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Conformal gravity with totally antisymmetric torsion

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We present a gauge theory of the conformal group in four spacetime dimensions with a nonvanishing torsion. In particular, we allow for a completely antisymmetric torsion, equivalent by Hodge duality to an axial vector whose presence does not spoil the conformal invariance of the theory, in contrast with claims of antecedent literature. The requirement of conformal invariance implies a differential condition (in particular, a Killing equation) on the aforementioned axial vector, which leads to a Maxwell-like equation in a four-dimensional curved background. We also give some preliminary results in the context of $\mathcal{N} = 1$ four-dimensional conformal supergravity in the geometric approach, showing that if we only allow for the constraint of vanishing supertorsion, all the other constraints imposed in the spacetime approach are a consequence of the closure of the Bianchi identities in superspace. This paves the way towards a future complete investigation of the conformal supergravity using the Bianchi identities in the presence of a nonvanishing (super)torsion.

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I. INTRODUCTION

In Ref. [1], it was shown that the locally scale-invariant Weyl theory of gravity is the gauge theory of the conformal group, where conformal transformations (conformal boosts) are gauged by a nonpropagating gauge field. In that theory, the authors adopted the formalism of Refs. [2,3] to construct a quadratic Lagrangian with the curvatures associated with the conformal group in four spacetime dimensions. They claimed that in order to produce a conformally invariant theory in this setup, it is necessary to set the torsion to zero.¹

In this work, in contrast with this claim, we show that it is actually possible to construct a gauge theory of the conformal group in four spacetime dimensions with a nonvanishing torsion component where proper conformal transformations are gauged by a nonpropagating gauge field (the Schouten 1-form field). In particular, we allow for

a totally antisymmetric torsion, equivalent by Hodge duality to an axial vector, and still get a conformal gauge theory whose Lagrangian is quadratic in the curvatures of the conformal algebra (as with that of Ref. [1], which is the same construction as in Refs. [2,3]). We explicitly show how to reproduce the Weyl Lagrangian in this framework in the presence of a nonvanishing, completely antisymmetric torsion and study the field equations of the theory. Let us also mention that for quadratic theories, in general, when working in the first-order or in the second-order formalism for the spin connection, one obtains different results. We will adopt the second-order formalism, which will allow us to end up with a fourth-order propagation equation for the graviton, the Lorentz connection now being torsionful.

In this setup, invariance under conformal boosts (also known as proper, or special, conformal transformations) implies a *Killing vector equation*—namely, a differential condition on the axial vector torsion which, upon further differentiation, leads to a Maxwell-like equation in a four-dimensional curved background. In the limit in which the torsion is set to zero, we recover the conformal theory of Ref. [1].

The first part of this work will be devoted to studying the purely bosonic gravitational theory. Subsequently, in view of a complete future investigation of the supersymmetric extension of this theory [4], we give some preliminary results we have obtained regarding conformal supergravity. We will adopt the geometric approach to supergravity (also called the *supergroup manifold* approach or *rheonomic* approach).²

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¹Actually, in Ref. [1], the explicit form of the spin connection contains the dilaton, which gives a torsion trace contribution. However, the latter can be consistently set to zero in the theory, as already observed in Ref. [1] and as we will also discuss in the present work. Therefore, let us refer to the theory of Ref. [1] as a torsion-free one.

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²For details on this formalism, see the original formulation in Ref. [5] and the pedagogical review [6].

As pioneering works on the structure of conformal supergravity at the linearized level, we refer the reader to Refs. [7,8]. The full conformal $\mathcal{N} = 1$ supergravity theory in $D = 3 + 1$ spacetime dimensions has been presented in Refs. [9,10] (see also the review work [11]) and subsequently rephrased in Ref. [12] in the geometric approach to supergravity. Interesting recent developments on \mathcal{N} -extended conformal supergravity and its spectrum in four dimensions have been recently obtained in Ref. [13]. In all these papers, together with the vanishing of (super) torsion, a set of additional constraints were also imposed. In particular, in Ref. [12], the constraints were implemented by the use of Lagrange multipliers. The constraint of vanishing supertorsion was justified by arguing that only in this case would the Lagrangian have been invariant under special conformal transformations. Since we will prove in the sequel of this work that, at least at the purely bosonic level, one can still recover invariance under (special) conformal transformations allowing for a nonvanishing axial vector torsion, we argue that something similar should presumably happen in the superconformal case. In view of future investigations in this direction, here we start a preliminary analysis at the level of Bianchi identities in the geometric approach, showing that, besides the vanishing supertorsion, all the aforementioned constraints can be directly obtained from the study of the Bianchi identities,³ just imposing the vanishing of the supertorsion (that is, the supersymmetric extension of the constraint imposed in Ref. [1], where torsion was indeed assumed to vanish).

The paper is organized as follows: In Sec. II, we give a brief review of the conformal setup in four spacetime dimensions. In Sec. III, we develop the gauge theory of the conformal group with a nonvanishing, completely antisymmetric torsion. Subsequently, in Sec. IV, we give some preliminary results regarding the extension to $\mathcal{N} = 1$, $D = 4$ conformal supergravity with vanishing supertorsion in the geometric approach. We conclude our work with some remarks and a discussion on future developments. In the Appendix, some useful formulas on gamma matrices in four dimensions are collected.

II. REVIEW OF THE GAUGING OF THE CONFORMAL GROUP

The conformal group [15] $O(4, 2)$ is locally isomorphic to $SU(2, 2)$. The corresponding algebra is generated by the set of generators $\mathbf{T}_A = \{\mathbf{J}_{ab}, \mathbf{P}_a, \mathbf{K}_a, \mathbf{D}\}$, where we have decomposed the adjoint index A of the conformal algebra with respect to the Lorentz indices $a, b, \dots = 0, 1, 2, 3$. \mathbf{J}_{ab} are the Lorentz rotations, \mathbf{P}_a the spacetime translations, \mathbf{K}_a the conformal boosts, and \mathbf{D} the dilatation (scale

³See also Ref. [14], where the authors used the geometric approach as in the present case, albeit in a different context, and the same conclusions can be reached after appropriate truncations.

transformation). In our conventions, the metric η_{ab} has the signature $(+, -, -, -)$.⁴

Let us introduce the gauge 1-form fields ω^{ab} (spin connection⁵), V^a (vierbein), K^a (special conformal 1-form field), and \mathfrak{D} (dilaton gauge field), which are respectively dual to the vector field generators of the conformal algebra, namely

$$\begin{aligned} \omega^{ab}(\mathbf{J}_{cd}) &= 2\delta_{cd}^{ab}, & V^a(\mathbf{P}_b) &= \delta_b^a, & K^a(\mathbf{K}_b) &= \delta_b^a, \\ \mathfrak{D}(\mathbf{D}) &= 1. \end{aligned} \quad (2.1)$$

We can then write the corresponding curvatures:

$$\begin{aligned} R^{ab} &\equiv \mathcal{R}^{ab} - 4V^{[a} \wedge K^{b]}, \\ T^a &\equiv \mathcal{D}V^a + \mathfrak{D} \wedge V^a, \\ \mathcal{T}^a &\equiv \mathcal{D}K^a - \mathfrak{D} \wedge K^a, \\ G &\equiv d\mathfrak{D} + 2V^a \wedge K_a, \end{aligned} \quad (2.2)$$

where $\mathcal{D} = d - \omega$ is the Lorentz covariant derivative, and

$$\mathcal{R}^{ab} = d\omega^{ab} - \omega^a_c \wedge \omega^{cb} \quad (2.3)$$

is the Riemann curvature.⁶ Setting the curvatures (2.2) to zero, the vanishing right-hand sides define the Maurer-Cartan equations, describing the ‘‘vacuum’’ (ground state), dual to the commutator algebra of the vector field generators $\{\mathbf{J}_{ab}, \mathbf{P}_a, \mathbf{K}_a, \mathbf{D}\}$ (as is well known, the d^2 closure of the Maurer-Cartan equations coincides with the Jacobi identities of the algebra). For the sake of convenience, let us also define

$$\begin{aligned} \hat{\mathcal{D}}V^a &\equiv \mathcal{D}V^a + \mathfrak{D} \wedge V^a = dV^a - \omega^a_b V^b + \mathfrak{D} \wedge V^a, \\ \hat{\mathcal{D}}K^a &\equiv \mathcal{D}K^a - \mathfrak{D} \wedge K^a = dK^a - \omega^a_b K^b - \mathfrak{D} \wedge K^a, \end{aligned} \quad (2.4)$$

where $\hat{\mathcal{D}}$ denotes the *Lorentz and scale* covariant differential. The length-scale weights of the 1-forms and of their corresponding curvatures are

⁴Regarding our conventions, throughout the paper we will use rigid latin indices $a, b, \dots = 0, 1, 2, 3$ instead of the world greek indices $\mu, \nu, \dots = 0, 1, 2, 3$, and will expand the p -forms in terms of the vierbein basis rather than in terms of differentials. For example, a generic 2-form will be expanded as $F^A = F^A_{bc} V^b V^c = F^A_{\mu\nu} dx^\mu \wedge dx^\nu$, where $V^a = V^a_\mu dx^\mu$. This choice is convenient for the extension of the theory to superspace using the geometric formalism where the p -forms are expanded in terms of the full supervierbein basis (V^a, ψ^α) , ψ^α being the gravitino 1-form. We will come back to a preliminary study of conformal supergravity in Sec. IV.

⁵We call ω^{ab} the spin connection antisymmetric in a, b , $\omega^{ab} = -\omega^{ba}$, which may (and in fact will) involve torsion.

⁶We will generally omit writing the wedge product between differential forms in order to lighten the notation.

$$[\omega^{ab}] = [\mathfrak{D}] = 0, \quad [V^a] = 1, \quad [K^a] = -1. \quad (2.5)$$

Exploiting Eq. (2.4), the curvatures (2.2) can be recast into the following simpler expressions:

$$\begin{aligned} R^{ab} &\equiv \mathcal{R}^{ab} - 4V^{[a}K^{b]}, \\ T^a &\equiv \hat{D}V^a, \\ \mathcal{T}^a &\equiv \hat{D}K^a, \\ G &\equiv d\mathfrak{D} + 2V^aK_a, \end{aligned} \quad (2.6)$$

and the Bianchi identities obeyed by the curvatures (2.6) are

$$\begin{aligned} DR^{ab} + 4(T^{[a}K^{b]} - V^{[a}T^{b]}) &= 0, \\ \hat{D}T^a + R^{ab}V_b - GV^a &= 0, \\ \hat{D}\mathcal{T}^a + R^{ab}K_b + GK^a &= 0, \\ dG - 2T^aK_a + 2V^aT_a &= 0, \end{aligned} \quad (2.7)$$

where

$$\begin{aligned} \hat{D}T^a &\equiv DT^a + \mathfrak{D} \wedge T^a, \\ \hat{D}\mathcal{T}^a &\equiv D\mathcal{T}^a - \mathfrak{D} \wedge \mathcal{T}^a. \end{aligned} \quad (2.8)$$

The conformal gauge transformations, associated with the conformal algebra, read

$$\begin{aligned} \delta\omega^{ab} &= \mathcal{D}\varepsilon^{ab} + 4\varepsilon^{[a}K^{b]} - 4V^{[a}\varepsilon_K^{b]}, \\ \delta V^a &= \hat{D}\varepsilon^a + \varepsilon^{ab}V_b - \varepsilon_{\mathfrak{D}}V^a, \\ \delta K^a &= \hat{D}\varepsilon_K^a + \varepsilon^{ab}K_b - \varepsilon_{\mathfrak{D}}K^a, \\ \delta\mathfrak{D} &= d\varepsilon_{\mathfrak{D}} - 2\varepsilon^aK_a + 2V_a\varepsilon_K^a, \end{aligned} \quad (2.9)$$

where ε^{ab} , ε^a , ε_K^a , and $\varepsilon_{\mathfrak{D}}$ are the Lorentz, translation, conformal boost, and dilatation parameters, respectively. Restricting ourselves to conformal boosts and dilatations, we have

$$\begin{aligned} \delta\omega^{ab} &= -4V^{[a}\varepsilon_K^{b]}, \\ \delta V^a &= -\varepsilon_{\mathfrak{D}}V^a, \\ \delta K^a &= \hat{D}\varepsilon_K^a - \varepsilon_{\mathfrak{D}}K^a, \\ \delta\mathfrak{D} &= d\varepsilon_{\mathfrak{D}} + 2V_a\varepsilon_K^a. \end{aligned} \quad (2.10)$$

Let us mention that the theory whose Lagrangian we are going to consider [see Eq. (3.2) in the following] is invariant under diffeomorphisms by construction, since it is written in terms of differential forms, but it is not invariant under spacetime translations. This is what commonly happens in gravitational theories. Thus, it is not a true ‘‘gauge’’ theory of the conformal group. However, we

shall adopt the terminology of ‘‘gauge theory of the conformal group’’ since it is widely used in the literature, keeping in mind that, in fact, we just have diffeomorphism invariance rather than invariance under spacetime translations.⁷

Finally, let us also recall that the curvatures (2.6) can be expanded along the vierbeins, which are dual to the spacetime translation generators. This amounts to the requirement of having the conformal symmetry of a theory defined on spacetime, and it is therefore a natural physical request to have a conformal gravity theory within our approach. Indeed, as Lorentz and scale symmetries are an exact invariance of the Lagrangian (which we will introduce in the following section), the coset

$$\frac{\text{SU}(2,2)}{\text{SO}(1,3) \otimes \text{O}(1,1)}$$

only depends on the vierbein and the K^a 1-form. On the other hand, since we shall see that the gauge field K^a can be expressed in terms of contractions of the Riemann tensor (more precisely, the Schouten tensor, as we will discuss in the sequel), the cotangent space is spanned in terms of the vierbein only.

Therefore, the aforesaid expansion of the curvatures along the vierbein basis reads

$$\begin{aligned} R^{ab} &= R^{ab}_{cd}V^cV^d, \\ T^a &= T^a_{bc}V^bV^c, \\ \mathcal{T}^a &= \mathcal{T}^a_{bc}V^bV^c, \\ G &= G_{ab}V^aV^b. \end{aligned} \quad (2.11)$$

We will now proceed with the development of a gauge theory of the conformal group with a nonvanishing torsion.

III. GAUGE THEORY OF THE CONFORMAL GROUP IN THE PRESENCE OF A NONVANISHING TORSION

We consider the same action introduced in Ref. [1], which is the only parity-conserving quadratic action that

⁷We recall that if we let the index A denote the coadjoint representation, an infinitesimal diffeomorphism of anholonomic parameter $\varepsilon^A = \varepsilon^\rho \mu_\rho^A$ on any gauge field of the algebra μ^A can be written as $\delta_\varepsilon \mu^A = D\varepsilon^A + \iota_\varepsilon R^A$, where D is the covariant derivative in the coadjoint representation. Therefore, the diffeomorphisms of the gauge fields differ from the gauge translations by a term proportional to the contraction of the curvature along an infinitesimal translation $\varepsilon^a \mathbf{P}_a$, where ε^a is the infinitesimal parameter. In the supersymmetric case, the contraction of the supercurvatures along a supersymmetric generator is also in general different from zero. Therefore, in the superconformal case, the superspace translations correspond to supersymmetry transformations.

can be constructed from the curvatures (2.6) without dimensional constants:

$$A = \int_{\mathcal{M}_4} \mathcal{L}, \quad (3.1)$$

where

$$\mathcal{L} = R^{ab} \wedge R^{cd} \epsilon_{abcd} \quad (3.2)$$

is the Lagrangian 4-form and \mathcal{M}_4 is the four-dimensional spacetime.

Let us first recall the well-known fact that the variation of the action with respect to the special conformal 1-form K^b gives an algebraic equation for the special conformal gauge field K^a . Indeed, varying the Lagrangian (3.2) with respect to K^b , we obtain the field equations

$$-8V^a R^{cd} \epsilon_{abcd} = 0, \quad (3.3)$$

which imply, using the expansion along the vierbein basis [Eq. (2.11)],

$$\begin{aligned} R &= 0, \\ \check{R}_{ab} &= 0, \end{aligned} \quad (3.4)$$

where $R = R^{ab}{}_{ab}$ and $\check{R}_{ab} = R^c{}_{acb}$. Taking Eq. (3.4) together with the definition of R^{ab} in Eq. (2.6) and writing $K_a = K_{ab}V^b$, one gets

$$K_{ab} = S_{ab}, \quad (3.5)$$

with S_{ab} being the Schouten 0-form tensor defined in four-dimensional spacetime as

$$S_{ab} \equiv \frac{1}{2} \left(\check{R}_{ab} - \frac{1}{6} \eta_{ab} \mathcal{R} \right), \quad (3.6)$$

where $\mathcal{R} = \mathcal{R}^{ab}{}_{ab}$ and \check{R}_{ab} are the scalar curvature and the Ricci tensor of ω^{ab} , respectively. Notice that, in the presence of a nonvanishing torsion, ω^{ab} also includes a contorsion component, implying that the Schouten tensor S_{ab} has a nonvanishing antisymmetric part, $S_{[ab]} = \check{R}_{[ab]}$, given entirely in terms of torsion.

Thus, we have obtained an algebraic equation for the nonpropagating gauge field K_{ab} —namely, Eq. (3.5)—which tells us that K_{ab} corresponds to the Schouten tensor. This is a well-known fact (see, for instance, Ref. [1]). However, some comment is in order on this point, and we make it in the following sections. In particular, we will show that the fact that K_{ab} corresponds to the Schouten can be actually deduced directly from a vacuum analysis, and we will also give the irreducible decomposition of \mathcal{R}_{abcd} , which will be useful in the sequel.

A. Curvature irreducible decomposition and Schouten tensor

Let us briefly discuss, before proceeding with our main results, that already at the vacuum level, as in fact expected, one can show that the field K^a cannot be anything other than the Schouten 1-form, $S^a = S^{ab}V_b$. This can be shown by taking into account the irreducible decomposition of the Riemann tensor $\mathcal{R}^{ab}{}_{cd}$ (here allowing also for the presence of torsion; see, for instance, Ref. [16] for details). Regarding the number of components, in four spacetime dimensions we have $\dim(\mathcal{R}^{ab}{}_{cd}) = 6 \times 6 = 20 \oplus 15 \oplus 1$, corresponding in terms of the $SL(4)$ representations to the dimensions of the following Young diagrams:

$$\begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \end{array} \otimes \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \end{array} = \begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \square \\ \hline \end{array} \oplus \begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \\ \hline \end{array} \oplus \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array} \quad (3.7)$$

Decomposing the $SL(4)$ representations with respect to $SO(1,3)$ in terms of their traceless plus trace parts, we find six irreducible pieces (irrepses):

$$\begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \square \\ \hline \end{array} \longrightarrow \begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \square \\ \hline \end{array} \oplus \begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \square \\ \hline \end{array} \oplus \bullet \quad (3.8)$$

$$\begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \\ \hline \end{array} \longrightarrow \begin{array}{|c|c|} \hline \square & \square \\ \hline \square & \square \\ \hline \end{array} \oplus \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \end{array} \quad (3.9)$$

$$\begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array} \longrightarrow \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array} \quad (3.10)$$

where the small ring on top of the diagrams on the right-hand sides means that the corresponding representation is traceless, while the bullet denotes the scalar representation.

The three irrepses on the right-hand side of Eq. (3.8) correspond to the 10-dimensional Weyl tensor W_{abcd} , the 9-dimensional traceless symmetric Ricci tensor $\check{\mathcal{R}}_{(ab)}$, and the scalar curvature \mathcal{R} , respectively, and, using the nomenclature of Ref. [16], are called WEYL + RICSIMF + SCALAR. The two irrepses on the right-hand side of Eq. (3.9) correspond to the 9-dimensional tensor, which has the same number of degrees of freedom as a symmetric traceless tensor,⁸ plus the 6-dimensional

⁸The corresponding representation is commonly referred to as *associated*.

traceless antisymmetric Ricci tensor $\check{\mathcal{R}}_{[ab]}$, shortly referred together to as PAIRCOM and RICANTI, respectively. Finally, on the right-hand side of Eq. (3.10), we have a pseudoscalar Hodge dual to $\mathcal{R}_{[abcd]}$ denoted as PSSCALAR.

Writing

$$\begin{aligned} \mathcal{R}_{abcd} = & \underbrace{\frac{1}{2}(\mathcal{R}_{abcd} + \mathcal{R}_{cdab})}_{\text{WEYL+RICSIMF+SCALAR}} \\ & \oplus \underbrace{\frac{1}{2}[(\mathcal{R}_{abcd} - \mathcal{R}_{cdab}) - \mathcal{R}_{[abcd]}]}_{\text{PAIRCOM+RICANTI}} \oplus \underbrace{\mathcal{R}_{[abcd]}}_{\text{PSSCALAR}}, \end{aligned} \quad (3.11)$$

the three ‘‘underbraced’’ expressions correspond to the left-hand sides of the three Eqs. (3.8)–(3.10), respectively. We enumerate the various representations, writing

$$\begin{aligned} \mathcal{R}^{ab}{}_{cd} = & \sum_{i=1}^6 \mathcal{R}^{ab}{}_{(i)cd} \\ = & \underbrace{\mathcal{R}^{ab}{}_{(1)cd}}_{\text{WEYL [10]}} + \underbrace{\mathcal{R}^{ab}{}_{(2)cd}}_{\text{PAIRCOM [9]}} + \underbrace{\mathcal{R}^{ab}{}_{(3)cd}}_{\text{PSSCALAR [1]}} + \underbrace{\mathcal{R}^{ab}{}_{(4)cd}}_{\text{RICSIMF [9]}} \\ & + \underbrace{\mathcal{R}^{ab}{}_{(5)cd}}_{\text{RICANTI [6]}} + \underbrace{\mathcal{R}^{ab}{}_{(6)cd}}_{\text{SCALAR [1]}}. \end{aligned} \quad (3.12)$$

where the numbers in the square brackets denote the components of each irrep (dimension of the irrep).

Let us observe that $\mathcal{R}^{ab}{}_{(2)cd}$, $\mathcal{R}^{ab}{}_{(3)cd}$, and $\mathcal{R}^{ab}{}_{(5)cd}$ are nonvanishing only in the presence of torsion (so that they are given in terms of torsion and its derivatives).

Now, we can exploit Eq. (3.12) in the vacuum of our theory, given by the vanishing right-hand side of Eq. (2.6). In particular, we have

$$\mathcal{R}^{ab} - 4V^{[a}K^{b]} = 0; \quad (3.13)$$

that is,

$$\mathcal{R}^{ab}{}_{cd} - 4\delta^{[a}{}_{[c}K^{b]}{}_{d]} = 0. \quad (3.14)$$

Using Eq. (3.12) and observing that the second term in Eq. (3.14) can be written only in terms of the irreducible pieces RICSIMF, SCALAR, and RICANTI, one can prove that K_{ab} must coincide with the Schouten tensor S_{ab} (that is, $K^a{}_b = S^a{}_b V^b$), the components of the latter being defined in Eq. (3.6).⁹ Therefore, the Maurer-Cartan equations obtained from Eq. (2.6) take the following form:

⁹Indeed, by tracing the b, d indices of both terms on the left-hand side of Eq. (3.14), one easily recovers $K_{ab} = S_{ab}$.

$$\begin{aligned} \mathcal{R}^{ab}{}_{(2)} & \equiv \mathcal{R}^{ab}{}_{(2)cd}V^cV^d = 0, \\ \mathcal{R}^{ab}{}_{(3)} & \equiv \mathcal{R}^{ab}{}_{(3)cd}V^cV^d = 0, \\ W^{ab} & \equiv \mathcal{R}^{ab}{}_{(1)} = \mathcal{R}^{ab} - 4V^{[a}S^{b]} = 0, \\ T^a & \equiv \hat{D}V^a = 0, \\ \mathcal{T}^a & \equiv \hat{D}S^a = 0, \\ G & \equiv d\mathfrak{D} + 2V^aS_a = 0, \end{aligned} \quad (3.15)$$

where $\mathcal{R}^{ab}{}_{(2)}$ and $\mathcal{R}^{ab}{}_{(3)}$ are the PAIRCOM and PSSCALAR 2-forms, respectively, and $W^{ab} = W^{ab}{}_{cd}V^cV^d$, with $W^{ab}{}_{cd}$ being the Weyl tensor.

One could then go out of the vacuum switching on the curvatures associated with the Maurer-Cartan equations (3.15) and write a quadratic Lagrangian in terms of these field strengths. We note that in the case of vanishing torsion ($T^a = 0 \Rightarrow \mathcal{R}^{ab}{}_{(2)} = \mathcal{R}^{ab}{}_{(3)} = 0$), such a Lagrangian reads

$$\mathcal{L}_W = W^{ab} \wedge W^{cd}\epsilon_{abcd}, \quad (3.16)$$

where $W^{ab} = W^{ab}(\omega)$, with $\omega = \omega(\mathfrak{D}, V)$ (note that we still have $\mathcal{R}^{ab}{}_{(5)cd} \neq 0$, due to the presence of \mathfrak{D}). The Lagrangian (3.16) coincides with the Lagrangian given in Ref. [1]. In that paper, the authors set the torsion T^a equal to zero right from the beginning, and plugging back into Eq. (3.2) the on-shell expression [Eq. (3.5)] for K_{ab} , they recover the same Weyl Lagrangian [Eq. (3.16)]. As was to be expected, the theory in Eq. (3.16) that we have constructed by gauging directly the Maurer-Cartan equations (3.15) coincides, when T^a is zero, with the theory of Ref. [1]. In fact, Eq. (3.5) shall be interpreted directly as a consequence of the structure of the vacuum of the theory quadratic in the Weyl tensor, which is indeed the conformal theory we are going to focus on.

In the sequel, we will show that, remarkably, the Lagrangian we will develop describing a gauge theory of the conformal group with a *nonvanishing torsion* is formally identical to the Lagrangian in Eq. (3.16), provided the curvatures are constructed from a torsionful connection. In view of this, let us proceed by first showing that it is still possible to get conformal invariance of the theory in the presence of a nonvanishing T^a . In other words, we are going to prove that the constraint of vanishing torsion introduced in Ref. [1] to get a conformally invariant theory can actually be relaxed.

B. Conformal invariance of the theory

The aim of this section is to see whether a nonvanishing T^a is allowed in a ‘‘gauge’’ theory of the conformal group. In particular, we will allow for a totally antisymmetric torsion, equivalent by Hodge duality to an axial vector. We will show that the requirement of conformal invariance of

the Lagrangian (3.2) constructed with the curvatures in Eq. (2.6) can still be fulfilled provided we require the vanishing of $R_{(2)}^{ab}$, $R_{(3)}^{ab}$, and $R_{(5)}^{ab}$ (notations for the irreps of R^{ab} are the same as for \mathcal{R}^{ab}). This will imply a differential condition on the completely antisymmetric part of the torsion, namely for the aforesaid Hodge-dual axial vector. Upon use of the on-shell conditions in Eq. (3.4) implying Eq. (3.5), the theory will *formally* reproduce the same Lagrangian (3.16), albeit with a torsionful spin connection.¹⁰

Let us show the above explicitly. The Lagrangian (3.2) is clearly scale invariant, as R^{ab} has zero scale weight. Nevertheless, the invariance under conformal boosts can be achieved only in a nontrivial way. In particular, one can prove that some constraints on the curvatures in Eq. (2.6) have to be imposed in order for Eq. (3.2) to be invariant under proper conformal transformations on spacetime. Indeed, to recover the invariance of Eq. (3.2) under conformal boosts, performing the variation $\delta K^d = \hat{D}\epsilon_K^d$ [see Eq. (2.9)], we must have

$$(R^{ab}\epsilon_{abcd}T^c)\epsilon_K^d = 0; \quad (3.17)$$

that is, using Eq. (2.11),

$$(R^{ab}{}_{lm}T^c{}_{pq}\epsilon^{lm pq}\epsilon_{abcd}\Omega^{(4)})\epsilon_K^d = 0, \quad (3.18)$$

where $\Omega^{(4)}$ is the four-dimensional volume element defined as $\Omega^{(4)} \equiv -\frac{1}{4!}\epsilon_{abcd}V^aV^bV^cV^d$. In Ref. [1], the authors claimed that Eq. (3.2) yields invariant results under proper conformal gauge transformations only if $T^a = 0$ (the vanishing of $R^{ab}{}_{lm}$ not being considered, as it would trivialize the theory). Actually, this is not the case, as we will show in the sequel.

In order to explain our claim in detail, we need the irreducible decomposition of the torsion in four spacetime dimensions (see, for instance, Refs. [16–18]). In four dimensions, the torsion tensor $T^a{}_{bc}$ has $24 = 20 \oplus 4$ components, and we may write its decomposition as

$$\square \otimes \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \end{array} = \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array} \oplus \square \oplus \begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array} \quad (3.19)$$

whose dimensions are 16, 4, and 4, respectively, corresponding to the decomposition

¹⁰Note that the same Lagrangian can be also obtained by directly gauging the Maurer-Cartan equations (3.15)—that is, switching on the corresponding curvatures going out of the vacuum, in the presence of a completely antisymmetric torsion, and with the aforementioned constraints.

$$T^a{}_{bc} = T^a{}_{\begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array}} + \frac{2}{3}\delta_{[b}^at_{c]} + T_{[abc]}. \quad (3.20)$$

In the following, we will denote the 16-dimensional representation as a tensor $Z^a{}_{bc}$ —that is,

$$T^a{}_{\begin{array}{|c|} \hline \square \\ \hline \square \\ \hline \square \\ \hline \end{array}} = Z^a{}_{bc}, \quad (3.21)$$

and the antisymmetric representation $T_{[abc]}$ as the axial vector \tilde{t}^d , namely

$$T_{[abc]} = -\frac{1}{6}\epsilon_{abcd}\tilde{t}^d, \quad (3.22)$$

while t^a appearing in the torsion trace part is an ordinary vector. Inserting the above decomposition of the torsion into Eq. (3.18), the latter becomes

$$\left[R^{ab}{}_{lm} \left(\frac{2}{3}\delta_p^at_q - \frac{1}{6}\epsilon^{cpqr}\tilde{t}_r + Z^c{}_{pq} \right) \epsilon^{lm pq}\epsilon_{abcd} \right] \epsilon_K^d = 0. \quad (3.23)$$

The necessary condition given by Eq. (3.23) consists of a set of four algebraic equations (recall that ϵ_K^d is arbitrary) in the curvatures R^{ab} and $T^a{}_{bc}$, for which we are now going to examine in detail some particular solutions.

We first observe that a sufficient constraint on the torsion to have conformal boost invariance is $\tilde{t}_a = Z^a{}_{bc} = 0$. Indeed, in this case, Eq. (3.23) yields

$$(-Rt_a + 2\check{R}_{ba}t^b)\epsilon_K^a = 0 \rightarrow Rt_a - 2\check{R}_{ba}t^b = 0, \quad (3.24)$$

which, for a nonvanishing torsion trace t_a , has as a particular solution

$$R = 0, \quad \check{R}_{ab} = 0. \quad (3.25)$$

The latter constraints coincide with the equations that one obtains when varying the Lagrangian (3.2) with respect to K^a (that is, when going on shell for K^a), namely with Eq. (3.4). However, let us recall here that the torsion trace t_a , even if perfectly allowed, as we have just seen that it does not spoil the conformal invariance of the theory, can actually be set to zero in a consistent way (see, e.g., Ref. [19] for details). Indeed, one can easily verify that the torsion is invariant under a shift of the dilaton $\mathfrak{D} = \mathfrak{D}_a V^a$ by a parameter X_a , namely $\mathfrak{D}_a \rightarrow \mathfrak{D}'_a = \mathfrak{D}_a + X_a$, provided that the spin connection $\omega_{ab|m} = \omega_{ab|\mu}V_m^\mu$ transforms as $\omega^{ab}{}_{|m} \rightarrow \omega'^{ab}{}_{|m} = \omega^{ab}{}_{|m} - 2\delta_m^{[a}X^{b]}$, and with the choice $X_a = \frac{2}{3}t_a$, the torsion trace gets reabsorbed into the dilaton (we have $\mathfrak{D}'_a = \mathfrak{D}_a + \frac{2}{3}t_a$, and \mathfrak{D}'_a will be again renamed as \mathfrak{D}_a in the following). Thus, from now on we set

$$t_a = 0. \quad (3.26)$$

It follows that the necessary condition for conformal invariance [Eq. (3.23)] becomes

$$\left[R^{ab}{}_{lm} \left(Z^c{}_{pq} - \frac{1}{6} \epsilon^{cpqr} \tilde{\gamma}_r \right) \epsilon^{lm pq} \epsilon_{abcd} \right] \epsilon_K^d = 0. \quad (3.27)$$

A simple solution of Eq. (3.27) can be found by assuming

$$Z^c{}_{pq} = 0. \quad (3.28)$$

Using Eq. (3.28), Eq. (3.27) becomes

$$2\epsilon^{bdlm} R^a{}_{dlm} \tilde{\gamma}_a - \epsilon^{cdlm} R_{cdlm} \tilde{\gamma}^b = 0. \quad (3.29)$$

We will now show that Eq. (3.28) and the ensuing condition in Eq. (3.29) lead to intriguing physical consequences on the surviving field $\tilde{\gamma}_a$. Indeed, for $\tilde{\gamma}_a \neq 0$, a possible solution of Eq. (3.29) is¹¹

$$R_{a[bcd]} = 0, \quad R_{[abcd]} = 0. \quad (3.30)$$

We shall focus on this particular solution. The latter implies

$$R^{ab}{}_{(2)|cd} + R^{ab}{}_{(5)|cd} = 0, \quad R^{ab}{}_{(3)|cd} = 0. \quad (3.31)$$

Furthermore, recalling Eq. (3.4), namely

$$\check{R}_{ab} = R = 0 \Leftrightarrow R^{ab}{}_{(4)|cd} = R^{ab}{}_{(5)|cd} = R^{ab}{}_{(6)|cd} = 0, \quad (3.32)$$

and Eq. (3.5), and plugging all of this into the definition of R^{ab} in Eq. (2.6), one can easily realize that we are left with

$$\begin{aligned} R^{ab}{}_{(2)|cd} = 0 &\Rightarrow \mathcal{R}^{ab}{}_{(2)|cd} = 0, \\ R^{ab}{}_{(3)|cd} = 0 &\Rightarrow \mathcal{R}^{ab}{}_{(3)|cd} = 0, \\ R^{ab}{}_{(4)|cd} = 0, \quad R^{ab}{}_{(5)|cd} = 0, \quad R^{ab}{}_{(6)|cd} = 0, &\quad (3.33) \end{aligned}$$

together with

$$\begin{aligned} \mathcal{R}^{ab}{}_{(4)|cd} + \mathcal{R}^{ab}{}_{(5)|cd} + \mathcal{R}^{ab}{}_{(6)|cd} &\equiv 4\delta^{[a} S^{b]}{}_{d]}, \\ \mathcal{R}^{ab}{}_{(1)|cd} = \mathcal{R}^{ab}{}_{(1)|cd} &\equiv W^{ab}{}_{cd}. \end{aligned} \quad (3.34)$$

¹¹Here, let us mention that, in fact, using Eq. (3.28) with Eq. (3.27), the latter boils down to

$$R^{ab}{}_{lm} \tilde{\gamma}_r \epsilon^{cpqr} \epsilon^{lm pq} \epsilon_{abcd} = 0,$$

which can be simplified by contracting either first $\epsilon^{cpqr} \epsilon_{lm pq}$ and then the result with ϵ_{abcd} , or first $\epsilon^{cpqr} \epsilon_{abcd}$ and then the result with $\epsilon^{lm pq}$. The respectively obtained equations may appear different at first sight, but by exploiting the symmetry properties of the irreps of R_{abcd} , one can verify that they are actually equivalent, both exhibiting, in particular, Eq. (3.30) as a possible solution.

Hence, since now we have

$$\begin{aligned} R^{ab}{}_{cd} &= R^{ab}{}_{(1)cd} = W^{ab}{}_{cd} \equiv \mathcal{R}^{ab}{}_{cd} - \sum_{i=2}^6 \mathcal{R}^{ab}{}_{(i)|cd} \\ &= \mathcal{R}^{ab}{}_{cd} - 4\delta^{[a} S^{b]}{}_{d]}, \end{aligned} \quad (3.35)$$

we may write

$$\begin{aligned} R^{ab} &= W^{ab} \equiv \mathcal{R}^{ab} - 4V^{[a} S^{b]} \\ &= \mathcal{R}^{ab} - 2\check{R}^{[a|c} V_c V^{b]} + \frac{1}{3} \mathcal{R} V^a V^b, \end{aligned} \quad (3.36)$$

which is formally identical to the torsionless W^{ab} , but now ω^{ab} contains a torsion part.

In conclusion, we have recovered invariance under conformal boosts of Eq. (3.2) by solving the necessary condition for conformal invariance [Eq. (3.27)] under the assumption $Z^a{}_{bc} = 0$, the only nonvanishing part of the torsion T^a being given by

$$T^a = T^a{}_{bc} V^b V^c = -\frac{1}{6} \epsilon^{abcd} \tilde{\gamma}_d V^b V^c. \quad (3.37)$$

Inserting the actual form (3.37) into Eq. (2.6), we are led to

$$\omega_{ab|m} = \omega_{ab|\mu} V_m^\mu = \hat{\omega}_{ab|m} - 2\eta_{m[a} \mathfrak{D}_{b]} - \frac{1}{6} \epsilon_{abmc} \tilde{\gamma}^c, \quad (3.38)$$

where $\hat{\omega}_{ab|m} = \hat{\omega}_{ab|\mu} V_m^\mu$, and the last term in Eq. (3.38) is the contribution due to the contorsion term.¹² Moreover, from the variation of the torsion definition in Eq. (2.6), we now have

$$\begin{aligned} \delta\omega_{ab|m} &= (\delta_m^l \delta_{[a}^q \delta_{b]}^p + \delta_m^q \delta_{[a}^p \delta_{b]}^l - \delta_m^p \delta_{[a}^l \delta_{b]}^q) \hat{D}_q (\delta V^p)_l \\ &\quad - 2\eta_{m[a} \delta_{b]}^c \delta \mathfrak{D}_c - \frac{1}{6} \epsilon_{abmc} \delta \tilde{\gamma}^c - \frac{1}{6} \epsilon_{abpc} \tilde{\gamma}^c \delta_m^l (\delta V^p)_l, \end{aligned} \quad (3.39)$$

and one can verify that the transformations of the fields in Eq. (3.39) are such that the variation of $\omega_{ab|m}$ under dilatations and conformal boosts is the same whether one determines it from the gauge prescription [Eq. (2.9)] or directly from Eq. (3.39).¹³ Thus, the Lagrangian

¹²Recall also that $\hat{\omega}_{ab|\mu} = (f_{\lambda|\mu\nu} + f_{\nu|\lambda\mu} - f_{\mu|\nu\lambda}) V_a^\lambda V_b^\nu$, with $f_{\lambda|\mu\nu} = V_\lambda^k \partial_{[\mu} V_{\nu]}^c \eta_{ck}$.

¹³Indeed, under dilatations we have $\delta_{\epsilon_{\mathfrak{D}}} V^p{}_l = -\epsilon_{\mathfrak{D}} \delta_l^p$, $\delta_{\epsilon_{\mathfrak{D}}} \mathfrak{D}_c = \partial_c \epsilon_{\mathfrak{D}} = \hat{D}_c \epsilon_{\mathfrak{D}}$, and $\delta_{\epsilon_{\mathfrak{D}}} \tilde{\gamma}^c = \epsilon_{\mathfrak{D}} \tilde{\gamma}^c$, which plugged into Eq. (3.39) lead to $\delta_{\epsilon_{\mathfrak{D}}} \omega_{ab|m}(V, \mathfrak{D}, \tilde{\gamma}) = 0$ (in particular, the torsion contributions cancel each other out), reproducing the same result that can be obtained from the gauge prescription in Eq. (2.9). Analogously, under conformal boosts, one has $\delta_{\epsilon_K} V^p{}_l = 0$, $\delta_{\epsilon_K} \mathfrak{D}_c = 2\epsilon_{K|c}$, and $\delta_{\epsilon_K} \tilde{\gamma}^c = 0$, implying $\delta_{\epsilon_K} \omega_{ab|m}(V, \mathfrak{D}, \tilde{\gamma}) = -4\eta_{m[a} \epsilon_{K|b]}$, which is the same variation that one can obtain from Eq. (2.9).

(3.2) remains scale and proper conformal invariant if $\omega = \omega(V, \mathfrak{D}, \tilde{t})$, as given in Eq. (3.38).

If we now substitute Eq. (3.36) into Eq. (3.2), we obtain a Lagrangian that is formally identical to Eq. (3.16) but involving torsion—namely,

$$\begin{aligned} \mathcal{L} &= W^{ab}W^{cd}\epsilon_{abcd} = -W^{abcd}W_{abcd}\Omega^{(4)} \\ &= \mathcal{R}^{ab}\mathcal{R}^{cd}\epsilon_{abcd} - 8\left(\check{\mathcal{R}}_{ab}\check{\mathcal{R}}^{ba} - \frac{1}{3}\mathcal{R}^2\right)\Omega^{(4)}, \end{aligned} \quad (3.40)$$

the torsion being now hidden in the torsionful spin connection.

Furthermore, at the level of the theory in Eq. (3.2), the dilaton can be consistently eliminated. Indeed, also in the present case (as already pointed out in Ref. [1] in the case of vanishing torsion), the kinetic term for \mathfrak{D} does not contribute, since in Eq. (3.40) we have the combination $\check{\mathcal{R}}_{ab}\check{\mathcal{R}}^{ba}$ rather than $\check{\mathcal{R}}_{ab}\check{\mathcal{R}}^{ab}$ (whose presence would instead yield such a kinetic term for the dilaton), which makes a difference because $\check{\mathcal{R}}_{ab}(\omega)$ is not symmetric. A further check of the nonpropagating nature of the dilaton can be ascertained from its equation of motion. Indeed, even if we retain \mathfrak{D} in the Lagrangian, one can verify that its equation of motion is actually the trivial identity. With these arguments, one may set¹⁴

$$\mathfrak{D} = 0 \quad (3.41)$$

from the start, but it is not immediately obvious how the Lagrangian remains invariant in this case, since now the variation of $\omega_{ab|m}$ under dilatations and conformal boosts determined from Eq. (3.39) is not the same as the one determined from the gauge prescription [Eq. (2.9)] anymore. This problem was solved in Ref. [1] (for the torsion-free theory), showing that the additional terms present in the $\delta\omega^{ab}$ variation give a vanishing contribution. The same mechanism holds true in our case. Indeed, when varying the Lagrangian, we get an additional variation $\delta'\mathcal{L} = \frac{\delta\mathcal{L}}{\delta\omega}\delta\omega'$, where $\delta\omega'$ is the difference between the gauge variation of ω and the variation found from the explicit form of $\omega = \omega(V, \tilde{t})$. In particular, we get

$$\begin{aligned} \delta'\mathcal{L} &= -16\left(\frac{1}{6}\epsilon^{clnr}\tilde{t}_r K^d{}_p - \delta^c{}_l T^d{}_{np}\right)\epsilon^{lnpq}\epsilon_{qbcd}\xi^b \\ &\Rightarrow \delta'\mathcal{L} = 16\left(\frac{1}{3}\epsilon_{blnp}K^{ln}\tilde{t}^p + 4T^l{}_{bl}\right)\xi^b, \end{aligned} \quad (3.42)$$

where ξ_b is either $\partial_b\epsilon_{\mathfrak{D}}$ for dilations or $2\epsilon_K^c$ for conformal boosts. However, from the Bianchi identity of R^{ab} , which yields, in particular, $T^a{}_{la} = -\frac{1}{12}\epsilon_{labn}K^{ab}\tilde{t}^n$, one can see

¹⁴Notice that, in the present case, setting $\mathfrak{D} = 0$ we still have a *torsionful spin connection*, as the torsion axial vector \tilde{t}^a does not vanish.

that $\delta'\mathcal{L}$ in Eq. (3.42) vanishes identically. Hence, one can set the dilaton to zero without spoiling the conformal invariance of the theory, and Eq. (3.40)—that is, Weyl gravity with a non-vanishing axial vector torsion—results in the gauge theory of the conformal group with non-vanishing torsion.

Having eliminated the dilaton, we are left with Eqs. (3.36) and (3.40), where now the dilaton contributions vanish. In particular, now we have

$$G = 2V^a S_a \Rightarrow G_{ab} = \check{\mathcal{R}}_{[ab]}. \quad (3.43)$$

Recalling that $\check{\mathcal{R}}_{[ab]}$ is a function of \tilde{t}_a only and, in particular, through Eq. (3.38) we have

$$\check{\mathcal{R}}_{[ab]} = -\frac{1}{12}\epsilon_{abcd}\mathcal{D}^c\tilde{t}^d, \quad (3.44)$$

Eq. (3.43) becomes

$$G_{ab} = \check{\mathcal{R}}_{[ab]} = -\frac{1}{12}\epsilon_{abcd}\mathcal{D}^c\tilde{t}^d \rightarrow \check{\mathcal{R}}_{[ab]} = -\frac{1}{6}\star\mathbb{T}_{ab}, \quad (3.45)$$

where the star symbol denotes the Hodge duality operator and we have defined the field strength

$$\mathbb{T}_{ab} \equiv \mathcal{D}_{[a}\tilde{t}_{b]}. \quad (3.46)$$

We note that this result directly follows from the conformal invariance of the theory.

Furthermore, let us observe that the conformal invariance constraints in Eq. (3.30) imply

$$\mathcal{R}_{a[bcd]} = -2\eta_{a[d}S_{bc]} \rightarrow \mathcal{R}_{a[bcd]} = -\eta_{a[d}\check{\mathcal{R}}_{bc]}, \quad (3.47)$$

$$\epsilon^{abcd}\mathcal{R}_{abcd} = 0. \quad (3.48)$$

We now show that Eqs. (3.47) and (3.48) imply a differential constraint on the axial vector part of the torsion. Indeed, from Eq. (3.45), recalling that now we have $\omega = \omega(V, \tilde{t})$ —that is, the spin connection also involves a contorsion part—one can show that Eq. (3.48) reduces to

$$\mathcal{D}^a\tilde{t}_a = 0. \quad (3.49)$$

Then, in Eq. (3.47), we express both sides of the equation in terms of the torsion by exploiting the fact that now the curvature and antisymmetric Ricci tensors are given entirely in terms of the totally antisymmetric torsion. Indeed, recall that the connection $\omega_{ab|c}$ involves, besides the usual Riemannian part, a contorsion term \mathcal{K}_{abc} related to the completely antisymmetric torsion as follows:

$$V^b\mathcal{K}^a{}_b = T^a \Rightarrow \mathcal{K}_{abc} = -\frac{1}{6}\epsilon_{abcd}\tilde{t}^d. \quad (3.50)$$

As a consequence, using Eq. (3.49) with the explicit form of Eq. (3.47), one is left with

$$\mathcal{D}_{(a}\tilde{t}_{b)} = 0, \quad (3.51)$$

which is a Killing equation for the axial vector \tilde{t}_a of the present conformal theory, as Eq. (3.49) holds. Hence, from Eq. (3.30), we obtain Eqs. (3.49) and (3.51) for the axial vector torsion \tilde{t}_a . These imply, together, as can be proven by further differentiation and using the fact that $[\mathcal{D}_a, \mathcal{D}_b]\tilde{t}^m = -2\mathcal{R}^m{}_{nab}\tilde{t}^b - 2T^n{}_{ab}\mathcal{D}_n\tilde{t}^m = -2\mathcal{R}^m{}_{nab}\tilde{t}^n + \frac{1}{3}\epsilon_{abnd}\tilde{t}^d\mathcal{D}^n\tilde{t}^m$,

$$\begin{aligned} \square\tilde{t}_b - 2\check{\mathcal{R}}_{ab}\tilde{t}^a - \frac{1}{3}\epsilon_{abcd}\tilde{t}^a\mathbb{T}^{cd} &= 0 \\ \rightarrow \square\tilde{t}_b - 2\check{\mathcal{R}}_{ab}\tilde{t}^a + 4\check{\mathcal{R}}_{[ab]}\tilde{t}^a &= 0, \end{aligned} \quad (3.52)$$

where \square denotes the covariant d'Alembertian with torsion, $\square\tilde{t}_b \equiv \mathcal{D}^a\mathcal{D}_a\tilde{t}_b$.

Thus, as we can see from Eq. (3.52), in our theory the axial vector torsion obeys (3.52), and this follows directly from the requirement of invariance of the Lagrangian under conformal boosts. Equation (3.52) can be regarded as a Maxwell-like equation in a curved four-dimensional background.¹⁵ Note, however, that this is not an equation of motion derived from the Lagrangian, but just the result of having required the conformal invariance in the presence of an axial vector torsion.

As a last comment, let us observe that, taking into account all the results obtained till now, the Bianchi identities in Eq. (2.7) can be rewritten as

$$\begin{aligned} \text{Bianchi for } R^{ab} (R^{ab} = W^{ab}): \mathcal{D}W^{ab} + 4(T^{[a}S^{b]} - V^{[a}C^{b]}) \\ = 0 \Leftrightarrow \mathcal{D}R^{ab} = 0, \\ \text{Bianchi for } T^a: \mathcal{D}T^a + W^{ab}V_b - GV^a = 0, \\ \text{Bianchi for } \mathcal{T}^a (\mathcal{T}^a = C^a): \mathcal{D}C^a + \mathcal{R}^{ab}S_b = 0, \\ \text{Bianchi for } G: dG - 2T^aS_a + 2V^aC_a = 0, \end{aligned} \quad (3.53)$$

¹⁵Observe that the field strength \mathbb{T}_{ab} of the axial vector actually determines the antisymmetric part of the Ricci tensor of ω as given by Eq. (3.45). In particular, the latter means that for the antisymmetric part of the Schouten in Eqs. (3.5) and (3.6), we have $S_{[ab]} = \check{\mathcal{R}}_{[ab]} = -\frac{1}{6}\star\mathbb{T}_{ab} = -\frac{1}{12}\epsilon_{abcd}\mathcal{D}^c\tilde{t}^d$, where the (torsionful) covariant derivative \mathcal{D} here, specifically, contributes only with the purely Levi-Civita part, since the torsion term in \mathcal{D} cancels in this expression. Allowing also for a trace part of the torsion, one would get, besides this, bilinear torsion terms, but here we have already shown that the torsion trace can be consistently set to zero. Moreover, let us mention that the axial vector \tilde{t}^a also contributes in a nontrivial way to the symmetric part of the Schouten tensor—that is, it contributes to $S_{(ab)}$ with a term proportional to $\tilde{t}_a\tilde{t}_b$.

where we encounter the vector-valued Cotton 2-form defined as

$$\begin{aligned} C^a &\equiv \mathcal{D}S^a \rightarrow C_{a|bc}V^bV^c = -\mathcal{D}_cS_{ab}V^bV^c + S_{al}T^l{}_{bc}V^bV^c \\ &= -\mathcal{D}_cS_{ab}V^bV^c - \frac{1}{6}S_{al}\epsilon^l{}_{bcd}\tilde{t}^dV^bV^c. \end{aligned} \quad (3.54)$$

In the case in which the torsion vanishes, we recover the properties that the Schouten tensor is symmetric and that the completely antisymmetric and trace parts of the Cotton tensor are zero.

As a final check of the consistency of the theory, we can check from the previous equations that the Bianchi identities are identically satisfied. Indeed, analyzing the torsion Bianchi in Eq. (3.53) by also taking into account Eq. (3.45) together with Eqs. (3.49) and (3.51) [recall that Eqs. (3.45), (3.49), and (3.51) come from conformal invariance of the theory, *not* from the analysis of the field equations], we get $0 = 0$. One can prove, with some algebraic manipulation and making use of Eq. (3.52), coming from the requirement of conformal invariance, that the same happens for the Bianchi of G . The Bianchi for \tilde{T}^a becomes the Cotton Bianchi (for a connection with torsion), since \tilde{T}^a coincides with the Cotton 2-form, while the Bianchi identity for R^{ab} is the Bianchi for the Weyl 2-form W^{ab} and simply leads to $\mathcal{R}^{ab} = 0$, the left-hand side of the latter being identically zero for any connection (here with a nonvanishing torsion).¹⁶

One can thus see that the Bianchi identities are, as expected, true identities, and they do not add any additional constraint to the theory. This result [Eq. (3.53)] was expected also from the vacuum analysis we have previously done.

C. Equations of motion

Since we are adopting the second-order formalism for ω —that is, $\omega = \omega(V, \mathfrak{D}, \tilde{t})$ —in particular by fixing the form of the torsion, we shall vary the Lagrangian (3.40) with respect to the independent fields V and \tilde{t} .¹⁷

As we will see in a while, in this setup we get a propagation equation of the graviton that has four derivatives (∂^4V), as in the case with vanishing torsion, which is indeed expected for a conformal gravity theory.

¹⁶The Weyl Bianchi written in the form $\mathcal{D}R^{ab} = 0$ and the Cotton Bianchi look formally the same as in the case of vanishing torsion (see, for instance, Ref. [20]).

¹⁷In principle, the spin connection ω^{ab} has a dilaton dependence, but, as we have previously discussed, the dilaton field \mathfrak{D} can be eliminated from the theory. Actually, as one can easily verify, even if \mathfrak{D} is allowed to appear in the Lagrangian, exploiting the explicit form for the variation of ω^{ab} [Eq. (3.39)] and the fact that the Weyl tensor is traceless, its dynamics is trivial, as expected.

Now, let us observe that the following applies to our analysis:

$$\begin{aligned}\delta_{\Phi_i}\mathcal{L} &= \frac{\delta\mathcal{L}}{\delta\omega}\frac{\delta\omega}{\delta\Phi_i}\delta\Phi_i + \frac{\delta\mathcal{L}}{\delta\Phi_i}\delta\Phi_i, \quad i = 1, 2, \\ \Phi_1 &= \tilde{t}, \quad \Phi_2 = V,\end{aligned}\quad (3.55)$$

schematically. The term $\frac{\delta\mathcal{L}}{\delta\Phi_i}\delta\Phi_i$ corresponds to the explicit variation of the various fields in the Lagrangian. Therefore, we may start by computing $\delta_\omega\mathcal{L} = \frac{\delta\mathcal{L}}{\delta\omega}\delta\omega$, where \mathcal{L} is the torsionful Weyl Lagrangian [Eq. (3.40)]. Recalling Eq. (3.36) and using

$$\begin{aligned}\delta_\omega\mathcal{R}^{ab} &= \mathcal{D}(\delta\omega^{ab}) = \mathcal{D}_c(\delta\omega^{ab})_d V^c V^d, \\ \delta_\omega(\mathcal{R}^{ab}{}_{cd}\delta_b^d) &= \mathcal{D}_c(\delta\omega^{ab})_b, \\ \delta_\omega(\mathcal{R}^{ab}{}_{cd}\delta_b^d\delta_a^c) &= \mathcal{D}_a(\delta\omega^{ab})_b,\end{aligned}\quad (3.56)$$

after partial integration and using the fact that the Weyl tensor is completely traceless, we find

$$\delta_\omega\mathcal{L} = 8\delta\omega_{ab|lm}(\mathcal{D}_l W^{ablm})\Omega^{(4)}, \quad (3.57)$$

where we have also used the fact that

$$\begin{aligned}\mathcal{D}_b\Omega^{(4)}V^b &= \frac{2}{3!}\epsilon_{defg}T^d{}_{bk}V^bV^kV^eV^fV^g \\ &= -\frac{1}{3\cdot 3!}\epsilon_{dbkq}\tilde{t}^q\epsilon_{defg}\epsilon^{kefg}V^b\Omega^{(4)} = 0;\end{aligned}\quad (3.58)$$

that is,

$$\mathcal{D}_b\Omega^{(4)} = 0. \quad (3.59)$$

Regarding the variation with respect to \tilde{t}^c , one finds that the explicit variation does not contribute. Hence, from implicit variation, one obtains

$$\frac{\delta\mathcal{L}}{\delta\omega}\frac{\delta\omega}{\delta\tilde{t}^c}\delta\tilde{t}^c = -\frac{4}{3}\epsilon_{abmc}\delta\tilde{t}^c(\mathcal{D}_l W^{ablm})\Omega^{(4)} = 0, \quad (3.60)$$

which vanishes identically, since antisymmetrization of W_{abcd} on three indices gives identically zero.¹⁸

Concerning the explicit variation with respect to V , here one can prove that we have

$$\begin{aligned}W^{abcd} &= -W^{abdc} = -W^{bacd}, \\ W^{abcd} &= W^{cdab}, \quad W^{a[bcd]} = 0, \quad W^{[abcd]} = 0.\end{aligned}$$

Note that they hold true also in the presence of torsion, since the Weyl tensor is an irrep of $\text{SO}(1, 3)$.

$$\begin{aligned}\frac{\delta\mathcal{L}}{(\delta V^p)_l}(\delta V^p)_l &= -4(\delta V^p)_l\left(2\check{\mathcal{R}}_{ac}\delta_m^c\eta_{pb} - 2\check{\mathcal{R}}_{ap}\eta_{bm} - \frac{2}{3}\mathcal{R}\eta_{am}\eta_{bp}\right)W^{ablm}\Omega^{(4)} \\ &= -8(\delta V^p)_l\check{\mathcal{R}}_{ac}W^{aplc}\Omega^{(4)},\end{aligned}\quad (3.61)$$

where we have also observed that the second and third terms inside the round brackets in the first line give a vanishing contribution, since they imply a tracing of the Weyl tensor. Then, using Eqs. (3.39) and (3.57) (and performing a partial integration), and (3.61), together with the symmetry properties of the Weyl tensor and Eq. (3.59), after some algebraic manipulation, we find that the field equation $\delta_V\mathcal{L} = 0$ reads

$$\begin{aligned}(\delta V^p)_l\left(\mathcal{D}_q\mathcal{D}_l W^{lpql} - \frac{1}{2}\check{\mathcal{R}}_{ac}W^{aplc} - \frac{1}{12}\epsilon^{abpc}\tilde{t}_c\mathcal{D}_l W_{ab}{}^{tl}\right) \\ = 0;\end{aligned}\quad (3.62)$$

namely,

$$\mathcal{D}_q\mathcal{D}_l W^{lpql} - \frac{1}{2}\check{\mathcal{R}}_{ac}W^{aplc} - \frac{1}{12}\epsilon^{abpc}\tilde{t}_c\mathcal{D}_l W_{ab}{}^{tl} = 0. \quad (3.63)$$

Notice that the p, l trace of the latter identically vanishes, due to the symmetry properties of the Weyl tensor. The first two terms in Eq. (3.63) are *formally* the same as in the absence of torsion, thus giving a fourth-order equation for the vierbein. At the linearized level, the kinetic term is actually the same as in the absence of torsion, while at higher levels the presence of contorsion in the spin connection gives higher-order corrections.

Hence, in our theory, we get a fourth-order equation for the vierbein, and in the limit in which \tilde{t}_a is set to zero, we recover the conformal theory of Ref. [1].¹⁹

IV. ON THE EXTENSION TO CONFORMAL SUPERGRAVITY

In this section, we give some preliminary results concerning conformal $\mathcal{N} = 1$, $D = 4$ supergravity in the geometric approach. As we have already mentioned, in the literature, besides vanishing supertorsion, some constraints have been implemented to recover super-conformal invariance (the same constraints have been implemented in Ref. [12], within the geometric approach, through Lagrange multipliers).

¹⁹Let us observe that, as we have previously mentioned, the same Eqs. (3.52) and (3.63) can be obtained by gauging Eq. (3.15) and implementing the constraints we have presented in our analysis in order to recover conformal invariance.

¹⁸The symmetry properties of the Weyl tensor read

Here we start a preliminary analysis at the level of Bianchi identities using the geometric approach, showing that all the aforementioned constrained can be directly obtained from the study of the Bianchi identities, just imposing the vanishing of the supertorsion (which can be viewed as the direct supersymmetric extension of the constraint of vanishing torsion imposed in Ref. [1]).²⁰

For the benefit of the reader and in order to establish our formalism, let us just recall the main basic points of the geometric approach to supergravity.

The gauge fields are now super 1-forms in superspace that can be expanded along the supervierbein (V^a, ψ^α) , with $\alpha = 1, \dots, 4$ and ψ^α being the gravitino 1-form.²¹ (Note that in the geometric approach, the superfields are never expanded in terms of the Grassmann coordinates.) Analogously, the supercurvatures are 2-forms which can be expanded along the basis of 2-forms, namely

$$R^A = R^A{}_{ab} V^a V^b + R^A{}_{aa} V^a \psi^\alpha + R^A{}_{a\beta} \psi^\alpha \psi^\beta, \quad (4.1)$$

where $R^A{}_{aa}$ and $R^A{}_{a\beta}$ are the outer components of R^A , while $R^A{}_{ab}$ are the inner ones.²² The important point is that, both in the Lagrangian approach as well as in the Bianchi identities approach, it turns out that all the outer components of the curvatures can be expressed algebraically in terms of the inner ones, thus allowing for the elimination of the spurious unphysical degrees of freedom from the theory.²³ Actually, this can be shown from both the study of the geometric Lagrangian and the sector-by-sector analysis of the Bianchi “identities.” Within the latter approach, the Bianchi identities become relations to be analyzed, performing their split in the different sectors $\psi\psi\psi$, $\psi\psi V$, ψVV , and VVV . This gives the expression of the outer components of the supercurvatures in terms of the inner ones, causing the theory on superspace to have the same physical content as the theory on spacetime.

²⁰For the original formulation of the geometric approach to supergravity in superspace and, in particular, of its application to the study of the Bianchi identities in superspace, we refer the reader to Refs. [5,6] (see also Ref. [12] and Appendixes A and B of Ref. [21]). Moreover, a concise review of the prescriptions on the supercurvatures in the geometric approach to supergravity is also given in Appendix A of Ref. [22].

²¹Here we are considering $\mathcal{N} = 1$, $D = 4$. Spinor indices are denoted by α, β, \dots , and in the sequel we will frequently omit them to lighten the notation.

²²The outer components of the curvatures are defined as those having at least one index along the ψ direction of superspace, while the components with indices only along the bosonic vierbein are called “inner.”

²³The relation between outer and inner components of the supercurvatures is also referred to as the “rheonomy principle.” Actually, this property is a consequence of the fact that the Lagrangian is constructed only in terms of differential 4-forms in superspace, with the exclusion of the Hodge duality operator.

Finally, we mention that since supersymmetry transformations are just Lie derivatives in superspace, they are easily derived from the (superspace) Lie derivative of the gauge fields using the formula in footnote 7, namely $\delta_\epsilon \mu^A = D\epsilon^A + \iota_\epsilon R^A$, where $D\epsilon^A$ is a gauge transformation and the contraction is made with a supersymmetry parameter.

We shall now apply the aforementioned prescription on the Bianchi identities to the case of conformal supergravity with vanishing supertorsion.

A. Bianchi identities of the superconformal group with vanishing supertorsion

The superconformal algebra [23,24] is generated by the set $\{\mathbf{J}_{ab}, \mathbf{P}_a, \mathbf{K}_a, \mathbf{D}, \mathbf{A}, \mathbf{Q}_\alpha, \mathbf{Q}_\beta\}$. We introduce the 1-form fields $\{\omega^{ab}, V^a, K^a, \mathfrak{D}, A, \psi^\alpha, \phi^\alpha\}$ (see also Refs. [9,10,12]), respectively dual to the vector field generators of the superconformal algebra as given by Eq. (2.1) together with

$$A(\mathbf{A}) = 1, \quad \psi^\alpha(\mathbf{Q}_\beta) = \delta_\beta^\alpha, \quad \phi^\alpha(\mathbf{Q}_\beta) = \delta_\beta^\alpha. \quad (4.2)$$

The scale weights of the U(1) gauge 1-form field A , of the gravitino 1-form ψ , and of the conformino 1-form ϕ are, respectively,

$$[A] = 0, \quad [\psi] = \frac{1}{2}, \quad [\phi] = -\frac{1}{2}. \quad (4.3)$$

The supercurvatures associated with the superconformal algebra are

$$\begin{aligned} R^{ab} &\equiv \mathcal{R}^{ab} - 4V^{[a} K^{b]} + \bar{\psi}\gamma^{ab}\phi, \\ T^a &\equiv \mathcal{D}V^a + \mathfrak{D} \wedge V^a - \frac{i}{2}\bar{\psi}\gamma^a\psi = \hat{\mathcal{D}}V^a - \frac{i}{2}\bar{\psi}\gamma^a\psi, \\ \mathcal{T}^a &\equiv \mathcal{D}K^a - \mathfrak{D} \wedge K^a + \frac{i}{2}\bar{\phi}\gamma^a\phi = \hat{\mathcal{D}}K^a + \frac{i}{2}\bar{\phi}\gamma^a\phi, \\ G &\equiv d\mathfrak{D} + 2V^a K_a - \bar{\psi}\phi, \\ F &\equiv dA + 2i\bar{\psi}\gamma^5\phi, \\ \rho &\equiv \mathcal{D}\psi + \frac{1}{2}\mathfrak{D} \wedge \psi - \frac{3i}{4}A\gamma^5\psi - i\gamma_a\phi V^a \\ &= \hat{\mathcal{D}}\psi - \frac{3i}{4}A\gamma^5\psi - i\gamma_a\phi V^a = \nabla\psi - i\gamma_a\phi V^a, \\ \sigma &\equiv \mathcal{D}\phi - \frac{1}{2}\mathfrak{D} \wedge \phi + \frac{3i}{4}A\gamma^5\phi + i\gamma_a\psi K^a \\ &= \hat{\mathcal{D}}\phi + \frac{3i}{4}A\gamma^5\phi + i\gamma_a\psi K^a = \nabla\phi + i\gamma_a\psi K^a, \end{aligned} \quad (4.4)$$

where ψ^α and ϕ^α are the gravitino and conformino 1-forms, dual to ordinary supersymmetry and conformal supersymmetry, respectively. We recall that $\mathcal{D} = d - \omega$ is the Lorentz covariant derivative, $\hat{\mathcal{D}}$ is the Lorentz plus scale covariant derivative, and we have also taken the opportunity to introduce a Lorentz plus scale plus U(1) covariant derivative ∇ . The matrices γ^a , γ^{ab} , and γ^5 are the usual gamma matrices

in four dimensions. Useful formulas on gamma matrices can be found in the Appendix.

The Bianchi identities obeyed by the supercurvatures [Eq. (4.4)] are

$$\begin{aligned}
 DR^{ab} + 4(T^{[a}K^{b]} - V^{[a}T^{b]}) + \bar{\phi}\gamma^{ab}\rho + \bar{\psi}\gamma^{ab}\sigma &= 0, \\
 \hat{D}T^a + R^{ab}V_b - GV^a - i\bar{\psi}\gamma^a\rho &= 0, \\
 \hat{D}T^a + R^{ab}K_b + GK^a + i\bar{\phi}\gamma^a\sigma &= 0, \\
 dG - 2T^aK_a + 2V^aT_a - \bar{\psi}\sigma + \bar{\phi}\rho &= 0, \\
 dF + 2i\bar{\psi}\gamma^5\sigma - 2i\bar{\phi}\gamma^5\rho &= 0, \\
 \nabla\rho + \frac{1}{4}\gamma_{ab}R^{ab}\psi - \frac{1}{2}G\psi + \frac{3i}{4}F\gamma^5\psi + i\gamma_a\sigma V^a - i\gamma_a\phi T^a &= 0, \\
 \nabla\sigma + \frac{1}{4}\gamma_{ab}R^{ab}\phi + \frac{1}{2}G\phi - \frac{3i}{4}F\gamma^5\phi - i\gamma_a\rho K^a + i\gamma_a\psi T^a &= 0,
 \end{aligned} \tag{4.5}$$

where

$$\begin{aligned}
 \hat{D}T^a &\equiv DT^a + \mathfrak{D} \wedge T^a, \\
 \hat{D}T^a &\equiv DT^a - \mathfrak{D} \wedge T^a, \\
 \nabla\rho &\equiv \mathcal{D}\rho + \frac{1}{2}\mathfrak{D} \wedge \rho - \frac{3i}{4}A\gamma^5\rho = d\rho - \frac{1}{4}\gamma_{ab}R^{ab}\rho \\
 &\quad + \frac{1}{2}\mathfrak{D} \wedge \rho - \frac{3i}{4}A\gamma^5\rho, \\
 \nabla\sigma &\equiv \mathcal{D}\sigma - \frac{1}{2}\mathfrak{D} \wedge \sigma + \frac{3i}{4}A\gamma^5\sigma = d\sigma \\
 &\quad - \frac{1}{4}\gamma_{ab}R^{ab}\sigma - \frac{1}{2}\mathfrak{D} \wedge \sigma + \frac{3i}{4}A\gamma^5\sigma.
 \end{aligned} \tag{4.6}$$

One can now apply the prescription on the Bianchi identities to the present case—that is, writing the supercurvatures expansion as given in Eq. (4.1) and differentiating it to compare the result with the Bianchi (4.5) expanded along the supervierbein basis. The closure of the resulting system of equations must occur sector by sector—that is, along the $\psi\psi\psi$, $\psi\psi V$, ψVV , and VVV sectors separately.

Imposing vanishing supertorsion ($T^a = 0$) from the very beginning, a careful analysis shows that the superspace curvatures must have the following parametrization²⁴:

²⁴The scale of the supercurvature components along the 2-vierbein sector is

$$\begin{aligned}
 [R^{ab}_{cd}, G_{ab}, F_{ab}] &= -2, & [T^a_{bc}] &= -1, & [T^a_{bc}] &= -3, \\
 [\rho_{ab}] &= -\frac{3}{2}, & [\sigma_{ab}] &= -\frac{5}{2}.
 \end{aligned}$$

When doing the explicit calculations, one can immediately simplify the starting general ansatz by exploiting scale weight arguments.

$$\begin{aligned}
 R^{ab} &= R^{ab}_{cd}V^cV^d + 2i\bar{\psi}\gamma_c\rho^{ab}V^c, \\
 T^a &= 0, \\
 \mathcal{T}^a &= \mathcal{T}^a_{bc}V^bV^c \\
 &\quad + \bar{\psi}\left(-\frac{1}{2}\sigma^{ab} - \frac{i}{2}\gamma^5\star\sigma_{ab} + \gamma^{(a}\gamma_m\sigma^{m|b)}\right)V_b, \\
 G &= G_{ab}V^aV^b, \\
 F &= F_{ab}V^aV^b, \\
 \rho &= \rho_{ab}V^aV^b, \\
 \sigma &= \sigma_{ab}V^aV^b + \left(-\frac{i}{2}\star F_{ab}\gamma^b + \frac{1}{2}F_{ab}\gamma^5\gamma^b\right)\psi V^a,
 \end{aligned} \tag{4.7}$$

where for any 0-form $U_{ab} = -U_{ba}$ we have denoted the corresponding Hodge dual as

$$\star U_{ab} = \frac{1}{2}\epsilon_{abcd}U^{cd}. \tag{4.8}$$

As previously observed, the supersymmetry transformation laws differ from the gauge transformations when the curvatures exhibit at least a gravitino ψ in their parametrization. In particular, in the case at hand this happens for R^{ab} , \mathcal{T}^a , and σ , which indeed have a ψ in their parametrization (for the explicit form of the supersymmetry transformations of the fields, we refer the reader to Refs. [9,12]).

Let us recall here that the quantities R^{ab}_{cd} , \mathcal{T}^a_{bc} , G_{ab} , F_{ab} , ρ_{ab} , and σ_{ab} appearing in the parametrization (4.7) are the so-called *supercovariant field strengths*, and they differ, in general, from the spacetime projections of the supercurvatures. Indeed, let us refer, e.g., to the Lorentz supercurvature. Taking the components of R^{ab} along $dx^\mu \wedge dx^\nu$, namely $R^{ab}_{\mu\nu} = R^{ab}_{cd}V^c_\mu V^d_\nu + 2i\bar{\psi}_{[\mu}\gamma_c\rho^{ab}V^c_{\nu]}$, we see that the spacetime components $R^{ab}_{\mu\nu}$ differ from the components along the purely bosonic supervierbein, $R^{ab}_{cd}V^c_\mu V^d_\nu$. The quantity $R^{ab}_{cd}V^c_\mu V^d_\nu \equiv R^{ab}_{\mu\nu|(\text{cov})} = R^{ab}_{cd}V^c_\mu V^d_\nu + 2i\bar{\psi}_{[\mu}\gamma_c\rho^{ab}V^c_{\nu]}$ is the supercovariant field strength. The same happens for the curvatures \mathcal{T}^a and σ . Instead, as in the present case, the parametrizations of G , F , and ρ only have components along two vierbein. Covariant and supercovariant components on spacetime are identified—that is, we have $G_{\mu\nu} = G_{ab}V^a_\mu V^b_\nu$, $F_{\mu\nu} = F_{ab}V^a_\mu V^b_\nu$, and $\rho_{\mu\nu} = \rho_{ab}V^a_\mu V^b_\nu$.

Besides the given parametrizations, one also obtains the following constraints:

$$\begin{aligned}
 G_{ab} &= \frac{1}{2}\star F_{ab}, & \check{R}_{[ab]} &= -G_{ab} = -\frac{1}{2}\star F_{ab}, \\
 \check{R}_{(ab)} &= 0, & R &= 0,
 \end{aligned} \tag{4.9}$$

$$\begin{aligned}
 \gamma^{ab}\sigma_{ab} &= 0, \\
 \gamma^a(\sigma_{ab} - i\gamma^5\star\sigma_{ab}) &= 0, \\
 \gamma_{m[a}\sigma_{m|b]} &= \frac{1}{2}(\sigma_{ab} - i\gamma^5\star\sigma_{ab}), \quad (4.10)
 \end{aligned}$$

and

$$\gamma_{[a}\rho_{bc]} = 0 \rightarrow \gamma_c\rho_{ab} = -2\gamma_{[a}\rho_{b]c}, \quad (4.11)$$

the latter implying

$$\begin{aligned}
 \gamma^a\rho_{ab} = 0 &\Rightarrow \gamma^{ab}\rho_{ab} = 0, \\
 \rho_{ab} + i\gamma^5\star\rho_{ab} &= 0. \quad (4.12)
 \end{aligned}$$

Notice that by using the first line of Eq. (4.10), after some algebraic manipulation, we find that the second line of Eq. (4.10) reduces to the trivial identity $0 = 0$.

Let us just give a brief summary of the main steps of the cumbersome calculations used to recover the above results. The parametrization of R^{ab} , G , and F , together with the constraints in Eq. (4.12) and the fact that σ does not have components along two ψ 's, can be obtained by analyzing the $\psi\psi\psi$ sector of the Bianchi for R^{ab} , G , and F , together with the $\psi\psi V$ sector of the Bianchi for T^a and ρ . Considering the $\psi\psi V$ sector of the Bianchi identities for G and F together with the $\psi\psi\psi$ sector of the Bianchi for σ , the ψVV sector of the Bianchi for ρ , the $\psi\psi V$ sector of the Bianchi for R^{ab} , and the VVV sector of the supertorsion Bianchi, one finds Eq. (4.9) and the parametrization of σ . Finally, the parametrization of T^a and the equations in Eq. (4.10) can be obtained by analyzing the ψVV sector of the Bianchi for G and F together with the $\psi\psi V$ sector of the Bianchi for σ .

The above results are in perfect agreement with those of Refs. [9,10] and [12]. In particular, the constraints in Eq. (4.12) are the ones used in Refs. [9,10] (together with $T^a = 0$). Moreover, the constraint $\gamma^a\rho_{ab} = 0$ in Eq. (4.12) is the same constraint fixed in Ref. [12] by using Lagrange multipliers in the Lagrangian in order to recover superconformal invariance of the theory. We conclude that the solution given by Eqs. (4.7), (4.9), (4.10), and (4.12) gives exactly the same results as in Refs. [9,10] and [12]. There the constraints were required by physical arguments in order to consistently have supersymmetry invariance, while here we have shown that they are a mere consequence of the geometrical structure of the theory expressed by the closure of the Bianchi identities.

Observe that the constraints derived from the Bianchi identities turn out to be necessary for their closure, in a way quite analogous to the requirement that in the absence of auxiliary fields the closure of the supergravity Bianchi identities only holds when the equations of motion are satisfied. However, in conformal supergravity, the parametrizations for the curvatures and the constraints

recovered so far do not imply the equations of motion. One could then be surprised that we need constraints to have closure, since after all, Bianchi identities, when no equation of motion is needed, are true identities. The point is that the Bianchi identities would be true identities if we analyzed them in the enlarged *superconformal* coset of the basis gauge fields $(V^a, K^a, \psi^a, \phi^a)$, the other gauge fields ω^{ab} , \mathfrak{D} , and A being the factorized 1-forms dual to the generators belonging to the fiber. However, we want to have a physical theory on the *ordinary* supercoset spanned only by (V^a, ψ^a) , which is a *cotangent submanifold* of the enlarged superconformal coset. The geometric constraints of the theory are then interpreted as the requirement needed in order to have a consistent projection from the superconformal coset into the ordinary superspace.

The fact that the study of the Bianchi identities leads to the constraint of conformal supergravity has been also inferred in Ref. [14] in the context of an off-shell formulation of $N = 2$ supergravity with tensor multiplets. Here we have further highlighted and clarified the geometric origin and meaning of the superconformal constraints in the case of $\mathcal{N} = 1$, $D = 4$ conformal supergravity with $T^a = 0$, whose understanding is rather fundamental in view of a future analysis including a nonvanishing supertorsion in the theory.

Let us also mention that there are no independent differentials in the \mathbf{K}_a and \mathcal{Q}_a directions [as can be also deduced by looking at Eqs. (4.4) and (4.5)] so that one can write, also using scale weight arguments,

$$\begin{aligned}
 K^a &= K^a{}_b V^b + \bar{\psi}\kappa^a, \\
 \phi &= \phi_a V^a, \quad (4.13)
 \end{aligned}$$

where the 0-forms $K^a{}_b$, κ^a , and ϕ_a are a tensor, a spinor vector, and another spinor vector, respectively. Recall that $K^a{}_b$ coincides with the spacetime components of K^a only when $\psi \rightarrow 0$, but since we are now in superspace, the spacetime components of K^a are given entirely by the supercovariant part of K^a [that is, $K^a{}_b$ in Eq. (4.13)] plus the component along ψ . When one formulates the Lagrangian for the theory, the above components of K^a and ϕ can be determined by studying the field equations of the theory (and this could be particularly useful in our future study, where we will try to include a nonvanishing supertorsion).

On the other hand, the aforementioned components can be also obtained by expanding the supercurvature definitions given in Eq. (4.4) and using the geometric parametrization in Eq. (4.7). For the conformino components ϕ_a , we get

$$\phi_b = \frac{2}{3}\gamma^a \left(i\rho_{(0)|ab} - \frac{1}{2}\gamma^5\star\rho_{(0)|ab} \right), \quad (4.14)$$

where we have exploited $\gamma^a\rho_{ab} = 0$ from Eq. (4.12) and used the definition of ρ given in Eq. (4.4), taking its

2-vierbein sector [which is the only sector appearing in the parametrization of ρ in Eq. (4.7)]—that is,²⁵

$$\rho_{ab} = \rho_{(0)|ab} + i\gamma_{[a}\phi_{b]}, \quad \rho_{(0)|ab} \equiv \nabla_{[a}\psi_{b]}. \quad (4.15)$$

Equation (4.14) coincides, up to normalization and conventions, with the expression for the conformino ($\phi_\mu = \phi_b V_\mu^b$) found in Ref. [9]. Finally, notice that using the other results on ρ_{ab} given in Eq. (4.12) and making some algebraic manipulation, we also obtain

$$\rho_{(0)|ab} + i\gamma^5 \star \rho_{(0)|ab} = 0. \quad (4.16)$$

Thus, using Eq. (4.16) in Eq. (4.14), we are left with

$$\phi_b = \frac{i}{3} \gamma^a \rho_{(0)|ab}. \quad (4.17)$$

Similar arguments can be applied to find the expression for $K^a = K^a_\mu dx^\mu = K^a_b V^b + \tilde{\psi} \kappa^a$ by looking at the definition of R^{ab} in Eq. (4.4) and using the parametrization for R^{ab} in Eq. (4.7). More precisely, defining

$$\mathcal{R}_{(0)}^{ab} \equiv \mathcal{R}^{ab} + \tilde{\psi} \gamma^{ab} \phi \quad (4.18)$$

in such a way that fermionic contributions are taken into account in a straightforward way by means of $\mathcal{R}_{(0)}^{ab}$, we get

$$\begin{aligned} R^{ab}{}_{\mu\nu} - \mathcal{R}^{ab}{}_{(0)|\mu\nu} &= -4V^{[a}{}_{[\mu} K^{b]}{}_{\nu]} \\ \Rightarrow K_{\mu\nu} = K_{a\mu} V_\nu^a &= \frac{1}{2} \left(\check{\mathcal{R}}_{(0)|\mu\nu} - \frac{1}{6} g_{\mu\nu} \mathcal{R}_{(0)} \right) \\ &\quad - \frac{1}{2} \check{R}_{[\mu\nu]} - \frac{i}{2} \tilde{\psi}^\lambda \gamma_\nu \rho_{\lambda\mu} \\ &= \frac{1}{2} \left(\check{\mathcal{R}}_{(0)|\mu\nu} - \frac{1}{6} g_{\mu\nu} \mathcal{R}_{(0)} \right) \\ &\quad + \frac{1}{4} \star F_{\mu\nu} - \frac{i}{2} \tilde{\psi}^\lambda \gamma_\nu \rho_{\lambda\mu}. \end{aligned} \quad (4.19)$$

The latter coincides, up to normalization and conventions, with the same expression found in Ref. [9] for $K_{\mu\nu}$.

We have thus shown that at the supersymmetric level, setting the supertorsion to zero, all the other constraints necessary for superconformal invariance and implemented in Ref. [12] through Lagrange multipliers here actually

²⁵We note that since there are no components of ρ along the outer basis (ψ, V) and (ψ, ψ) , we have $\rho_{\mu\nu} = \rho_{ab} V_\mu^a V_\nu^b$, and therefore we can identify the a, b indices with spacetime anholonomic indices related to each other by the four-dimensional vierbein. This observation explains the meaning of the subsequent equation $\rho_{(0)|ab} \equiv \nabla_{[a}\psi_{b]}$ in Eq. (4.15), which would be senseless if a, b were interpreted as superspace indices along $V^a V^b$, since the 1-form ψ in superspace is independent of V^a by definition.

follow geometrically from the study of the various sectors of the Bianchi identities. Therefore, we expect the Bianchi identities to be a key feature in order to explore the possible construction of a conformal supergravity theory with a nonvanishing supertorsion, dictating, in this framework, the constraints that one must impose on the theory in order to recover superconformal invariance.

V. CONCLUSIONS

In this paper, we have shown that, in contrast with the claim of Ref. [1], it is actually possible to construct a gauge theory of the conformal group in four spacetime dimensions with a nonvanishing torsion component. In particular, we have allowed for a nonvanishing axial vector torsion and found a sufficient condition to write a gauge theory for the conformal group. In this setup, invariance under proper special conformal transformations (conformal boosts) implies a Killing equation for the axial vector torsion that, upon further differentiation, leads to a Maxwell-like propagation equation [precisely, to the d'Alembert equation (3.52)] in a curved background for the aforementioned axial vector. In the limit in which \tilde{t}_a is set to zero, we recover the conformal theory of Ref. [1].

We have then given some preliminary results regarding conformal $\mathcal{N} = 1, D = 4$ supergravity. In particular, we have shown that the constraints introduced in Ref. [12] in the geometric approach by the use of Lagrange multipliers in order to recover superconformal invariance of the theory can be, in fact, directly obtained in the same geometric approach from the study of the Bianchi identities, just assuming vanishing supertorsion.

This paves the way for future investigations that will be devoted to deepening the analysis on conformal supergravity (this work is currently in progress [4]). In particular, since we have seen that at the purely bosonic level there exists the possibility of introducing a nonvanishing, completely antisymmetric torsion without spoiling conformal invariance, we argue that something similar may also occur in the superconformal case. As a further remark, from a first glance we can say that (some of) the supersymmetric constraints arising from the requirement of superconformal invariance would certainly be different from the ones obtained in the case in which the supertorsion is set to zero in order to get a superconformal theory, and this in particular might cause something unexpected to happen.

Finally, let us say that our findings could also prove useful in the development of a possible four-dimensional extension of the theories presented in Refs. [25,26] in the context of modified/alternative theories of gravity. A detailed study in this direction could also unveil some peculiar features of non-Riemannian degrees of freedom, together with a clearer understanding of the potential relations occurring at the dynamical level among them.

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APPENDIX: USEFUL FORMULAS ON GAMMA MATRICES

We are working with Majorana spinors, satisfying $\bar{\lambda} = \lambda^T C$, where C is the charge conjugation matrix.

- (1) *Symmetric gamma matrices:* $C\gamma_a, C\gamma_{ab}, C\gamma_5\gamma_{ab}$.
- (2) *Antisymmetric gamma matrices:* $C, C\gamma_5, C\gamma_5\gamma_a$.
- (3) *Clifford algebra:*

$$\begin{aligned}
 \{\gamma_a, \gamma_b\} &= 2\eta_{ab}, & [\gamma_a, \gamma_b] &= 2\gamma_{ab}, & \gamma_5 &\equiv -i\gamma_0\gamma_1\gamma_2\gamma_3, \\
 \gamma_0^\dagger &= \gamma_0, & \gamma_0\gamma_i^\dagger\gamma_0 &= \gamma_i & (i = 1, 2, 3), & \gamma_5^\dagger &= \gamma_5, \\
 \epsilon_{abcd}\gamma^{cd} &= 2i\gamma_{ab}\gamma_5, & \gamma_{ab}\gamma_5 &= \gamma_5\gamma_{ab}, & \gamma_a\gamma_5 &= -\gamma_5\gamma_a, \\
 \gamma_m\gamma^{ab}\gamma^m &= 0, & \gamma_{ab}\gamma_m\gamma^{ab} &= 0, & \gamma_{ab}\gamma_{cd}\gamma^{ab} &= 4\gamma_{cd}, & \gamma_m\gamma^a\gamma^m &= -2\gamma^a, \\
 \gamma^a\gamma_a &= 4, & \gamma_b\gamma^{ab} &= -3\gamma^a, & \gamma^{ab}\gamma_b &= 3\gamma^a, \\
 \gamma^{ab}\gamma^c &= 2\gamma^{[a}\delta_c^{b]} + i\epsilon^{abcd}\gamma_5\gamma_d, & \gamma^c\gamma^{ab} &= -2\gamma^{[a}\delta_c^{b]} + i\epsilon^{abcd}\gamma_5\gamma_d, \\
 \gamma_{ab}\gamma_{cd} &= i\epsilon^{abcd}\gamma_5 - 4\delta_{[c}^{[a}\gamma_{d]}^{b]} - 2\delta_{cd}^{ab}.
 \end{aligned} \tag{A1}$$

- (4) *Useful Fierz identities for $\mathcal{N} = 1$* (for the 1-form spinor ψ):

$$\begin{aligned}
 \psi\bar{\psi} &= \frac{1}{4}\gamma_a\bar{\psi}\gamma^a\psi - \frac{1}{8}\gamma_{ab}\bar{\psi}\gamma^{ab}\psi, \\
 \gamma_a\psi\bar{\psi}\gamma^a\psi &= 0, \\
 \gamma_{ab}\psi\bar{\psi}\gamma^{ab}\psi &= 0.
 \end{aligned} \tag{A2}$$

Irreducible 3- ψ representations:

$$\begin{aligned}
 \Xi_{(12)}^a &\equiv \psi\bar{\psi}\gamma^a\psi, \\
 \Xi_{(8)}^{ab} &\equiv \psi\bar{\psi}\gamma^{ab}\psi + \gamma^{[a}\Xi_{(12)}^{b]}.
 \end{aligned} \tag{A3}$$

They satisfy $\gamma_a\Xi_{(12)}^a = 0$, $\gamma_a\Xi_{(8)}^{ab} = 0$, and we further have

$$\gamma_{ab}\psi\bar{\psi}\gamma^a\psi = -\gamma^a\psi\bar{\psi}\gamma_{ab}\psi = -\gamma_5\gamma^a\psi\bar{\psi}\gamma_{ab}\gamma_5\psi = \Xi_b^{(12)}. \tag{A4}$$

- (5) *Some spinor identities:*

$$\begin{aligned}
 \bar{\psi}\xi &= (-1)^{pq}\bar{\xi}\psi, \\
 \bar{\psi}(S)\xi &= -(-1)^{pq}\bar{\xi}(S)\psi, \\
 \bar{\psi}(AS)\xi &= (-1)^{pq}\bar{\xi}(AS)\psi,
 \end{aligned} \tag{A5}$$

where (S) is a symmetric matrix, while (AS) is an antisymmetric one; ψ and ξ are, respectively, a generic p -form spinor and a generic q -form spinor.

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