, the adhesive material is applied. The latter is composed by flexible electrodes insulated between a double polypropylene layer which is characterized by a high dielectric constant. The resulting compliant surface in combination with the flexible foam guarantees a high level of adaptability that guarantees a better contact between the mating system and the debris surface. A secondary component of the system is the low rigidity passive damping joint. The main function of this device is to reduce impact forces and to dissipate the relative velocities and oscillations between the debris and the chaser vehicle after contact. The joint is based on elastomeric elements whose deformation determines internal energy dissipation. The joint can be designed so that it allows up to 6 degrees of freedom in a limited deflection range. The rigidity and damping performance of this component depend on material selection and geometry; a prototype version of this component is expected to present a flectional rigidity of 1 Nm/rad and a radial rigidity of 30 kN/m, while the damping ratio can reach the value of 0.2.

Considering a debris of about 2,000 kg, a preliminary sizing of the capture mechanism (for a 2 m^2 electro-adhesive surface) shows a peak power requirement of 10 W and an overall mass requirement of about 20 kg. Starting from the contact time instant, the capture procedure can be split into four phases (see Figure 8):

1. Impact;
2. Relative motion damping;
3. De-tumbling;
4. Attitude stabilization.

At the beginning of the capture sequence (1), the adhesive material is activated and put in contact with the debris surface. The polymeric foam substrate adapts to the local features of the target debris and guarantees a high quality contact at the interface. In less than a second the attraction force is established and the two bodies are connected. Preliminary estimations show that attraction pressures up to 10 kPa normally and up to 4 kPa in shear are feasible suggesting also the compatibility with large objects manipulation.

The initial relative motion of the objects determines impact forces that stress the adhesion interface and are transmitted to the service satellite. The adhesion mechanism has to generate a force larger than the impact loads in order to securely hold the debris while the spacecraft structure has to withstand to transmitted loads. In this phase the
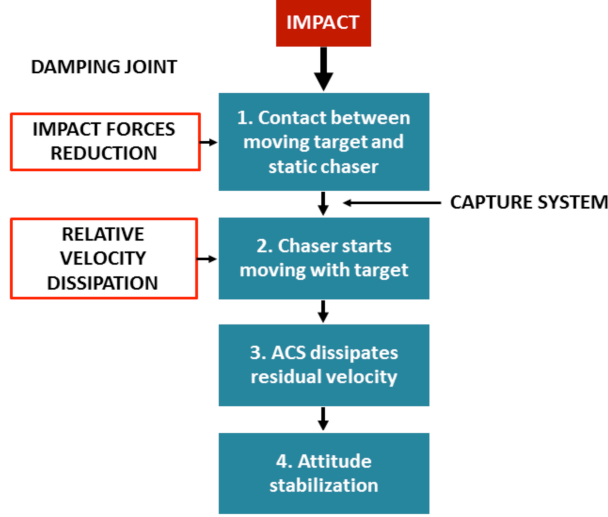
Figure 8: Capture procedure main phases.

damping joint plays a key role reducing the impulsive loads in the system. The low rigidity of this component reduces the peak forces in the transient dynamics, thus reducing the requisites of the adhesion system increasing the chances of successful docking. In the next phase (2) the two objects move together with a residual relative velocity. The damping joint dissipates the relative kinetic energy and the oscillations decay over time. After the relative motion is completely damped (3), the service spacecraft ADCS actuators (e.g. torque rods) can de-tumble the two body system. When the angular momentum is completely dumped, the system attitude is stabilized (4). Next it is possible to connect the HPM to the debris by means of the specifically designed docking mechanism. An interesting point of connection is the nozzle of Cosmos-3M, due to its high resistance to thermal, fluid dynamic and mechanical stresses. The so called *corkscrew system*, described in [44], could be a possible solution.

The main advantage of the proposed capture approach is that the adhesion mechanism does not require any particular structural feature to perform the grasping. In addition, since also large relative velocities between the target and the adhesive surface are tolerable, there is no need to identify a contact point that is slowly moving with respect to the chaser vehicle. These points have the following consequences: (a) there is no need to align the spacecraft with the axis of rotation of the debris, (b) the robotic arm limited work envelope is not an issue since it does not need to reach a particular point on the target. Furthermore, the possibility to dock virtually everywhere on the debris surface allows to place the end-effector closer to the center of mass of the target, in a more convenient position for handling. Finally, the larger tolerance on the docking position and relative velocities reduce the requisites on the spacecraft navigation and orbital control, as well as on the robotic arm trajectory, potentially reducing the resources dedicated to close proximity approach phases.

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4. HPM Sizing and De-orbiting Mass Budgets

The impact of propulsion maneuver on mass budget, system volume, and cost depends on many aspects such as the size of the target, the propulsion technology, and the type of reentry. Capability of throttling and reignition may represent a stringent requirement for the adequate control of the final disposing maneuver whereas compact design is important for easier docking to the target and for dynamic stability of the final assembly (de-orbiting module and target). Compact volume may request a higher average propellant density but may collide with ΔV requirements for a controlled reentry, needed by large systems. Thrust level should stem from a tradeoff choice between the risk of debris fragmentation and mission duration (correlated to propellant storability and collision risk during maneuver). Several innovative proposals are under development nowadays with varying time frame of realization, however, most of them need in-orbit demonstration of reliability and applicability on a real mission. Out of this group, it is worth mentioning the use of tethers, as single spaceships as well as in fleet, to perform uncontrolled de-orbiting even on multiple subjects [56, 57]. Other options, for the time being, appeal to systems already studied or realized in onboard de-orbiting devices, such as drag augmentation techniques (deployed sails or inflating balloons) or proven propulsion devices [58]. In this respect, a cost analysis for the de-orbiting of a 1.2 metric ton IRS-1C satellite was presented for different propulsion options, suggesting that chemical rockets can be a viable solution [59]. Within this pool of technologies, solid propellants represent a simple, reliable, and proven technology but feature low specific impulse and limited flexibility, while liquid propellants fill the gaps left by the solid propellants but larger volumes and higher degree of complexity are requested. Moreover, storability of the propellant must be carefully considered. Thus, hybrid rocket technology for de-orbiting applications is considered a valuable option due to the high specific impulse obtainable, intrinsic safety and, especially, thrust throttleability, possibility of green propellant use and low cost technology. Throttleability is important for rendezvous maneuvers with space targets. A hybrid rocket engine typically features the oxidizer in the liquid or gaseous state, while the fuel is in the solid state. Its safety is guaranteed by no-contact between fuel and oxidizer, except during the combustion phase. Hybrid rocket engines can also be built with a particular geometry, using a tangentially oxidizer injection, resulting very compact and highly efficient in combustion, thanks to the oxidizer flow that provides a vortex combustion. This particular kind of hybrid rocket engine results very small in size. Such characteristics can be the right solution for space debris mitigation, by supplementing with this engine the new satellites that will reach space in the future. However, in our view, this technology is very promising even in the field of space debris remediation, making possible the active removal in LEO of large intact objects (several metric tons), by placing on their surface one HPM for the reentry maneuver and, for attitude control, few small thrusters spilling the HPM liquid oxidizer and burning it as a monopropellant (dual-mode use) [40, 41].

Overall, a hybrid propulsion module represents a solution that mediates benefits and drawbacks from both liquid and solid rocket technology. On one side, it is bestowed throttleability and reignition capability typical of liquids, specific impulse levels which fall in between the performance of solid and liquid propulsion, and a higher mean propellant density due to the use of a solid fuel. However, a technological gap exists due to late development and lack of in-orbit demonstration. In the simplest possible configura-
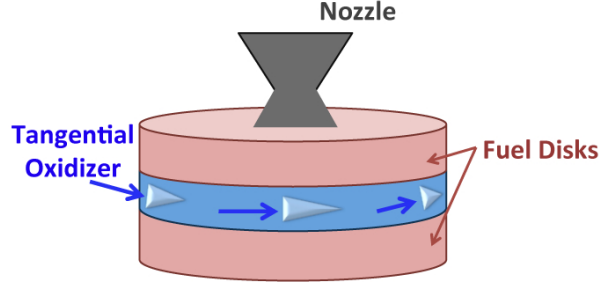
Figure 9: Vortex Flow Pancake engine scheme.

tion, a hybrid rocket is made by a center-perforated solid fuel placed in the combustion chamber where an injector blows in a liquid or gaseous oxidizer. Low regression rate is the main drawback of this combustion process but different means are considered for the enhancement of mass burning rate spanning from the use of advanced additives to different injection approaches (vortex combustion and planar vortex pancake) [60, 61, 62]. These advanced designs of the combustion chamber, see Figure 9, provides high combustion efficiency, low performance variation during combustion, and - in the case of solid metal additives - reduced emission of condensed combustion products (CCP) thanks to the vortex effect.
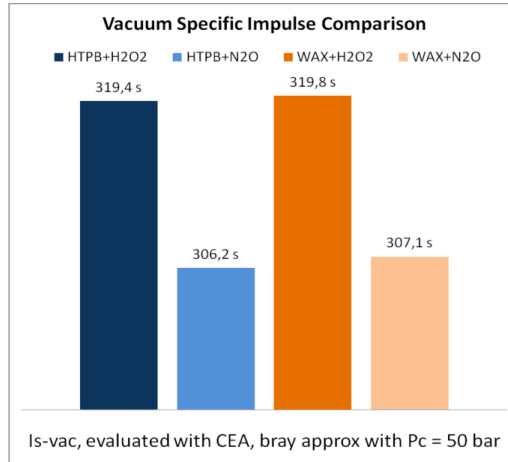

Figure 10: Ideal vacuum specific impulse comparison between HTPB and WAX burning in H_2O_2 or N_2O.

For the development of the HPM, the attention is focused on HTPB (hydroxyl-terminated polybutadiene) as fuel and N_2O or H_2O_2 as oxidizers. These combinations of propellants provides vacuum specific impulses (I_{s-vac}) over 300 s and significant volumetric specific impulses (I_v), due to the high density of the oxidizers, especially for hydrogen peroxide (Figures 10 and 11). In view of their good compromise between performance, costs and toxicity, hydrogen peroxide and nitrous oxide seems to be the best choice for

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this kind of application. From the preliminary performance analysis, conducted by the CEA software comparing HTPB and WAX solid fuels, it turns out that the ideal I_{s-vac} values for HTPB and WAX are similar, with a slight advantage for the latter one while, in terms of I_v, HTPB performs better due to its higher density. However, considering a multi-burn mission and special solid fuel geometries, WAX is not a suitable fuel, due to its low mechanical properties.
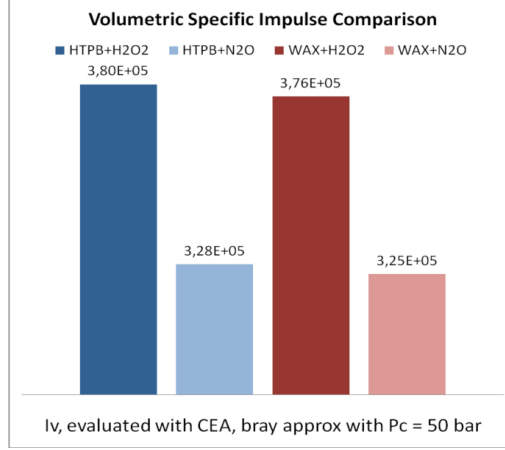

Figure 11: Ideal volumetric specific impulse comparison between HTPB and WAX burning in H_2O_2 or N_2O.

Turning to the choice of the oxidizer, catalytic decomposition hydrogen peroxide provides oxygen-rich hot gases up to 1,000 K. Considering that ignition of HTPB solid fuel requires about 800 K, it is possible to develop a simple and reliable re-ignition system. Moreover, with a single tank of H_2O_2, it is possible to feed both the primary propulsion system and a set of Reaction Control System (RCS) catalytic micro-thrusters. Though H_2O_2 is notorious for its storability issues, due to its decomposition inside tanks, high level of peroxide purity and the use of appropriate materials have demonstrated that risks can be avoided and the rate of dissociation can be reduced appreciably [63].

Controlled de-orbiting of a large object like Cosmos-3M second stage requires a ΔV of the order of 500 m/s. A tentative mission considers an initial thrust impulse for transfer to a lower orbit (Hohmann maneuver), then a further burn to lower the perigee below 100 km altitude and a final firing for the controlled atmospheric re-entry. The preliminary HPM size allows to plot the HPM mass required for the active removal of abandoned objects with masses up to 8,000 kg, Figure 12, in an altitude range between 700 and 1,000 km. The ratio between the propellant mass required and HPM mass is about 0.6.

Each altitude corresponds to the velocity increment required for de-orbiting, taking into account a safety margin of 10% increase over the maneuver value. For example, to de-orbit a Cosmos-3M second stage, with a mass of 1,400 kg, from an altitude of 1,000 km, a velocity increment of 552 m/s is required; this mission can be performed by an HPM with a mass of 515 kg, fitted into VEGA's payload case. Instead ENVISAT represents an even larger target. The ΔV requirement is lower due to orbital considerations but propellant mass budget and engine size increase, obtaining a HPM with a full mass of
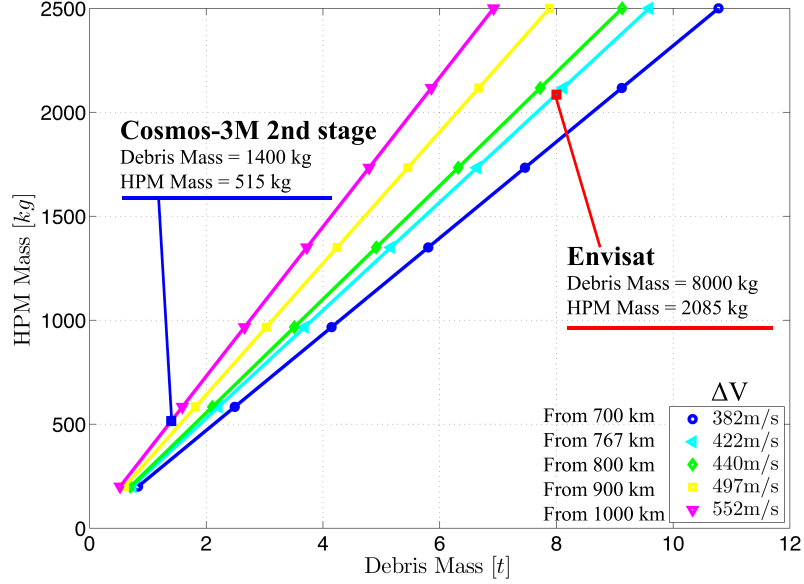
Figure 12: Debris mass vs. HPM mass, for several velocity requirements, corresponding to different debris altitudes.

2,085 kg. Launch can be provided only by Ariane 5ES or Soyuz Space Launcher. The great number of Cosmos stages still on orbit, in a range between 700 and 1,000 km, means a higher probability of catastrophic impacts with other spacecrafts or satellites. A space cleaning program could reasonably start with the removing of these objects.

For the HPM preliminary sizing, in order to de-orbit a Cosmos stage at 1,000 km altitude, the combination HTPB+H_2O_2 at an operating chamber pressure of 60 bar, with an ideal I_s of 320 s, is considered. To provide a velocity increment of 552 m/s a HPM is necessary with a mass of 513 kg, the classical solid fuel configuration (cylindrical fuel with axial oxidizer injection) is considered, or a mass of 534 kg in the case of vortex flow pancake configuration (two flat fuel disks with tangential injection). At this preliminary stage the two engine masses are comparable, since the Tsiolkovsky equation requires the same propellant mass, for a fixed value of ΔV and I_s. Proper experiments about combustion configurations and engine firing tests will provide the effective performance parameters for axial and vortex combustion. This should increase the gap between final HPM mass, probably in favor of the vortex one. However, from a preliminary sizing, the classical HPM has a diameter of 25 cm and a total length (including nozzle) of 176 cm. Instead the vortex HPM results more compact, with a diameter of 53 cm and a total length of 91 cm. If the oxidizer tanks are placed at the sides of the rocket, the final HPM diameter will be 144 cm for classical configuration and 170 cm for the vortex one (see Figure 13) [44]. Spherical tanks, with an internal membrane for pressurization with gaseous N_2, are considered. This preliminary HPM sizing takes into account also the mass of a RCS system for attitude control of the debris, during de-orbiting. Ignition is performed by a catalytic cell in which the hydrogen peroxide decomposes generating hot gases, then expanded in the combustion chamber.
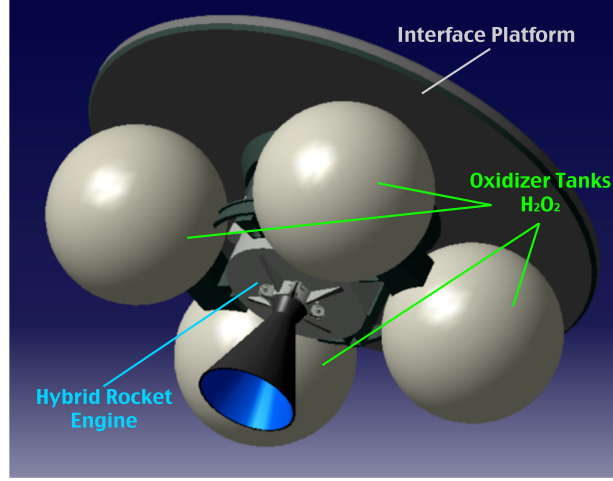
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Figure 13: Hybrid propulsion Module (HPM) conceptual sketch for a Vortex Flow Pancake configuration.

4.1. System Mass Budget

In order to perform a preliminary service platform mass estimate, a single debris removal mission is considered first. As already stated, very attractive targets for active removal are the Russian Cosmos-3M second stages, mainly concentrated around two inclinations, 74° and 83°, and altitudes between 650 and 1,050 km. For a preliminary estimate of the propellant mass needed for rendezvous with the target, a service platform launched with VEGA is assumed. Nevertheless, different launchers could be included in the analysis as well. Considering Cosmos-3M, TLE data and VEGA required nominal performance [64], it is assumed that the service platform is injected directly into the target nominal orbit plane in a 700 km parking orbit. Then, the Mass Ratio, MR_1, needed to transfer the servicing platform to the Cosmos-3M altitude (around 1,000 km) is estimated to be about 1.1, assuming again a two body model, a Hohmann transfer orbit and a reasonable hybrid engine performance ($I_s = 300$ s). It shall be outlined that the estimated mass ratio includes a 30% margin in the needed ΔV to consider also attitude control and close proximity maneuvers [65]. The total mass of the service platform maneuvering toward the target includes the bus mass, m_{bus}, the propellant mass for maneuvering, m_{prop}, and the payload mass, m_{HPM}, which consists of one single HPM mass. Since a single debris removal is being considered, after rendezvous and target de-orbiting, the service platform has to be de-orbited in turn. To this end, an additional Mass Ratio, MR_2, of about 1.2, is estimated, considering a 10% margin in the needed ΔV. Once the mass ratios relevant to the two considered maneuvers are known, the service platform mass can be estimated by using Eq. 2, that relates the space system initial mass to the service platform bus mass and the HPM mass.

$$MR_i = e^{\frac{\Delta V_i}{I_s g_0}} \tag{1}$$

$$m_{in} = MR_1(MR_2 m_{bus} + m_{HPM}) \tag{2}$$

Since the ADR vehicle has to carry in orbit all the avionics needed for mission operation and debris target rendezvous and capture, from preliminary system consideration and historical data relevant to rendezvous missions [46, 47], the bus mass can be estimated in the range 400-500 kg. Thus, using an average value of 450 kg, the initial system mass for one single debris removal is of the order of 1,290 kg, with a dry mass of about 1,085 kg and a propellant mass of about 205 kg. The ratio between dry and wet masses is about 0.84 as in [46, 47]. It shall be outlined that a margin of 30% has been added to the dry mass as in [65].

In order to consider a multi removal mission, Eq. 2 can be generalized. Specifically, if the removal of two debris targets is required, one finds

$$m_{in} = (MR_1 MR_2 MR_3)m_{bus} + MR_1(1 + MR_2)m_{HPM} \qquad (3)$$

now MR_2 is the Mass Ratio needed for the second target rendezvous (a target at 850 km altitude is considered), and MR_3 is the Mass Ratio needed for service platform controlled de-orbiting. If we assume that no orbit plane changes are required for the second rendezvous, a MR_2 of about 1.03 can be estimated (30% margin included in the needed ΔV). In this case, the system wet mass is about 1,863 kg, with a dry mass of about 1,585 kg (30% margin included) and a propellant total mass of about 278 kg. Table 2 summarizes the results relevant to the preliminary ADR vehicle mass budget. The value of the ratio between the total removed mass and the ADR vehicle wet mass suggests that for targets with mass lower than 1,000 kg 3 to 5 removals per year might be feasible for targets with mass lower than 1,000 kg.

	1# Removal	2# Removal
Total Removed Mass	1,400	2,800
ADR Wet Mass	$\sim 1,300$	$\sim 1,900$
ADR Dry Mass/ADR Wet Mass	~ 0.84	~ 0.85
Total Removed Mass/Total HPM Mass	~ 2.8	~ 2.8
Total Removed Mass/ADR Wet Mass	~ 1.1	~ 1.5

Table 2: ADR vehicle budget results.

With regard to a multiple removal mission, it shall be outlined that in order to minimize the total required ΔV, possible methods for prioritizing and categorizing the debris of interest, producing subsets of the overall population, have to be implemented (for example they could be grouped considering altitude, orbit inclination and RAAN). Within a subset, the order in which the debris targets are visited has to be selected to reduce propellant budgets. Moreover, to limit orbit plane changes, the beneficial effects of the non-spherical (J2) perturbation could be exploited, waiting for orbit plane alignments [66].

5. Conclusions

In this paper the active removal of a large debris target from LEO by exploiting hybrid propulsion technology was investigated. A demonstration mission concept was developed in which, following debris target identification, rendezvous and capture, a

de-orbiting kit consisting of a hybrid propulsion module is attached to the debris for a controlled re-entry. For a preliminary analysis of the various technologies involved in the removal mission Russian Cosmos-3M second stages were considered as candidate targets for removal with a VEGA-class launcher. Critical issues relevant to the debris active removal system were investigated, with particular focus on debris de-orbiting strategy, possible approaches and technologies for debris rendezvous, capture and mating, and hybrid propulsion module sizing and de-orbiting performance estimate. For the capture system design an innovative concept based on electro-adhesive capture was introduced to cope with a non cooperating target. This solution, indeed, shall allow high flexibility in capture point selection and high adaptability to the debris surface. For hybrid propulsion module sizing, different design solutions were compared and performance parameters evaluated as a function of the mass to be removed. Preliminary results show that the ratio between the removed mass and the engine total mass ranges from 3 to 4. In addition, for two removals the ratio between the total removed mass and the removal system wet mass is about 1.5. These results demonstrate that the removal of up two Cosmos-3M second stages might be feasible within a single mission with a VEGA-class launcher. Instead, for the removal of very large objects, like the ENVISAT satellite, or performing a mission with multi-removal capability of 3 to 5 objects per year, launches with Ariane or Soyuz shall be considered.

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