

Three-dimensional non-relativistic extended supergravity with cosmological constant

Original

Three-dimensional non-relativistic extended supergravity with cosmological constant / Concha, P.; Ravera, L.; Rodriguez, E.. - In: THE EUROPEAN PHYSICAL JOURNAL. C, PARTICLES AND FIELDS. - ISSN 1434-6044. - ELETTRONICO. - 80:12(2020), pp. 1-19. [10.1140/epjc/s10052-020-08685-2]

Availability:

This version is available at: 11583/2876405 since: 2022-01-12T12:44:53Z

Publisher:

Springer Science and Business Media Deutschland GmbH

Published

DOI:10.1140/epjc/s10052-020-08685-2

Terms of use:

This article is made available under terms and conditions as specified in the corresponding bibliographic description in the repository

Publisher copyright

Springer postprint/Author's Accepted Manuscript

This version of the article has been accepted for publication, after peer review (when applicable) and is subject to Springer Nature's AM terms of use, but is not the Version of Record and does not reflect post-acceptance improvements, or any corrections. The Version of Record is available online at: <http://dx.doi.org/10.1140/epjc/s10052-020-08685-2>

(Article begins on next page)



Three-dimensional non-relativistic extended supergravity with cosmological constant

Patrick Concha^{1,a} , Lucrezia Ravera^{2,3,b}, Evelyn Rodríguez^{4,c}

¹ Departamento de Matemática y Física Aplicadas, Universidad Católica de la Santísima Concepción, Alonso de Ribera 2850, Concepción, Chile

² DISAT, Politecnico di Torino, Corso Duca degli Abruzzi 24, 10129 Turin, Italy

³ INFN, Sezione di Torino, Via P. Giuria 1, 10125 Turin, Italy

⁴ Departamento de Ciencias, Facultad de Artes Liberales, Universidad Adolfo Ibáñez, Viña del Mar, Chile

Received: 29 August 2020 / Accepted: 18 November 2020
© The Author(s) 2020

Abstract In this paper, we present two novel non-relativistic superalgebras which correspond to supersymmetric extensions of the enlarged extended Bargmann algebra. The three-dimensional non-relativistic Chern–Simons supergravity actions invariant under the aforementioned superalgebras are constructed. The new non-relativistic superalgebras allow to accommodate a cosmological constant in a non-relativistic supergravity theory. Interestingly, we show that one of the non-relativistic supergravity theories presented here leads to the recently introduced Maxwellian exotic Bargmann supergravity when the flat limit $\ell \rightarrow \infty$ is considered. Besides, we show that both descriptions can be written in terms of a supersymmetric extension of the Nappi–Witten algebra or the extended Newton–Hooke superalgebra.

Contents

1	Introduction	1
2	Enlarged extended Bargmann gravity	2
3	Three-dimensional supergravity based on a supersymmetric extension of the Enlarged Extended Bargmann algebra	3
3.1	Supersymmetric extension of the Enlarged Extended Bargmann algebra	3
3.2	Non-relativistic Chern–Simons supergravity action	4
4	Three-dimensional non-standard enlarged extended Bargmann supergravity	4
4.1	Non-standard enlarged extended Bargmann superalgebra	4

4.2	Non-standard non-relativistic extended supergravity action	4
5	Discussion	5
Appendix A:	Invariant tensor and semigroup expansion method	5
Appendix B:	Gauge transformations	5
References		5

1 Introduction

Non-relativistic (NR) supergravity theories have recently been studied and their construction has only been approached in three spacetime dimensions [1–6]. In this way, to formulate a NR (super)gravity theory invariant under a certain NR (super)symmetry remains as a challenging but at the same time interesting problem. At the bosonic level, the NR theories have been of particular interest due to their relation to condensed matter systems [7–17] and NR effective field theories [18–22]. In presence of supersymmetry, moving toward this kind of theories can be useful, for instance, to approach supersymmetric field theories on non-relativistic curved backgrounds via localization [23–25]. Also from the purely theoretical point of view, NR supergravity is an attractive topic that deserves to be developed. The construction of NR (super)algebras and the formulation of NR (super)gravity theories are of special interest for the present work. As it happens in NR gravity theories, the suitable construction of a NR supergravity action has been shown to require the introduction of extra bosonic and fermionic generators. In particular, working in three spacetime dimensions, the appropriate construction of a Chern–Simons (CS) action is possible due to the presence of additional generators which allow to have a non-degenerate invariant tensor. The non-degeneracy of the

^a e-mail: patrick.concha@ucsc.cl (corresponding author)

^b e-mail: lucrezia.ravera@polito.it

^c e-mail: evelyn.rodriguez@edu.uai.cl

invariant tensor is a fundamental ingredient to get a well-defined CS (super)gravity action [26–28].

Recently, the NR version of a three-dimensional CS gravity theory invariant under a particular enlargement of the extended Bargmann algebra [29–35] was presented in [36]. Such algebra, called as enlarged extended Bargmann (EEB) algebra by the authors, was obtained by considering the NR contraction of the $[\text{AdS-Lorentz}] \oplus \mathfrak{u}_1 \oplus \mathfrak{u}_1 \oplus \mathfrak{u}_1$ algebra. Such novel NR symmetry offers us an alternative way to accommodate a cosmological constant to a three-dimensional NR gravity theory different from those studied in [37–45]. Interestingly, as was shown in [36], the Maxwellian exotic Bargmann (MEB) algebra recently introduced in [46] can be obtained from the EEB gravity theory when the flat limit $\ell \rightarrow \infty$ is considered (where ℓ is the length parameter related to the cosmological constant through $\Lambda \propto \pm \frac{1}{\ell^2}$; the limit $\ell \rightarrow \infty$ corresponds to $\Lambda \rightarrow 0$).

The AdS-Lorentz (AdS- \mathcal{L}) (super)symmetry has been introduced in [47–49] and has been useful in three and higher dimensions in diverse (super)gravity contexts. On the gravity side, this symmetry and its generalizations were used to recover (pure) Lovelock gravities from CS and Born-Infeld theories in different odd and even dimensions, respectively [50–52]. Subsequently, in the case of a three-dimensional gravity theory invariant under the AdS- \mathcal{L} algebra, it has been shown that the asymptotic symmetry of conserved charges at null infinity is given by a semi-simple enlargement of \mathfrak{bms}_3 [53]. On the other hand, the supersymmetric extension of the AdS- \mathcal{L} algebra allowed to introduce a generalized supersymmetric cosmological constant term in a four-dimensional supergravity theory [54–58]. In three space-time dimensions, the construction of the \mathcal{N} -extended AdS- \mathcal{L} supergravity theory was presented in [59, 60].

As pointed out in [36], an open problem was to extend their results at the supersymmetric level. Very recently, a NR version of the Maxwell CS supergravity theory was presented in [6] while the construction of a possible EEB supergravity theory remained unexplored till now. In this work, we approach the aforesaid problem, and present two different supersymmetric extensions of the EEB algebra by expanding different relativistic $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebras. In particular, the Lie algebra expansion method [61–64] has resulted to be a powerful tool in NR symmetries [36, 65–75]. Here, we show that the semigroup expansion method [64] allows us to obtain not only new NR superalgebras but also provides with the non-vanishing components of the invariant tensor allowing to construct the corresponding NR CS supergravity actions. We show that, although both supersymmetric descriptions allow us to introduce a cosmological constant to a NR supergravity action, only one supersymmetric extension contains a well-defined vanishing cosmological constant limit.

The paper is organized as follows: in Sect. 2, a brief review of the EEB algebra and the corresponding formulation of

the three-dimensional CS gravity action are presented. Sections 3 and 4 contain the main results of the work. First, a supersymmetric extension of the EEB is presented. Then, a NR CS supergravity action based on the EEB superalgebra is constructed. The supersymmetry transformation laws are also provided. In Sect. 4, we provide with an alternative supersymmetric extension of the EEB algebra which we call non-standard EEB superalgebra. The construction of the CS supergravity action is also studied. Section 5 concludes our work with discussions and some possible future approaches.

2 Enlarged extended Bargmann gravity

In this section, we briefly review the enlarged extended Bargmann algebra considered in [36] and the associated CS gravity theory constructed in the same paper in three space-time dimensions. The particular enlargement of the Extended Bargmann algebra presented in [36] was denoted as EEB algebra and its non-trivial commutation relations read

$$\begin{aligned} [\tilde{G}_a, \tilde{P}_b] &= -\epsilon_{ab} \tilde{M}, & [\tilde{G}_a, \tilde{Z}_b] &= -\epsilon_{ab} \tilde{T}, \\ [\tilde{H}, \tilde{Z}_a] &= \frac{1}{\ell^2} \epsilon_{ab} \tilde{P}_b, \\ [\tilde{H}, \tilde{G}_a] &= \epsilon_{ab} \tilde{P}_b, & [\tilde{J}, \tilde{Z}_a] &= \epsilon_{ab} \tilde{Z}_b, \\ [\tilde{Z}_a, \tilde{Z}_b] &= -\frac{1}{\ell^2} \epsilon_{ab} \tilde{T}, \\ [\tilde{J}, \tilde{P}_a] &= \epsilon_{ab} \tilde{P}_b, & [\tilde{H}, \tilde{P}_a] &= \epsilon_{ab} \tilde{Z}_b, \\ [\tilde{P}_a, \tilde{Z}_b] &= -\frac{1}{\ell^2} \epsilon_{ab} \tilde{M}, \\ [\tilde{J}, \tilde{G}_a] &= \epsilon_{ab} \tilde{G}_b, & [\tilde{P}_a, \tilde{P}_b] &= -\epsilon_{ab} \tilde{T}, \\ [\tilde{Z}, \tilde{P}_a] &= \frac{1}{\ell^2} \epsilon_{ab} \tilde{P}_b, \\ [\tilde{G}_a, \tilde{G}_b] &= -\epsilon_{ab} \tilde{S}, & [\tilde{Z}, \tilde{G}_a] &= \epsilon_{ab} \tilde{Z}_b, \\ [\tilde{Z}, \tilde{Z}_a] &= \frac{1}{\ell^2} \epsilon_{ab} \tilde{Z}_b, \end{aligned} \quad (2.1)$$

where $a = 1, 2$, $\epsilon_{ab} = \epsilon_{0ab}$ and $\epsilon^{ab} = \epsilon^{0ab}$, and where ℓ is a length parameter. Such algebra turns out to be the NR version of the relativistic $[\text{AdS-}\mathcal{L}] \oplus \mathfrak{u}(1) \oplus \mathfrak{u}(1) \oplus \mathfrak{u}(1)$ algebra. The AdS- \mathcal{L} algebra reads

$$\begin{aligned} [J_A, J_B] &= \epsilon_{ABC} J^C, & [J_A, P_B] &= \epsilon_{ABC} P^C, \\ [J_A, Z_B] &= \epsilon_{ABC} Z^C, & [P_A, P_B] &= \epsilon_{ABC} Z^C, \\ [Z_A, Z_B] &= \frac{1}{\ell^2} \epsilon_{ABC} Z^C, & [P_A, Z_B] &= \frac{1}{\ell^2} \epsilon_{ABC} P^C, \end{aligned} \quad (2.2)$$

with $A = 0, 1, 2$, where J_A are the Lorentz generators, P_A the spacetime translation generators, and Z_A are extra Abelian charges. The aforementioned $U(1)$ -enlargement (given in

terms of the $U(1)$ generators Y_1 , Y_2 , and Y_3) of the relativistic AdS- \mathcal{L} symmetry is necessary to assure a finite and non-degenerate invariant tensor allowing to construct a well-defined NR CS gravity theory in three spacetime dimensions. As it was shown in [36], the EEB algebra (2.1) appears as a contraction of the relativistic algebra (2.2). Indeed, the NR algebra is obtained considering the limit $\xi \rightarrow \infty$ in the following redefinition of the $[\text{AdS-}\mathcal{L}] \oplus \mathfrak{u}(1) \oplus \mathfrak{u}(1) \oplus \mathfrak{u}(1)$ generators:

$$\begin{aligned} J_0 &= \frac{\tilde{J}}{2} + \xi^2 \tilde{S}, & J_a &= \xi \tilde{G}_a, & Y_2 &= \frac{\tilde{J}}{2} - \xi^2 \tilde{S}, \\ P_0 &= \frac{\tilde{H}}{2\xi} + \xi \tilde{M}, & P_a &= \tilde{P}_a, & Y_1 &= \frac{\tilde{H}}{2\xi} - \xi \tilde{M}, \\ Z_0 &= \frac{\tilde{Z}}{2\xi^2} + \tilde{T}, & Z_a &= \frac{\tilde{Z}_a}{\xi}, & Y_3 &= \frac{\tilde{Z}}{2\xi^2} - \tilde{T}, \end{aligned} \quad (2.3)$$

together with the rescaling of the length parameter $\ell \rightarrow \xi \ell$. Furthermore, as in the case of its relativistic counterpart, a vanishing cosmological constant limit $\ell \rightarrow \infty$ can be applied to the EEB algebra (2.1), leading to the Maxwellian Extended Bargmann (MEB) algebra of [46].

The EEB algebra admits the following non-vanishing components of the invariant tensor:

$$\begin{aligned} \langle \tilde{G}_a \tilde{G}_b \rangle &= \tilde{\alpha}_0 \delta_{ab}, & \langle \tilde{G}_a \tilde{P}_b \rangle &= \tilde{\alpha}_1 \delta_{ab}, \\ \langle \tilde{G}_a \tilde{Z}_b \rangle &= \tilde{\alpha}_2 \delta_{ab}, & \langle \tilde{P}_a \tilde{P}_b \rangle &= \tilde{\alpha}_2 \delta_{ab}, \\ \langle \tilde{J} \tilde{S} \rangle &= -\tilde{\alpha}_0, & \langle \tilde{J} \tilde{M} \rangle &= -\tilde{\alpha}_1, \\ \langle \tilde{H} \tilde{S} \rangle &= -\tilde{\alpha}_1, & \langle \tilde{J} \tilde{T} \rangle &= -\tilde{\alpha}_2, \\ \langle \tilde{H} \tilde{M} \rangle &= -\tilde{\alpha}_2, & \langle \tilde{S} \tilde{Z} \rangle &= -\tilde{\alpha}_2, \\ \langle \tilde{Z}_a \tilde{Z}_b \rangle &= \frac{\tilde{\alpha}_2}{\ell^2} \delta_{ab}, & \langle \tilde{Z}_a \tilde{P}_b \rangle &= \frac{\tilde{\alpha}_1}{\ell^2} \delta_{ab}, \\ \langle \tilde{Z} \tilde{M} \rangle &= -\frac{\tilde{\alpha}_1}{\ell^2}, & \langle \tilde{T} \tilde{H} \rangle &= -\frac{\tilde{\alpha}_1}{\ell^2}, \\ \langle \tilde{Z} \tilde{T} \rangle &= -\frac{\tilde{\alpha}_2}{\ell^2}, \end{aligned} \quad (2.4)$$

where $\tilde{\alpha}_0$, $\tilde{\alpha}_1$, and $\tilde{\alpha}_2$ are arbitrary constants which are related to the relativistic parameters as

$$\alpha_0 = \tilde{\alpha}_0 \xi^2, \quad \alpha_1 = \tilde{\alpha}_1 \xi, \quad \alpha_2 = \tilde{\alpha}_2. \quad (2.5)$$

Such rescaling of the coupling constants reproduces the most general NR non-degenerate invariant tensor in the limit $\xi \rightarrow \infty$. Let us note that the non-degeneracy of the invariant tensor is related to the Physical requirement that the CS action involves a kinematical term for each field and the equation of motions imply that all curvatures vanish. On the other hand, one can see that the limit $\ell \rightarrow \infty$ applied to the components

of the invariant tensor written in (2.4) leads to the invariant tensor of the MEB algebra [46].

The CS form for a connection A constructed with the invariant non-degenerate bilinear form defines an action for the relativistic gauge theory based on the symmetry under consideration as

$$I_{\text{CS}} = \int \left\langle A \wedge dA + \frac{2}{3} A \wedge A \wedge A \right\rangle = \int \left\langle A \wedge dA + \frac{1}{3} A \wedge [A, A] \right\rangle. \quad (2.6)$$

In terms of the NR generators and fields of the EEB algebra, the gauge connection one-form of [36], $A = A^A T_A$, where $T_A = \{\tilde{J}, \tilde{G}_a, \tilde{H}, \tilde{P}_a, \tilde{Z}, \tilde{Z}_a, \tilde{M}, \tilde{S}, \tilde{T}\}$, is given by

$$A = \omega \tilde{J} + \omega^a \tilde{G}_a + \tau \tilde{H} + e^a \tilde{P}_a + k \tilde{Z} + k^a \tilde{Z}_a + m \tilde{M} + s \tilde{S} + t \tilde{T}. \quad (2.7)$$

The EEB curvature two-form is then

$$\begin{aligned} F &= R(\omega) \tilde{J} + R^a(\omega^b) \tilde{G}_a + R(\tau) \tilde{H} + R^a(e^b) \tilde{P}_a \\ &\quad + R(k) \tilde{Z} + R^a(k^b) \tilde{Z}_a + R(m) \tilde{M} \\ &\quad + R(s) \tilde{S} + R(t) \tilde{T}, \end{aligned} \quad (2.8)$$

with¹

$$\begin{aligned} R(\omega) &= d\omega, \\ R^a(\omega^b) &= d\omega^a + \epsilon^{ac} \omega \omega_c, \\ R(\tau) &= d\tau, \\ R^a(e^b) &= de^a + \epsilon^{ac} \omega e_c + \epsilon^{ac} \tau \omega_c + \frac{1}{\ell^2} \epsilon^{ac} k e_c + \frac{1}{\ell^2} \epsilon^{ac} \tau k_c, \\ R(k) &= dk, \\ R^a(k^b) &= dk^a + \epsilon^{ac} \omega k_c + \epsilon^{ac} \tau e_c + \epsilon^{ac} k \omega_c + \frac{1}{\ell^2} \epsilon^{ac} k k_c, \\ R(m) &= dm + \epsilon^{ac} e_a \omega_c + \frac{1}{\ell^2} \epsilon^{ac} e_a k_c, \\ R(s) &= ds + \frac{1}{2} \epsilon^{ac} \omega_a \omega_c, \\ R(t) &= dt + \epsilon^{ac} \omega_a k_c + \frac{1}{2} \epsilon^{ac} e_a e_c + \frac{1}{2\ell^2} \epsilon^{ac} k_a k_c. \end{aligned} \quad (2.9)$$

Then, the NR three-dimensional CS action based on the EEB algebra [36], up to boundary terms, is obtained considering the gauge connection one-form (2.7) and the non-vanishing components of the invariant tensor (2.4) in the general expression of the CS action (2.6). The NR CS graviy action reads

$$\begin{aligned} I_{\text{EEB}} &= \int \tilde{\alpha}_0 \left[\omega_a R^a(\omega^b) - 2s R(\omega) \right] \\ &\quad + \tilde{\alpha}_1 \left[e_a R^a(\omega^b) + \omega_a R^a(e^b) - 2m R(\omega) - 2\tau ds \right] \end{aligned}$$

¹ In the sequel, we will omit the wedge product \wedge between differential forms.

$$\begin{aligned}
& + \frac{1}{\ell^2} e_a R^a(k^b) + \frac{1}{\ell^2} k_a R^a(e^b) - \frac{2}{\ell^2} \tau dt \\
& - \frac{2}{\ell^2} m R(k) \Big] + \tilde{\alpha}_2 \Big[e_a R^a(e^b) + k_a R^a(\omega^b) \\
& + \omega_a R^a(k^b) + \frac{1}{\ell^2} k_a R^a(k^b) \\
& - 2s R(k) - 2m R(\tau) - 2t R(\omega) - \frac{2}{\ell^2} t R(k) \Big].
\end{aligned} \quad (2.10)$$

The CS gravity theory based on the EEB symmetry can be seen as an alternative NR gravity model in presence of a cosmological constant. The NR CS action (2.10) contains three sectors proportional to different arbitrary constants $\tilde{\alpha}_0$, $\tilde{\alpha}_1$, and $\tilde{\alpha}_2$. The first contribution, proportional to $\tilde{\alpha}_0$, is the CS action for the NR Exotic Gravity. The second and third term, proportional, respectively, to $\tilde{\alpha}_1$ and $\tilde{\alpha}_2$, reproduce the enlarged extended Bargmann gravity with the explicit presence of the k_a gauge field. Observe that the limit $\ell \rightarrow \infty$ taken in the term proportional to $\tilde{\alpha}_1$ reproduces the CS action for the Extended Bargmann algebra [3]. On the other hand, the limit $\ell \rightarrow \infty$ in the sector proportional to $\tilde{\alpha}_2$ leads to the CS action for the NR Maxwell algebra [46]. Let us note that the term proportional to $\tilde{\alpha}_1$ is not the extended Newton–Hooke gravity Lagrangian, although it leads to the extended Bargmann gravity Lagrangian in the $\ell \rightarrow \infty$ limit. In particular, the additional gauge fields related to the EEB algebra appearing in the $\tilde{\alpha}_1$ term vanish in the flat ($\ell \rightarrow \infty$) limit.

3 Three-dimensional supergravity based on a supersymmetric extension of the Enlarged Extended Bargmann algebra

Here, we present a supersymmetric extension of the EEB algebra which allows us to construct a NR supergravity action. Consequently, we develop the aforementioned NR supergravity action by exploiting the CS construction in three dimensions. In order to have a proper NR CS supergravity action based on a supersymmetric extension of the EEB algebra, one requires to find a NR superalgebra which not only contains the EEB algebra as a subalgebra but also admits an invariant supertrace. Such task can be properly accomplished by considering the S -expansion procedure [64] (see also [76, 77] for recent developments on the same method; more details on the S -expansion procedure are given in Appendix A). Indeed, an EEB superalgebra can be obtained by applying the S -expansion to the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra [60] with a suitable semigroup S .

3.1 Supersymmetric extension of the Enlarged Extended Bargmann algebra

The relativistic $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra is characterized by the presence of the AdS- \mathcal{L} bosonic generators $\{J_A, P_A, Z_A\}$, the fermionic spinor charges $\{Q_\alpha^i, \Sigma_\alpha^i\}$, and three $\mathfrak{so}(2)$ generators T, \mathcal{U} , and \mathcal{B} . Such generators satisfy the following non-vanishing commutation relations [60]:

$$\begin{aligned}
[J_A, J_B] &= \epsilon_{ABC} J^C, & [J_A, P_B] &= \epsilon_{ABC} P^C, \\
[J_A, Z_B] &= \epsilon_{ABC} Z^C, & [P_A, P_B] &= \epsilon_{ABC} Z^C, \\
[P_A, P_B] &= \epsilon_{ABC} Z^C, & [Z_A, Z_B] &= \frac{1}{\ell^2} \epsilon_{ABC} Z^C, \\
[P_A, Z_B] &= \frac{1}{\ell^2} \epsilon_{ABC} P^C, \\
[J_A, Q_\alpha^i] &= -\frac{1}{2} (\gamma_A)_\alpha^\beta Q_\beta^i, & [J_A, \Sigma_\alpha^i] &= -\frac{1}{2} (\gamma_A)_\alpha^\beta \Sigma_\beta^i, \\
[P_A, Q_\alpha^i] &= -\frac{1}{2} (\gamma_A)_\alpha^\beta \Sigma_\beta^i, \\
[P_A, \Sigma_\alpha^i] &= -\frac{1}{2\ell^2} (\gamma_A)_\alpha^\beta Q_\beta^i, & [Z_A, Q_\alpha^i] &= -\frac{1}{2\ell^2} (\gamma_A)_\alpha^\beta Q_\beta^i, \\
[Z_A, \Sigma_\alpha^i] &= -\frac{1}{2\ell^2} (\gamma_A)_\alpha^\beta \Sigma_\beta^i, \\
[T, Q_\alpha^i] &= \frac{1}{2} \epsilon^{ij} Q_\alpha^j, & [T, \Sigma_\alpha^i] &= \frac{1}{2} \epsilon^{ij} \Sigma_\alpha^j, \\
[\mathcal{U}, Q_\alpha^i] &= \frac{1}{2} \epsilon^{ij} \Sigma_\alpha^j, \\
[\mathcal{U}, \Sigma_\alpha^i] &= \frac{1}{2\ell^2} \epsilon^{ij} Q_\alpha^j, & [\mathcal{B}, Q_\alpha^i] &= \frac{1}{2\ell^2} \epsilon^{ij} Q_\alpha^j, \\
[\mathcal{B}, \Sigma_\alpha^i] &= \frac{1}{2\ell^2} \epsilon^{ij} \Sigma_\alpha^j,
\end{aligned} \quad (3.1)$$

along with the following anti-commutators,

$$\begin{aligned}
\{Q_\alpha^i, Q_\beta^j\} &= -\delta_{ij} (\gamma^A C)_{\alpha\beta} P_A - C_{\alpha\beta} \epsilon^{ij} \mathcal{U}, \\
\{Q_\alpha^i, \Sigma_\beta^j\} &= -\delta_{ij} (\gamma^A C)_{\alpha\beta} Z_A - C_{\alpha\beta} \epsilon^{ij} \mathcal{B}, \\
\{\Sigma_\alpha^i, \Sigma_\beta^j\} &= -\frac{\delta_{ij}}{\ell^2} (\gamma^A C)_{\alpha\beta} P_A - \frac{1}{\ell^2} C_{\alpha\beta} \epsilon^{ij} \mathcal{U},
\end{aligned} \quad (3.2)$$

where $A, B, C = 0, 1, 2$ are the Lorentz indices which are raised and lowered with the Minkowski metric $\eta_{AB} = (-1, 1, 1)$, $\alpha, \beta = 1, 2$,² and $i, j = 1, 2$ label the number of supercharges. Here, γ^A are Dirac matrices in three spacetime dimensions and C is the charge conjugation matrix,

$$C_{\alpha\beta} = C^{\alpha\beta} = \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}, \quad (3.3)$$

which satisfies $C^T = -C$ and $C\gamma^A = (C\gamma^A)^T$. As it was discussed in [60], the presence of $\mathfrak{so}(2)$ symmetry generators is not arbitrary and assures the non-degeneracy of the

² In the following, we will frequently omit these spinorial indices in order to lighten the notation.

invariant tensor. In particular, the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra admits the following non-vanishing components of an invariant tensor,

$$\begin{aligned} \langle J_A J_B \rangle &= \alpha_0 \eta_{AB}, \quad \langle J_A P_B \rangle = \alpha_1 \eta_{AB}, \quad \langle J_A Z_B \rangle = \alpha_2 \eta_{AB}, \\ \langle P_A P_B \rangle &= \alpha_2 \eta_{AB}, \quad \langle P_A Z_B \rangle = \frac{\alpha_1}{\ell^2} \eta_{AB}, \quad \langle Z_A Z_B \rangle = \frac{\alpha_2}{\ell^2} \eta_{AB}, \\ \langle \mathcal{T} \mathcal{T} \rangle &= \alpha_0, \quad \langle \mathcal{T} \mathcal{U} \rangle = \alpha_1, \quad \langle \mathcal{T} \mathcal{B} \rangle = \alpha_2, \\ \langle \mathcal{U} \mathcal{U} \rangle &= \alpha_2, \quad \langle \mathcal{U} \mathcal{B} \rangle = \frac{\alpha_1}{\ell^2}, \quad \langle \mathcal{B} \mathcal{B} \rangle = \frac{\alpha_2}{\ell^2}, \\ \langle Q_\alpha^i Q_\beta^j \rangle &= \alpha_1 C_{\alpha\beta} \delta^{ij}, \quad \langle Q_\alpha^i \Sigma_\beta^j \rangle = \alpha_2 C_{\alpha\beta} \delta^{ij}, \\ \langle \Sigma_\alpha^i \Sigma_\beta^j \rangle &= \frac{\alpha_1}{\ell^2} C_{\alpha\beta} \delta^{ij}. \end{aligned} \quad (3.4)$$

One can notice that the flat limit $\ell \rightarrow \infty$ leads to the $\mathcal{N} = 2$ Maxwell supergravity theory [60] in which the internal symmetry generator \mathcal{B} becomes a central charge.

A supersymmetric extension of the EEB algebra can be obtained by expanding the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra (3.1) and (3.2). Indeed, as we shall see, the S -expansion of a relativistic superalgebra considering $S_E^{(2)}$ as the relevant semigroup reproduces a NR contraction. This is due to the fact that the S -expansion method can be seen as a generalization of the Inönü–Wigner contraction process [78], when the semigroup under consideration belongs to the $S_E^{(N)}$ family. In particular, it would be interesting to analyze which kind of superalgebras can be found by considering $S_E^{(2)}$ for $N > 2$. Here, we shall focus on the semigroup $S_E^{(2)}$, which allows us to obtain not only a well-defined EEB superalgebra but also its non-degenerate invariant tensor.

Let $S_E^{(2)} = \{\lambda_0, \lambda_1, \lambda_2, \lambda_3\}$ be the relevant semigroup whose elements satisfy the following multiplication law:

$$\begin{array}{c|cccc} \lambda_3 & \lambda_3 & \lambda_3 & \lambda_3 & \lambda_3 \\ \lambda_2 & \lambda_2 & \lambda_3 & \lambda_3 & \lambda_3 \\ \lambda_1 & \lambda_1 & \lambda_2 & \lambda_3 & \lambda_3 \\ \lambda_0 & \lambda_0 & \lambda_1 & \lambda_2 & \lambda_3 \\ \hline & \lambda_0 & \lambda_1 & \lambda_2 & \lambda_3 \end{array} \quad (3.5)$$

Here $\lambda_3 = 0_s$ is the zero element of the semigroup such that $0_s \lambda_k = 0_s$. Before applying the S -expansion to the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra we will consider a particular subspace decomposition of the Lie superalgebra. Let $V_0 = \{J_0, P_0, Z_0, \mathcal{T}, \mathcal{U}, \mathcal{B}, Q_\alpha^+, \Sigma_\alpha^+\}$ and $V_1 = \{J_a, P_a, Z_a, Q_\alpha^-, \Sigma_\alpha^-\}$ be the subspaces decomposition of the $\mathcal{N} = 2$ superalgebra (3.1) and (3.2), where we have split the Lorentz index as $A = \{0, a\}$ with $a = 1, 2$ and where we have defined

$$Q_\alpha^\pm = \frac{1}{\sqrt{2}} (Q_\alpha^1 \pm \epsilon_{\alpha\beta} Q_\beta^2), \quad \Sigma_\alpha^\pm = \frac{1}{\sqrt{2}} (\Sigma_\alpha^1 \pm \epsilon_{\alpha\beta} \Sigma_\beta^2). \quad (3.6)$$

One can see that such decomposition satisfies

$$[V_0, V_0] \subset V_0, \quad [V_0, V_1] \subset V_1, \quad [V_1, V_1] \subset V_0. \quad (3.7)$$

Let us consider now $S_E^{(2)} = S_0 \cup S_1$ as the semigroup decomposition where

$$\begin{aligned} S_0 &= \{\lambda_0, \lambda_2, \lambda_3\}, \\ S_1 &= \{\lambda_1, \lambda_3\}. \end{aligned} \quad (3.8)$$

Then, the decomposition (3.8) is said to be resonant since it satisfies the same structure than the subspaces, that is

$$S_0 \cdot S_0 \subset S_0, \quad S_0 \cdot S_1 \subset S_1, \quad S_1 \cdot S_1 \subset S_0. \quad (3.9)$$

Following the definitions of [64], after extracting a resonant subalgebra of the $S_E^{(2)}$ -expansion of the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra and applying a 0_s -reduction, one finds a novel NR expanded superalgebra spanned by the set of generators:

$$\begin{aligned} \{\tilde{J}, \tilde{G}_a, \tilde{S}, \tilde{H}, \tilde{P}_a, \tilde{M}, \tilde{Z}, \tilde{Z}_a, \tilde{T}, \tilde{Y}_1, \tilde{Y}_2, \tilde{U}_1, \tilde{U}_2, \tilde{B}_1, \tilde{B}_2, \\ \tilde{Q}_\alpha^+, \tilde{Q}_\alpha^-, \tilde{R}_\alpha, \tilde{\Sigma}_\alpha^+, \tilde{\Sigma}_\alpha^-, \tilde{W}_\alpha\}. \end{aligned} \quad (3.10)$$

The NR generators are related to the relativistic ones through the semigroup elements as

$$\begin{aligned} \tilde{J} &= \lambda_0 J_0, & \tilde{S} &= \lambda_2 J_0, & \tilde{G}_a &= \lambda_1 J_a, \\ \tilde{H} &= \lambda_0 P_0, & \tilde{M} &= \lambda_2 P_0, & \tilde{P}_a &= \lambda_1 P_a, \\ \tilde{Z} &= \lambda_0 Z_0, & \tilde{T} &= \lambda_2 Z_0, & \tilde{Z}_a &= \lambda_1 Z_a, \\ \tilde{Q}_\alpha^+ &= \lambda_0 Q_\alpha^+, & \tilde{R}_\alpha &= \lambda_2 Q_\alpha^+, & \tilde{Q}_\alpha^- &= \lambda_1 Q_\alpha^-, \\ \tilde{\Sigma}_\alpha^+ &= \lambda_0 \Sigma_\alpha^+, & \tilde{W}_\alpha &= \lambda_2 \Sigma_\alpha^+, & \tilde{\Sigma}_\alpha^- &= \lambda_1 \Sigma_\alpha^-, \\ \tilde{Y}_1 &= \lambda_0 \mathcal{T}, & \tilde{U}_1 &= \lambda_0 \mathcal{U}, & \tilde{B}_1 &= \lambda_0 \mathcal{B}, \\ \tilde{Y}_2 &= \lambda_2 \mathcal{T}, & \tilde{U}_2 &= \lambda_2 \mathcal{U}, & \tilde{B}_2 &= \lambda_2 \mathcal{B}. \end{aligned} \quad (3.11)$$

The NR generators satisfy the bosonic subalgebra (2.1) along with the following commutation relations:

$$\begin{aligned} [\tilde{J}, \tilde{Q}_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{Q}_\beta^\pm, & [\tilde{J}, \tilde{R}_\alpha] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{H}, \tilde{Q}_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{\Sigma}_\beta^\pm, \\ [\tilde{J}, \tilde{\Sigma}_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{\Sigma}_\beta^\pm, & [\tilde{J}, \tilde{W}_\alpha] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{W}_\beta, \\ [\tilde{H}, \tilde{\Sigma}_\alpha^\pm] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{Q}_\beta^\pm, \\ [\tilde{Z}, \tilde{Q}_\alpha^\pm] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{Q}_\beta^\pm, & [\tilde{H}, \tilde{R}_\alpha] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{W}_\beta, \\ [\tilde{Z}, \tilde{\Sigma}_\alpha^\pm] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{\Sigma}_\beta^\pm, \\ [\tilde{S}, \tilde{Q}_\alpha^+] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, & [\tilde{M}, \tilde{Q}_\alpha^+] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{W}_\beta, \\ [\tilde{T}, \tilde{Q}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{S}, \tilde{\Sigma}_\alpha^+] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{W}_\beta, & [\tilde{M}, \tilde{\Sigma}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{T}, \tilde{\Sigma}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{W}_\beta, \end{aligned}$$

$$\begin{aligned}
[\tilde{H}, \tilde{W}_\alpha] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, & [\tilde{Z}, \tilde{R}_\alpha] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\
[\tilde{Z}, \tilde{W}_\alpha] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{W}_\beta, \\
[\tilde{G}_a, \tilde{Q}_\alpha^+] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{Q}_\beta^-, & [\tilde{G}_a, \tilde{Q}_\alpha^-] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{R}_\beta, \\
[\tilde{P}_a, \tilde{Q}_\alpha^+] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{\Sigma}_\beta^-, \\
[\tilde{G}_a, \tilde{\Sigma}_\alpha^+] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{\Sigma}_\beta^-, & [\tilde{G}_a, \tilde{\Sigma}_\alpha^-] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{W}_\beta, \\
[\tilde{P}_a, \tilde{Q}_\alpha^-] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{W}_\beta^-, \\
[\tilde{P}_a, \tilde{\Sigma}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{Q}_\beta^-, & [\tilde{Z}_a, \tilde{Q}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{Q}_\beta^-, \\
[\tilde{Z}_a, \tilde{Q}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{R}_\beta, \\
[\tilde{P}_a, \tilde{\Sigma}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{R}_\beta, & [\tilde{Z}_a, \tilde{\Sigma}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{\Sigma}_\beta^-, \\
[\tilde{Z}_a, \tilde{\Sigma}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{W}_\beta, & & (3.12)
\end{aligned}$$

while the additional bosonic generators $\{Y_1, Y_2, U_1, U_2, B_1, B_2\}$, which are expansions of $\mathfrak{so}(2)$ internal symmetry generators, satisfy the following commutators:

$$\begin{aligned}
[\tilde{Y}_1, \tilde{Q}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^+, & [\tilde{Y}_1, \tilde{Q}_\alpha^-] &= -\frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^-, \\
[\tilde{Y}_1, \tilde{R}_\alpha] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, \\
[\tilde{Y}_1, \tilde{\Sigma}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{\Sigma}_\beta^+, & [\tilde{Y}_1, \tilde{\Sigma}_\alpha^-] &= -\frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{\Sigma}_\beta^-, \\
[\tilde{Y}_1, \tilde{W}_\alpha] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{W}_\beta, \\
[\tilde{Y}_2, \tilde{Q}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, & [\tilde{U}_1, \tilde{Q}_\alpha^-] &= -\frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{\Sigma}_\beta^-, \\
[\tilde{U}_1, \tilde{Q}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{\Sigma}_\beta^+, \\
[\tilde{Y}_2, \tilde{\Sigma}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{W}_\beta, & [\tilde{U}_1, \tilde{R}_\alpha] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{W}_\beta, \\
[\tilde{U}_2, \tilde{Q}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{W}_\beta, \\
[\tilde{U}_1, \tilde{\Sigma}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^+, & [\tilde{U}_1, \tilde{\Sigma}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^-, \\
[\tilde{B}_1, \tilde{R}_\alpha] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, \\
[\tilde{B}_1, \tilde{Q}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^+, & [\tilde{B}_1, \tilde{Q}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^-, \\
[\tilde{B}_1, \tilde{W}_\alpha] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{W}_\beta, \\
[\tilde{B}_1, \tilde{\Sigma}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{\Sigma}_\beta^+, & [\tilde{B}_1, \tilde{\Sigma}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{\Sigma}_\beta^-, \\
[\tilde{U}_1, \tilde{W}_\alpha] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, \\
[\tilde{U}_2, \tilde{\Sigma}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, & [\tilde{B}_2, \tilde{Q}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, \\
[\tilde{B}_2, \tilde{\Sigma}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{W}_\beta. & & (3.13)
\end{aligned}$$

On the other hand, the fermionic generators satisfy the following anti-commutation relations:

$$\begin{aligned}
\{\tilde{Q}_\alpha^-, \tilde{Q}_\beta^-\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{M} + (\gamma^0 C)_{\alpha\beta} \tilde{U}_2, \\
\{\tilde{Q}_\alpha^+, \tilde{Q}_\beta^+\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{H} - (\gamma^0 C)_{\alpha\beta} \tilde{U}_1, \\
\{\tilde{Q}_\alpha^-, \tilde{\Sigma}_\beta^-\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{T} + (\gamma^0 C)_{\alpha\beta} \tilde{B}_2, \\
\{\tilde{Q}_\alpha^+, \tilde{\Sigma}_\beta^+\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{Z} - (\gamma^0 C)_{\alpha\beta} \tilde{B}_1, \\
\{\tilde{Q}_\alpha^+, \tilde{R}_\beta\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{M} - (\gamma^0 C)_{\alpha\beta} \tilde{U}_2, \\
\{\tilde{Q}_\alpha^+, \tilde{W}_\beta\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{T} - (\gamma^0 C)_{\alpha\beta} \tilde{B}_2, \\
\{\tilde{\Sigma}_\alpha^+, \tilde{R}_\beta\} &= -(\gamma^0 C)_{\alpha\beta} \tilde{T} - (\gamma^0 C)_{\alpha\beta} \tilde{B}_2, \\
\{\tilde{\Sigma}_\alpha^-, \tilde{\Sigma}_\beta^-\} &= -(\gamma^0 C)_{\alpha\beta} \frac{\tilde{M}}{\ell^2} + (\gamma^0 C)_{\alpha\beta} \frac{\tilde{U}_2}{\ell^2}, \\
\{\tilde{\Sigma}_\alpha^+, \tilde{\Sigma}_\beta^+\} &= -(\gamma^0 C)_{\alpha\beta} \frac{\tilde{H}}{\ell^2} - (\gamma^0 C)_{\alpha\beta} \frac{\tilde{U}_1}{\ell^2}, \\
\{\tilde{\Sigma}_\alpha^+, \tilde{W}_\beta\} &= -(\gamma^0 C)_{\alpha\beta} \frac{\tilde{M}}{\ell^2} - (\gamma^0 C)_{\alpha\beta} \frac{\tilde{U}_2}{\ell^2}, \\
\{\tilde{Q}_\alpha^+, \tilde{Q}_\beta^-\} &= -(\gamma^a C)_{\alpha\beta} \tilde{P}_a, \\
\{\tilde{Q}_\alpha^\pm, \tilde{\Sigma}_\beta^\mp\} &= -(\gamma^a C)_{\alpha\beta} \tilde{Z}_a, \\
\{\tilde{\Sigma}_\alpha^+, \tilde{\Sigma}_\beta^-\} &= -\frac{1}{\ell^2} (\gamma^a C)_{\alpha\beta} \tilde{P}_a. & & (3.14)
\end{aligned}$$

The superalgebra given by (2.1), (3.12), (3.13), and (3.14) will be denoted as the *enlarged Extended Bargmann* superalgebra and, as we can see, it properly contains the EEB algebra [53] as bosonic subalgebra. Let us note that the presence of the \tilde{R} and \tilde{W} generators is similar to what happens in the Extended Bargmann superalgebra presented in [3] and the Extended Newtonian superalgebra of [5] in which a \tilde{R} generator is considered. One can see that the supersymmetric extension of the EEB algebra requires the presence of six additional bosonic generators $\tilde{Y}_1, \tilde{Y}_2, \tilde{U}_1, \tilde{U}_2, \tilde{B}_1$, and \tilde{B}_2 which act non-trivially on the fermionic charges $\tilde{Q}_\alpha^\pm, \tilde{\Sigma}_\alpha^\pm, \tilde{R}_\alpha$, and \tilde{W}_α . Interestingly, both \tilde{B}_1 and \tilde{B}_2 become central in the vanishing cosmological constant limit. Indeed, the flat limit $\ell \rightarrow \infty$ of the EEB superalgebra reproduces the Maxwellian version of the extended Bargmann superalgebra introduced in [6], and corresponds to the supersymmetric extension of the MEB algebra first presented in [46]. Let us stress that the EEB superalgebra obtained here has been obtained through an $S_E^{(2)}$ -expansion of a relativistic superalgebra. As we shall see in the next section, the supersymmetric extension of the EEB algebra presented here allowing the construction a well-defined CS supergravity action is not unique. Indeed, a different supersymmetric extension of the

AdS- \mathcal{L} algebra can be expanded with the same semigroup $S_E^{(2)}$ to get a different EEB superalgebra.

Before studying the construction of a NR CS supergravity action invariant under the novel NR superalgebra, we will show that the present NR structure can be rewritten in a different way. Indeed, the Nappi–Witten symmetry [79, 80], which can be seen as a central extension of the homogeneous part of the Galilei algebra, appears considering an appropriate redefinition of the generators. In particular, the EEB superalgebra can be written as three copies of the Nappi–Witten algebra, two of which are augmented by supersymmetry, endowed with $u(1)$ generators. Indeed, the aforesaid structure is revealed by considering the following redefinition:

$$\begin{aligned} G_a &= \frac{1}{2} (\ell^2 \tilde{Z}_a + \ell \tilde{P}_a), & G_a^* &= \frac{1}{2} (\ell^2 \tilde{Z}_a - \ell \tilde{P}_a), \\ \hat{G}_a &= \tilde{G}_a - \ell^2 \tilde{Z}_a, \\ J &= \frac{1}{2} (\ell^2 \tilde{Z} + \ell \tilde{H}), & J^* &= \frac{1}{2} (\ell^2 \tilde{Z} - \ell \tilde{H}), \\ \hat{J} &= \tilde{J} - \ell^2 \tilde{Z}, \\ S &= \frac{1}{2} (\ell^2 \tilde{T} + \ell \tilde{M}), & S^* &= \frac{1}{2} (\ell^2 \tilde{T} - \ell \tilde{M}), \\ \hat{S} &= \tilde{S} - \ell^2 \tilde{T}, \\ T_1 &= \frac{1}{2} (\ell^2 \tilde{B}_1 + \ell \tilde{U}_1), & T_1^* &= \frac{1}{2} (\ell^2 \tilde{B}_1 - \ell \tilde{U}_1), \\ \hat{T}_1 &= \tilde{T}_1 - \ell^2 \tilde{B}_1, \\ T_2 &= \frac{1}{2} (\ell^2 \tilde{B}_2 + \ell \tilde{U}_2), & T_2^* &= \frac{1}{2} (\ell^2 \tilde{B}_2 - \ell \tilde{U}_2), \\ \hat{T}_2 &= \tilde{T}_2 - \ell^2 \tilde{B}_2, \\ Q_\alpha^+ &= \frac{1}{2} (\ell^{1/2} \tilde{Q}_\alpha^+ + \ell^{3/2} \tilde{\Sigma}_\alpha^+), & Q_\alpha^- &= \frac{1}{2} (\ell^{1/2} \tilde{Q}_\alpha^- + \ell^{3/2} \tilde{\Sigma}_\alpha^-), \\ \mathcal{R}_\alpha &= \frac{1}{2} (\ell^{1/2} \tilde{R}_\alpha + \ell^{3/2} \tilde{W}_\alpha), \\ \mathcal{G}_\alpha^+ &= \frac{i}{2} (\ell^{1/2} \tilde{Q}_\alpha^+ - \ell^{3/2} \tilde{\Sigma}_\alpha^+), & \mathcal{G}_\alpha^- &= \frac{i}{2} (\ell^{1/2} \tilde{Q}_\alpha^- - \ell^{3/2} \tilde{\Sigma}_\alpha^-), \\ \mathcal{W}_\alpha &= \frac{i}{2} (\ell^{1/2} \tilde{R}_\alpha - \ell^{3/2} \tilde{W}_\alpha). \end{aligned} \quad (3.15)$$

One can notice that set of generators $\{\hat{G}_a, \hat{J}, \hat{S}, \hat{T}_1, \hat{T}_2\}$ satisfies the Nappi–Witten algebra [79] endowed with $u(1)$ generators \hat{T}_1 and \hat{T}_2 ,

$$\begin{aligned} [\hat{J}, \hat{G}_a] &= \epsilon_{ab} \hat{G}_b, \\ [\hat{G}_a, \hat{G}_b] &= -\epsilon_{ab} \hat{S}. \end{aligned} \quad (3.16)$$

On the other hand, the set of generators $\{G_a, J, S, T_1, T_2, Q_\alpha^+, Q_\alpha^-, \mathcal{R}_\alpha\}$ satisfies a novel supersymmetric extension of the Nappi–Witten algebra,

$$\begin{aligned} [J, G_a] &= \epsilon_{ab} G_b, & [G_a, G_b] &= -\epsilon_{ab} S, \\ [J, Q_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta Q_\beta^\pm, & [J, \mathcal{R}_\alpha] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \mathcal{R}_\beta, \\ [G_a, Q_\alpha^+] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta Q_\beta^-, & [G_a, Q_\alpha^-] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \mathcal{R}_\beta, \end{aligned}$$

$$\begin{aligned} [S, Q_\alpha^+] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \mathcal{R}_\beta, & [T_1, Q_\alpha^\pm] &= \pm \frac{1}{2} (\gamma_0)_\alpha^\beta Q_\beta^\pm, \\ [T_2, Q_\alpha^+] &= \frac{1}{2} (\gamma_0)_\alpha^\beta \mathcal{R}_\beta, & [T_1, \mathcal{R}_\alpha] &= \frac{1}{2} (\gamma_0)_\alpha^\beta \mathcal{R}_\beta, \\ \{Q_\alpha^+, Q_\beta^-\} &= -(\gamma^a C)_{\alpha\beta} G_a, \\ \{Q_\alpha^+, Q_\beta^+\} &= -(\gamma^0 C)_{\alpha\beta} J - (\gamma^0 C)_{\alpha\beta} T_1, \\ \{Q_\alpha^-, Q_\beta^-\} &= -(\gamma^0 C)_{\alpha\beta} S + (\gamma^0 C)_{\alpha\beta} T_2, \\ \{Q_\alpha^+, \mathcal{R}_\beta\} &= -(\gamma^0 C)_{\alpha\beta} S - (\gamma^0 C)_{\alpha\beta} T_2. \end{aligned} \quad (3.17)$$

Furthermore, one can see that the generators $\{G_a^*, J^*, S^*, T_1^*, T_2^*, \mathcal{G}_\alpha^+, \mathcal{G}_\alpha^-, \mathcal{W}_\alpha\}$ satisfy an additional copy of a Nappi–Witten superalgebra (3.17). The present supersymmetric extension of the Nappi–Witten algebra contains additional bosonic generators $\{T_1, T_2\}$ which are required to assure the Jacobi identities. Furthermore, as in the case of the extended Bargmann superalgebra [3] and the extended Newtonian superalgebra [5], the Nappi–Witten superalgebra has three fermionic charges given by Q_α^+ , Q_α^- , and \mathcal{R}_α . Such feature could be useful to establish some Lie algebra expansion between the Nappi–Witten superalgebra (3.17) and known NR superalgebras, similarly as in the bosonic case [36, 67]. Interestingly, the EEB superalgebra inherits the same structure than its relativistic version, in which case the AdS- \mathcal{L} superalgebra can be rewritten as the direct sum of the Lorentz algebra and two super-Lorentz.

Let us note that the Nappi–Witten algebra (3.16) endowed with $u(1)$ generators along with the Nappi–Witten superalgebra (3.17) can be rewritten as an extended Newton–Hooke superalgebra [5]. Indeed, such structure can be obtained by considering the following redefinition of the generators of (3.16) and (3.17):

$$\begin{aligned} \check{G}_a &= G_a - \hat{G}_a, \check{P}_a = \frac{1}{\ell} (G_a + \hat{G}_a), \check{Q}_\alpha^+ = \sqrt{\frac{2}{\ell}} Q_\alpha^+, \\ \check{J} &= J + \hat{J}, \check{H} = \frac{1}{\ell} (J - \hat{J}), \check{Q}_\alpha^- = \sqrt{\frac{2}{\ell}} Q_\alpha^-, \\ \check{S} &= S + \hat{S}, \check{M} = \frac{1}{\ell} (S - \hat{S}), \check{R}_\alpha = \sqrt{\frac{2}{\ell}} \mathcal{R}_\alpha, \\ \check{T}_1 &= T_1 - \hat{T}_1, \check{U}_1 = \frac{1}{\ell} (T_1 + \hat{T}_1), \\ \check{T}_2 &= T_2 - \hat{T}_2, \check{U}_2 = \frac{1}{\ell} (T_2 + \hat{T}_2). \end{aligned} \quad (3.18)$$

One can see that the set of generators $\{\check{G}_a, \check{P}_a, \check{J}, \check{H}, \check{S}, \check{M}, \check{T}_1, \check{T}_2, \check{U}_1, \check{U}_2, \check{Q}_\alpha^+, \check{Q}_\alpha^-, \check{R}_\alpha\}$ satisfies an extended Newton–Hooke superalgebra,

$$\begin{aligned} [\check{J}, \check{G}_a] &= \epsilon_{ab} \check{G}_b, & [\check{G}_a, \check{G}_b] &= -\epsilon_{ab} \check{S}, & [\check{J}, \check{P}_a] &= \epsilon_{ab} \check{P}_b, \\ [\check{H}, \check{G}_a] &= \epsilon_{ab} \check{P}_b, & [\check{G}_a, \check{P}_b] &= -\epsilon_{ab} \check{M}, & [\check{H}, \check{P}_a] &= \frac{1}{\ell^2} \epsilon_{ab} \check{G}_b, \\ [\check{P}_a, \check{P}_b] &= -\frac{1}{\ell^2} \epsilon_{ab} \check{S}, & [\check{J}, \check{Q}_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \check{Q}_\beta^\pm, \end{aligned}$$

$$\begin{aligned}
[\check{H}, \check{Q}_\alpha^\pm] &= -\frac{1}{2\ell} (\gamma_0)_\alpha^\beta \check{Q}_\beta^\pm, \\
[\check{J}, \check{R}_\alpha] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \check{R}_\beta, \quad [\check{H}, \check{R}_\alpha] = -\frac{1}{2\ell} (\gamma_0)_\alpha^\beta \check{R}_\beta, \\
[\check{G}_a, \check{Q}_\alpha^\pm] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \check{Q}_\beta^\pm, \\
[\check{P}_a, \check{Q}_\alpha^\pm] &= -\frac{1}{2\ell} (\gamma_a)_\alpha^\beta \check{Q}_\beta^\pm, \quad [\check{G}_a, \check{Q}_\alpha^\pm] = -\frac{1}{2} (\gamma_a)_\alpha^\beta \check{R}_\beta, \\
[\check{P}_a, \check{Q}_\alpha^\pm] &= -\frac{1}{2\ell} (\gamma_a)_\alpha^\beta \check{R}_\beta, \\
[\check{S}, \check{Q}_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \check{R}_\beta, \quad [\check{M}, \check{Q}_\alpha^\pm] = -\frac{1}{2\ell} (\gamma_0)_\alpha^\beta \check{R}_\beta, \\
[\check{T}_1, \check{Q}_\alpha^\pm] &= \pm \frac{1}{2} (\gamma_0)_{\alpha\beta} \check{Q}_\beta^\pm, \\
[\check{U}_1, \check{Q}_\alpha^\pm] &= \pm \frac{1}{2\ell} (\gamma_0)_{\alpha\beta} \check{Q}_\beta^\pm, \quad [\check{T}_2, \check{Q}_\alpha^\pm] = \frac{1}{2} (\gamma_0)_{\alpha\beta} \check{R}_\beta, \\
[\check{U}_2, \check{Q}_\alpha^\pm] &= \frac{1}{2\ell} (\gamma_0)_{\alpha\beta} \check{R}_\beta, \\
[\check{T}_1, \check{R}_\alpha] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \check{R}_\beta, \quad [\check{U}_1, \check{R}_\alpha] = \frac{1}{2\ell} (\gamma_0)_{\alpha\beta} \check{R}_\beta, \\
\{\check{Q}_\alpha^+, \check{Q}_\beta^-\} &= -\frac{1}{\ell} (\gamma^a C)_{\alpha\beta} \check{G}_a - (\gamma^a C)_{\alpha\beta} \check{P}_a, \\
\{\check{Q}_\alpha^+, \check{Q}_\beta^+\} &= -\frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \check{J} - (\gamma^0 C)_{\alpha\beta} \check{H} \\
&\quad - \frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \check{T}_1 - (\gamma^0 C)_{\alpha\beta} \check{U}_1, \\
\{\check{Q}_\alpha^-, \check{Q}_\beta^-\} &= -\frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \check{S} - (\gamma^0 C)_{\alpha\beta} \check{M} \\
&\quad + \frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \check{T}_2 + (\gamma^0 C)_{\alpha\beta} \check{U}_2, \\
\{\check{Q}_\alpha^+, \check{R}_\beta\} &= -\frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \check{S} - (\gamma^0 C)_{\alpha\beta} \check{M} \\
&\quad - \frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \check{T}_2 - (\gamma^0 C)_{\alpha\beta} \check{U}_2,
\end{aligned}$$

which involves two more generators (\check{T}_1 and \check{T}_2 , which act non-trivially on the fermionic generators \check{Q}^\pm and \check{R}) with respect to the extended Newton–Hooke superalgebra of [5]. As we can see, in (3.19) both \check{U}_1 and \check{U}_2 act non-trivially on the fermionic generators \check{Q}^\pm and \check{R} in the presence of the length parameter ℓ . In particular, both of these NR generators become central in the flat limit $\ell \rightarrow \infty$ which reproduces a central extension of the extended Bargmann superalgebra [3] endowed with two additional generators \check{T}_1 and \check{T}_2 . Let us note that the extended Newton–Hooke superalgebra (3.19) can be seen as the NR counterpart of the $\mathcal{N} = 2$ AdS superalgebra [81],

$$\begin{aligned}
[J_A, J_B] &= \epsilon_{ABC} J^C, & [J_A, P_B] &= \epsilon_{ABC} P^C, \\
[P_A, P_B] &= \frac{1}{\ell^2} \epsilon_{ABC} J^C, & [J_A, Q_\alpha^i] &= -\frac{1}{2} (\gamma_A)_\alpha^\beta Q_\beta^i, \\
[P_A, Q_\alpha^i] &= -\frac{1}{2\ell} (\gamma_A)_\alpha^\beta Q_\beta^i, & [T, Q_\alpha^i] &= \frac{1}{2} \epsilon^{ij} Q_\alpha^j
\end{aligned}$$

$$\begin{aligned}
[U, Q_\alpha^i] &= \frac{1}{2\ell^2} \epsilon^{ij} Q_\alpha^j, \\
\{Q_\alpha^i, Q_\beta^j\} &= -\frac{\delta_{ij}}{\ell} (\gamma^A C)_{\alpha\beta} J_A - \delta_{ij} (\gamma^A C)_{\alpha\beta} P_A \\
&\quad - C_{\alpha\beta} \epsilon^{ij} \left(\frac{1}{\ell} T + U \right).
\end{aligned} \quad (3.20)$$

Indeed, the extended Newton–Hooke superalgebra (3.19) can alternatively be obtained after applying a resonant $S_E^{(2)}$ -expansion of the $\mathcal{N} = 2$ AdS superalgebra and performing a 0_s -reduction.

Thus, the EEB superalgebra can alternatively be rewritten as the direct sum of the extended Newton–Hooke superalgebra (3.19) and the super Nappi–Witten one (3.17). Such feature is inherited from its relativistic version, in which case the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra can be written as the direct sum of the super AdS algebra and the super-Lorentz one.

3.2 Non-relativistic Chern–Simons supergravity action

We now construct a NR CS supergravity action based on the EEB superalgebra given by (2.1), (3.12), (3.13), and (3.14). Although the superalgebra seems simpler written as copies of the Nappi–Witten (super)algebra, the motivation to consider the EEB structure given by (2.1), (3.12), (3.13), and (3.14) is twofold. First, as we shall see, it directly offers us an alternative NR supergravity theory in presence of a cosmological constant. Second, it allows to establish a flat limit leading in a manifest way to the Maxwellian version of the extended Bargmann supergravity [6].

As we have already mentioned, a crucial ingredient to construct a CS action is the invariant tensor. Interestingly, the same semigroup allowing to obtain the new NR superalgebra (3.19) can be used to find the non-vanishing components of the EEB invariant tensor (for further details see Appendix A). Indeed, the invariant tensor for the EEB superalgebra can be obtained in terms of the invariant tensor for the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra given by (3.4). One can then show that the non-vanishing components of the invariant tensor for the EEB superalgebra are given by (2.4) along with

$$\begin{aligned}
\langle \check{Y}_1 \check{Y}_2 \rangle &= \tilde{\alpha}_0, & \langle \check{Y}_1 \check{U}_2 \rangle &= \tilde{\alpha}_1, \\
\langle \check{U}_1 \check{Y}_2 \rangle &= \tilde{\alpha}_1, \\
\langle \check{Y}_1 \check{B}_2 \rangle &= \tilde{\alpha}_2, & \langle \check{U}_1 \check{U}_2 \rangle &= \tilde{\alpha}_2, \\
\langle \check{B}_1 \check{Y}_2 \rangle &= \tilde{\alpha}_2, \\
\langle \check{U}_1 \check{B}_2 \rangle &= \frac{\tilde{\alpha}_1}{\ell^2}, & \langle \check{B}_1 \check{U}_2 \rangle &= \frac{\tilde{\alpha}_1}{\ell^2}, \\
\langle \check{B}_1 \check{B}_2 \rangle &= \frac{\tilde{\alpha}_2}{\ell^2}, \\
\langle \check{Q}_\alpha^- \check{Q}_\beta^- \rangle &= 2\tilde{\alpha}_1 C_{\alpha\beta}, & \langle \check{Q}_\alpha^+ \check{R}_\beta \rangle &= 2\tilde{\alpha}_1 C_{\alpha\beta},
\end{aligned}$$

$$\begin{aligned}\langle \tilde{Q}_\alpha^- \tilde{\Sigma}_\beta^- \rangle &= 2\tilde{\alpha}_2 C_{\alpha\beta}, \\ \langle \tilde{\Sigma}_\alpha^+ \tilde{R}_\beta \rangle &= 2\tilde{\alpha}_2 C_{\alpha\beta}, \quad \langle \tilde{Q}_\alpha^+ \tilde{W}_\beta \rangle = 2\tilde{\alpha}_2 C_{\alpha\beta}, \\ \langle \tilde{\Sigma}_\alpha^+ \tilde{W}_\beta \rangle &= \frac{2\tilde{\alpha}_1}{\ell^2} C_{\alpha\beta}, \\ \langle \tilde{\Sigma}_\alpha^- \tilde{\Sigma}_\beta^- \rangle &= \frac{2\tilde{\alpha}_1}{\ell^2} C_{\alpha\beta},\end{aligned}\quad (3.21)$$

where the NR generators are related to the relativistic ones through the semigroup elements as in (3.11). Here, the NR constants $\tilde{\alpha}_i$ are related to the relativistic parameters appearing in (3.4) as

$$\tilde{\alpha}_0 = \lambda_2 \alpha_0, \quad \tilde{\alpha}_1 = \lambda_2 \alpha_1, \quad \tilde{\alpha}_2 = \lambda_2 \alpha_2. \quad (3.22)$$

On the other hand, the gauge connection one-form A for the EEB superalgebra reads

$$\begin{aligned}A &= \omega \tilde{J} + \omega^a \tilde{G}_a + \tau \tilde{H} + e^a \tilde{P}_a + k \tilde{Z} + k^a \tilde{Z}_a + m \tilde{M} \\ &\quad + s \tilde{S} + t \tilde{T} + y_1 \tilde{Y}_1 + y_2 \tilde{Y}_2 + b_1 \tilde{B}_1 \\ &\quad + b_2 \tilde{B}_2 + u_1 \tilde{U}_1 + u_2 \tilde{U}_2 + \bar{\psi}^+ \tilde{Q}^+ + \bar{\psi}^- \tilde{Q}^- \\ &\quad + \bar{\xi}^+ \tilde{\Sigma}^+ + \bar{\xi}^- \tilde{\Sigma}^- + \bar{\rho} \tilde{R} + \bar{\chi} \tilde{W}.\end{aligned}\quad (3.23)$$

The corresponding curvature two-form $F = dA + A \wedge A = dA + \frac{1}{2} [A, A]$ written in terms of the generators is given by

$$\begin{aligned}F &= F(\omega) \tilde{J} + F^a(\omega^b) \tilde{G}_a + F(\tau) \tilde{H} + F^a(e^b) \tilde{P}_a \\ &\quad + F(k) \tilde{Z} + F^a(k^b) \tilde{Z}_a + F(m) \tilde{M} \\ &\quad + F(s) \tilde{S} + F(t) \tilde{T} + F(y_1) \tilde{Y}_1 \\ &\quad + F(y_2) \tilde{Y}_2 + F(b_1) \tilde{B}_1 + F(b_2) \tilde{B}_2 + F(u_1) \tilde{U}_1 \\ &\quad + F(u_2) \tilde{U}_2 + \nabla \bar{\psi}^+ \tilde{Q}^+ + \nabla \bar{\psi}^- \tilde{Q}^- + \nabla \bar{\xi}^+ \tilde{\Sigma}^+ \\ &\quad + \nabla \bar{\xi}^- \tilde{\Sigma}^- + \nabla \bar{\rho} \tilde{R} + \nabla \bar{\chi} \tilde{W}.\end{aligned}\quad (3.24)$$

In particular, the bosonic curvature two-forms are given by

$$\begin{aligned}F(\omega) &= R(\omega), \\ F^a(\omega^b) &= R^a(\omega^b), \\ F(\tau) &= R(\tau) + \frac{1}{2} \bar{\psi}^+ \gamma^0 \psi^+ + \frac{1}{2\ell^2} \bar{\xi}^+ \gamma^0 \xi^+, \\ F^a(e^b) &= R^a(e^b) + \bar{\psi}^+ \gamma^a \psi^- + \frac{1}{\ell^2} \bar{\xi}^+ \gamma^a \xi^-, \\ F(k) &= R(k) + \bar{\psi}^+ \gamma^0 \xi^+, \\ F^a(k^b) &= R^a(k^b) + \bar{\psi}^+ \gamma^a \xi^- + \bar{\psi}^- \gamma^a \xi^+, \\ F(m) &= R(m) + \frac{1}{2} \bar{\psi}^- \gamma^0 \psi^- + \bar{\psi}^+ \gamma^0 \rho + \frac{1}{2\ell^2} \bar{\xi}^- \gamma^0 \xi^- \\ &\quad + \frac{1}{\ell^2} \bar{\xi}^+ \gamma^0 \chi, \\ F(s) &= R(s), \\ F(t) &= R(t) + \bar{\psi}^- \gamma^0 \xi^- + \bar{\psi}^+ \gamma^0 \chi + \bar{\xi}^+ \gamma^0 \rho,\end{aligned}\quad (3.25)$$

where $R(\omega)$, $R^a(\omega^b)$, $R(\tau)$, $R^a(e^b)$, $R(k)$, $R^a(k^b)$, $R(m)$, $R(s)$, and $R(t)$ are the EEB curvatures defined in (2.9), together with

$$\begin{aligned}F(y_1) &= dy_1, \\ F(y_2) &= dy_2, \\ F(b_1) &= db_1 + \bar{\psi}^+ \gamma^0 \xi^+, \\ F(b_2) &= db_2 - \bar{\psi}^- \gamma^0 \xi^- + \bar{\psi}^+ \gamma^0 \chi + \bar{\xi}^+ \gamma^0 \rho, \\ F(u_1) &= du_1 + \frac{1}{2} \bar{\psi}^+ \gamma^0 \psi^+ + \frac{1}{2\ell^2} \bar{\xi}^+ \gamma^0 \xi^+, \\ F(u_2) &= du_2 - \frac{1}{2} \bar{\psi}^- \gamma^0 \psi^- + \bar{\psi}^+ \gamma^0 \rho - \frac{1}{2\ell^2} \bar{\xi}^- \gamma^0 \xi^- \\ &\quad + \frac{1}{\ell^2} \bar{\xi}^+ \gamma^0 \chi.\end{aligned}\quad (3.26)$$

Let us observe that, if we restrict ourselves to the purely bosonic sector and excluding the extra bosonic gauge field $\{y_1, y_2, u_1, u_2, b_1, b_2\}$, we properly recover the two-form curvatures associated with the EEB algebra introduced in [36]. On the other hand, the covariant derivatives of the spinor 1-form fields read

$$\begin{aligned}\nabla \psi^+ &= d\psi^+ + \frac{1}{2} \omega \gamma_0 \psi^+ - \frac{1}{2} y_1 \gamma_0 \psi^+ + \frac{1}{2\ell^2} \tau \gamma_0 \xi^+ \\ &\quad + \frac{1}{2\ell^2} k \gamma_0 \psi^+ - \frac{1}{2\ell^2} u_1 \gamma_0 \xi^+ - \frac{1}{2\ell^2} b_1 \gamma_0 \psi^+, \\ \nabla \psi^- &= d\psi^- + \frac{1}{2} \omega \gamma_0 \psi^- + \frac{1}{2} \omega^a \gamma_a \psi^+ + \frac{1}{2} y_1 \gamma_0 \psi^- + \frac{1}{2\ell^2} \tau \gamma_0 \xi^- \\ &\quad + \frac{1}{2\ell^2} e^a \gamma_a \xi^+ + \frac{1}{2\ell^2} k^a \gamma_a \psi^+ \\ &\quad + \frac{1}{2\ell^2} k \gamma_0 \psi^- + \frac{1}{2\ell^2} u_1 \gamma_0 \xi^- + \frac{1}{2\ell^2} b_1 \gamma_0 \psi^-, \\ \nabla \xi^+ &= d\xi^+ + \frac{1}{2} \omega \gamma_0 \xi^+ + \frac{1}{2} \tau \gamma_0 \psi^+ - \frac{1}{2} y_1 \gamma_0 \xi^+ \\ &\quad - \frac{1}{2} u_1 \gamma_0 \psi^+ + \frac{1}{2\ell^2} k \gamma_0 \xi^+ - \frac{1}{2\ell^2} b_1 \gamma_0 \xi^+, \\ \nabla \xi^- &= d\xi^- + \frac{1}{2} \omega \gamma_0 \xi^- + \frac{1}{2} \tau \gamma_0 \psi^- + \frac{1}{2} e^a \gamma_a \psi^+ \\ &\quad + \frac{1}{2} \omega^a \gamma_a \xi^+ + \frac{1}{2} y_1 \gamma_0 \xi^- + \frac{1}{2} u_1 \gamma_0 \psi^- \\ &\quad + \frac{1}{2\ell^2} k^a \gamma_a \xi^+ + \frac{1}{2\ell^2} k \gamma_0 \xi^- + \frac{1}{2\ell^2} b_1 \gamma_0 \xi^-, \\ \nabla \rho &= d\rho + \frac{1}{2} \omega \gamma_0 \rho + \frac{1}{2} \omega^a \gamma_a \psi^- + \frac{1}{2} s \gamma_0 \psi^+ - \frac{1}{2} y_2 \gamma_0 \psi^+ \\ &\quad - \frac{1}{2} y_1 \gamma_0 \rho + \frac{1}{2\ell^2} e^a \gamma_a \xi^- \\ &\quad + \frac{1}{2\ell^2} k^a \gamma_a \psi^- + \frac{1}{2\ell^2} m \gamma_0 \xi^+ + \frac{1}{2\ell^2} \tau \gamma_0 \chi \\ &\quad + \frac{1}{2\ell^2} t \gamma_0 \psi^+ + \frac{1}{2\ell^2} k \gamma_0 \rho \\ &\quad - \frac{1}{2\ell^2} u_1 \gamma_0 \chi - \frac{1}{2\ell^2} u_2 \gamma_0 \xi^+ \\ &\quad - \frac{1}{2\ell^2} b_1 \gamma_0 \rho - \frac{1}{2\ell^2} b_2 \gamma_0 \psi^+, \\ \nabla \chi &= d\chi + \frac{1}{2} \omega \gamma_0 \chi + \frac{1}{2} \omega^a \gamma_a \xi^- + \frac{1}{2} e^a \gamma_a \psi^- \\ &\quad + \frac{1}{2} \tau \gamma_0 \rho + \frac{1}{2} s \gamma_0 \xi^+ + \frac{1}{2} m \gamma_0 \psi^+\end{aligned}$$

$$\begin{aligned}
& -\frac{1}{2}y_2\gamma_0\xi^+ - \frac{1}{2}y_1\gamma_0\chi - \frac{1}{2}u_2\gamma_0\psi^+ \\
& -\frac{1}{2}u_1\gamma_0\rho + \frac{1}{2\ell^2}k^a\gamma_a\xi^- + \frac{1}{2\ell^2}t\gamma_0\xi^+ + \frac{1}{2\ell^2}k\gamma_0\chi \\
& -\frac{1}{2\ell^2}b_1\gamma_0\chi - \frac{1}{2\ell^2}b_2\gamma_0\xi^+. \quad (3.27)
\end{aligned}$$

Notice that taking the $\ell \rightarrow \infty$ limit of the curvature two-forms (3.25), (3.26), and (3.27) one recovers the curvatures associated with the MEB superalgebra of [6].

A CS supergravity action based on the EEB superalgebra given by (2.1), (3.12), (3.13), and (3.14) can be constructed by combining the non-zero components of the invariant tensor (2.4) and (3.21) with the gauge connection one-form A (3.23). The NR CS supergravity action reads, up to boundary terms, as follows:

$$I_{\text{NR}} = \tilde{\alpha}_0 I_0 + \tilde{\alpha}_1 I_1 + \tilde{\alpha}_2 I_2, \quad (3.28)$$

where

$$\begin{aligned}
I_0 &= \int \omega_a R^a(\omega^b) - 2sR(\omega) + 2y_1 dy_2, \\
I_1 &= \int e_a R^a(\omega^b) + \omega_a R^a(e^b) - 2mR(\omega) - 2\tau ds \\
&\quad + \frac{1}{\ell^2}e_a R^a(k^b) + \frac{1}{\ell^2}k_a R^a(e^b) \\
&\quad - \frac{2}{\ell^2}\tau dt - \frac{2}{\ell^2}mR(k) + 2y_1 du_2 + 2u_1 dy_2 + \frac{2}{\ell^2}u_1 db_2 \\
&\quad + \frac{2}{\ell^2}b_1 du_2 - 2\bar{\psi}^+ \nabla \rho \\
&\quad - 2\bar{\rho} \nabla \psi^+ - 2\bar{\psi}^- \nabla \psi^- - \frac{2}{\ell^2}\bar{\xi}^+ \nabla \chi \\
&\quad - \frac{2}{\ell^2}\bar{\chi} \nabla \xi^+ - \frac{2}{\ell^2}\bar{\xi}^- \nabla \xi^-, \\
I_2 &= \int e_a R^a(e^b) + k_a R^a(\omega^b) + \omega_a R^a(k^b) \\
&\quad + \frac{1}{\ell^2}k_a R^a(k^b) - 2sR(k) - 2mR(\tau) \\
&\quad - 2tR(\omega) - \frac{2}{\ell^2}tR(k) + 2y_1 db_2 + 2u_1 du_2 + 2y_2 db_1 \\
&\quad + \frac{2}{\ell^2}b_1 db_2 - 2\bar{\psi}^- \nabla \xi^- \\
&\quad - 2\bar{\xi}^- \nabla \psi^- - 2\bar{\psi}^+ \nabla \chi - 2\bar{\chi} \nabla \psi^+ - 2\bar{\xi}^+ \nabla \rho - 2\bar{\rho} \nabla \xi^+. \quad (3.29)
\end{aligned}$$

The CS action (3.28) obtained here describes the so-called Enlarged Extended Bargmann supergravity theory. Let us note that the NR CS supergravity action (3.28) contains three independent sectors proportional to $\tilde{\alpha}_0$, $\tilde{\alpha}_1$, and $\tilde{\alpha}_2$. In particular, I_0 corresponds to the CS action for the NR exotic gravity coupled to the extra gauge fields y_1 and y_2 . The CS actions I_1 and I_2 describe the EEB CS supergravity in presence of the cosmological constant and the gauge field k_a . We observe that taking the flat limit $\ell \rightarrow \infty$ of (3.28) we

recover the CS MEB supergravity action of [6].³ In particular, the vanishing cosmological constant limit applied to the CS action I_1 reproduces the extended Bargmann supergravity CS action coupled to the additional gauge fields y_1 , y_2 , u_1 , and u_2 . The new NR CS supergravity action (3.28) generalizes the extended Newton–Hooke supergravity theory [5] in the sense that it not only reproduces the extended Bargmann supergravity in the flat limit, but also the Maxwellian version proportional to I_2 .

It is important to emphasize that in order to obtain the CS supergravity action through the S -expansion method we have restricted ourselves exclusively to the CS terms related to the λ_2 element of the semigroup $S_E^{(2)}$. Indeed, all the NR parameter $\tilde{\alpha}_0$, $\tilde{\alpha}_1$, and $\tilde{\alpha}_2$ are defined as in (3.22) in terms of the relativistic parameters and the λ_2 element. On the other hand, it is possible to obtain diverse exotic-like contributions to the CS action which are proportional to the λ_0 element. Nevertheless, since we are interested in the supersymmetric extension of the EEB gravity theory [36], we shall omit such exotic-like terms which would imply that the bosonic sector is no more EEB gravity. It would be interesting to study the Physical implications of such additional contributions.

Let us note that the CS supergravity action (3.28) can alternatively be obtained by expanding directly the relativistic $\mathcal{N} = 2$ AdS- \mathcal{L} CS supergravity action [60],

$$\begin{aligned}
I_R &= \int \alpha_0 \left[\omega_A d\omega^A + \frac{1}{3}\epsilon_{ABC}\omega^A\omega^B\omega^C + \tau dt \right] \\
&\quad + \alpha_1 \left[2e_A R^A + \frac{1}{3\ell^2}\epsilon_{ABC}e^A e^B e^C + \frac{2}{\ell^2}e_A F^A \right. \\
&\quad \left. - \bar{\psi}^i \nabla \hat{\psi}^i - \frac{1}{\ell^2}\bar{\xi}^i \nabla \hat{\xi}^i + u d\tau + u db \right] \\
&\quad + \alpha_2 \left[2\sigma_A R^A + e_A T^A + \frac{2}{\ell^2}e_A F^A + \frac{1}{\ell^2}\epsilon_{ABC}e^A \sigma^B e^C \right. \\
&\quad \left. - \bar{\psi}^i \nabla \hat{\xi}^i - \bar{\xi}^i \nabla \hat{\psi}^i + b d\tau + u du + \frac{1}{\ell^2}b db \right], \quad (3.30)
\end{aligned}$$

where

$$\begin{aligned}
R^A &= d\omega^A + \frac{1}{2}\epsilon^{ABC}\omega_B\omega_C, \\
T^A &= de^A + \epsilon^{ABC}\omega_B e_C, \\
F^A &= d\sigma^A + \epsilon^{ABC}\omega_B \sigma_C + \frac{1}{2\ell^2}\epsilon^{ABC}\sigma_B \sigma_C, \\
\nabla \hat{\psi}^i &= d\hat{\psi}^i + \frac{1}{2}\omega^A \gamma_A \hat{\psi}^i + \tau \epsilon^{ij} \hat{\psi}^j + \frac{1}{\ell^2}b \epsilon^{ij} \hat{\psi}^j \\
&\quad + \frac{1}{\ell^2}u \epsilon^{ij} \hat{\xi}^j, \\
\nabla \hat{\xi}^i &= d\hat{\xi}^i + \frac{1}{2}\omega^A \gamma_A \hat{\xi}^i + \frac{1}{2}e^A \gamma_A \hat{\psi}^i + \tau \epsilon^{ij} \hat{\xi}^j + u \epsilon^{ij} \hat{\psi}^j \\
&\quad + \frac{1}{\ell^2}b \epsilon^{ij} \hat{\xi}^j. \quad (3.31)
\end{aligned}$$

³ This also fixes a sign misprint appearing in [6].

Indeed, we can express the NR gauge fields in terms of the relativistic ones and the semigroup elements as

$$\begin{aligned} \omega &= \lambda_0 \omega_0, & s &= \lambda_2 \omega_0, & \omega_a &= \lambda_1 \omega_a, \\ \tau &= \lambda_0 e_0, & m &= \lambda_2 e_0, & e_a &= \lambda_1 e_a, \\ k &= \lambda_0 \sigma_0, & t &= \lambda_2 \sigma_0, & k_a &= \lambda_1 \sigma_a, \\ \psi_\alpha^+ &= \lambda_0 \hat{\psi}_\alpha^+, & \rho_\alpha &= \lambda_2 \hat{\psi}_\alpha^+, & \psi_\alpha^- &= \lambda_1 \hat{\psi}_\alpha^-, \\ \xi_\alpha^+ &= \lambda_0 \hat{\xi}_\alpha^+, & \chi_\alpha &= \lambda_2 \hat{\xi}_\alpha^+, & \xi_\alpha^- &= \lambda_1 \hat{\xi}_\alpha^-, \\ y_1 &= \lambda_0 \mathfrak{t}, & u_1 &= \lambda_0 \mathfrak{u}, & b_1 &= \lambda_0 \mathfrak{b}, \\ y_2 &= \lambda_2 \mathfrak{t}, & u_2 &= \lambda_2 \mathfrak{u}, & b_2 &= \lambda_2 \mathfrak{b}, \end{aligned} \quad (3.32)$$

where we have split the Lorentz index as $A = \{0, a\}$ with $a = 1, 2$ and where we have defined

$$\hat{\psi}_\alpha^\pm = \frac{1}{\sqrt{2}} \left(\hat{\psi}_\alpha^1 \pm \epsilon_{\alpha\beta} \hat{\psi}_\beta^2 \right), \quad \hat{\xi}_\alpha^\pm = \frac{1}{\sqrt{2}} \left(\hat{\xi}_\alpha^1 \pm \epsilon_{\alpha\beta} \hat{\xi}_\beta^2 \right). \quad (3.33)$$

Then, the EEB supergravity action is obtained considering the expanded gauge fields (3.32), the expanded parameters (3.22), the multiplication law (3.5), and the O_s -reduction property.

As an ending remark let us note that, since the invariant tensor is non-degenerate, the field equations from the NR CS supergravity action (3.28) imply the vanishing of the curvature two-forms (3.25), (3.26), and (3.27) associated with the EEB superalgebra. The aforementioned curvatures transform covariantly with respect to the supersymmetry transformation laws given in Appendix B.

4 Three-dimensional non-standard enlarged extended Bargmann supergravity

Let us focus now on an alternative supersymmetric extension of the EEB algebra which we call as non-standard EEB superalgebra. The new structure appears as an S -expansion of a different $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra and, unlike the EEB superalgebra presented in the previous section, contains only three fermionic charges. The construction of a CS supergravity action based on the aforementioned superalgebra is also discussed.

4.1 Non-standard enlarged extended Bargmann superalgebra

An alternative supersymmetric extension of the EEB algebra can be defined by expanding a $\mathcal{N} = 2$ extension of the relativistic non-standard AdS- \mathcal{L} superalgebra [59, 82]. The aforesaid $\mathcal{N} = 2$ non-standard supersymmetric extension of the AdS- \mathcal{L} algebra is spanned by the set of generators

$\{J_A, P_a, Z_a, \mathcal{U}, \mathcal{T}, \mathcal{B}, Q_\alpha^i\}$ which satisfy the following (anti-)commutation relations:

$$\begin{aligned} [J_A, J_B] &= \epsilon_{ABC} J^C, & [J_A, P_B] &= \epsilon_{ABC} P^C, \\ [J_A, Z_B] &= \epsilon_{ABC} Z^C, & [P_A, P_B] &= \epsilon_{ABC} Z^C, \\ [Z_A, Z_B] &= \frac{1}{\ell^2} \epsilon_{ABC} Z^C, & [P_A, Z_B] &= \frac{1}{\ell^2} \epsilon_{ABC} P^C, \\ [J_A, Q_\alpha^i] &= -\frac{1}{2} (\gamma_A)_\alpha^\beta Q_\beta^i, & [P_A, Q_\alpha^i] &= -\frac{1}{2\ell} (\gamma_A)_\alpha^\beta Q_\beta^i, \\ [Z_A, Q_\alpha^i] &= -\frac{1}{2\ell^2} (\gamma_A)_\alpha^\beta Q_\beta^i, & [\mathcal{T}, Q_\alpha^i] &= \frac{1}{2} \epsilon^{ij} Q_\beta^j, \\ [\mathcal{U}, Q_\alpha^i] &= \frac{1}{2\ell^2} \epsilon^{ij} Q_\beta^j, \\ \{Q_\alpha^i, Q_\beta^j\} &= -\frac{\delta_{ij}}{\ell} (\gamma^A C)_{\alpha\beta} P_A - \delta_{ij} (\gamma^A C)_{\alpha\beta} Z_A \\ &\quad - C_{\alpha\beta} \epsilon^{ij} \left(\mathcal{U} + \frac{1}{\ell} \mathcal{B} \right), \end{aligned} \quad (4.1)$$

where, as usual, $i = 1, 2$ denotes the number of supercharges and $A, B, \dots = 0, 1, 2$ are the Lorentz indices. Let us note that \mathcal{B} is a central charge, while \mathcal{U} and \mathcal{T} are $\mathfrak{so}(2)$ internal symmetry generators. The presence of the central charge and internal symmetry generators are required to guarantee the non-degeneracy of the invariant tensor. In particular, the non-vanishing components of a non-degenerate invariant tensor of the superalgebra (4.1) are given by

$$\begin{aligned} \langle J_A J_B \rangle &= \alpha_0 \eta_{AB}, & \langle J_A P_B \rangle &= \alpha_1 \eta_{AB}, \\ \langle J_A Z_B \rangle &= \alpha_2 \eta_{AB}, & \langle P_A P_B \rangle &= \alpha_2 \eta_{AB}, \\ \langle P_A Z_B \rangle &= \frac{\alpha_1}{\ell^2} \eta_{AB}, & \langle Z_A Z_B \rangle &= \frac{\alpha_2}{\ell^2} \eta_{AB}, \\ \langle \mathcal{T} \mathcal{T} \rangle &= \alpha_0, & \langle \mathcal{T} \mathcal{U} \rangle &= \alpha_2, \\ \langle \mathcal{U} \mathcal{U} \rangle &= \frac{1}{\ell^2} \alpha_2, & \langle \mathcal{T} \mathcal{B} \rangle &= \alpha_1, \\ \langle \mathcal{U} \mathcal{B} \rangle &= \frac{1}{\ell^2} \alpha_1, & \langle \mathcal{B} \mathcal{B} \rangle &= -\frac{1}{\ell} \alpha_1, \\ \langle Q_\alpha^i Q_\beta^j \rangle &= 2 \left(\frac{\alpha_1}{\ell} + \alpha_2 \right) C_{\alpha\beta} \delta^{ij}. \end{aligned} \quad (4.2)$$

Although such superalgebra is well-defined since it allows us to reproduce a proper three-dimensional $\mathcal{N} = 2$ CS supergravity action in presence of a cosmological constant, its flat limit $\ell \rightarrow \infty$ is problematic. Indeed, the vanishing cosmological constant limit reproduces an $\mathcal{N} = 2$ exotic Maxwell supersymmetric CS action due to the behavior of the P_A generator which is no more expressed as a bilinear expression of the fermionic generators Q^i . Such feature is responsible of the label “non-standard” in the Maxwell superalgebra [47, 83]. A dual version of the non-standard super-Maxwell algebra being the supersymmetric extension of the Hietarinta–Maxwell algebra [84, 85] could overcome such difficulty by interchanging the role of the P_A and Z_A generators. Nevertheless, such approach will not be considered here.

Here, we shall see that a new and consistent NR superalgebra is obtained by considering an S -expansion of the non-standard superalgebra (4.1). Let us consider $S_E^{(2)} = \{\lambda_0, \lambda_1, \lambda_2, \lambda_3\}$ as the relevant abelian semigroup whose elements satisfy the multiplication law (3.5) with $\lambda_3 = 0_S$ being the zero element of the semigroup. Let $S_E^{(2)} = S_0 \cup S_1$ be a semigroup decomposition where

$$\begin{aligned} S_0 &= \{\lambda_0, \lambda_2, \lambda_3\}, \\ S_1 &= \{\lambda_1, \lambda_3\}. \end{aligned} \quad (4.3)$$

The decomposition (4.3) is said to be resonant since it satisfies the same structure than the subspaces $V_0 = \{J_0, P_0, Z_0, \mathcal{T}, \mathcal{U}, \mathcal{B}, Q_\alpha^+\}$ and $V_1 = \{J_a, P_a, Z_a, Q_\alpha^-\}$ of the $\mathcal{N} = 2$ superalgebra (4.1) with $a = 1, 2$ and where we have defined

$$Q_\alpha^\pm = \frac{1}{\sqrt{2}} \left(Q_\alpha^1 \pm \epsilon_{\alpha\beta} Q_\beta^2 \right). \quad (4.4)$$

A new NR superalgebra is obtained after applying a resonant $S_E^{(2)}$ -expansion of (4.1) and performing a 0_S -reduction. In particular, the expanded generators are related to the $\mathcal{N} = 2$ super AdS- \mathcal{L} ones through the semigroup elements as

$$\begin{aligned} \tilde{J} &= \lambda_0 J_0, & \tilde{S} &= \lambda_2 J_0, & \tilde{G}_a &= \lambda_1 J_a, \\ \tilde{H} &= \lambda_0 P_0, & \tilde{M} &= \lambda_2 P_0, & \tilde{P}_a &= \lambda_1 P_a, \\ \tilde{Z} &= \lambda_0 Z_0, & \tilde{T} &= \lambda_2 Z_0, & \tilde{Z}_a &= \lambda_1 Z_a, \\ \tilde{Q}_\alpha^+ &= \lambda_0 Q_\alpha^+, & \tilde{R}_\alpha &= \lambda_2 Q_\alpha^+, & \tilde{Q}_\alpha^- &= \lambda_1 Q_\alpha^-, \\ \tilde{Y}_1 &= \lambda_0 \mathcal{T}, & \tilde{U}_1 &= \lambda_0 \mathcal{U}, & \tilde{B}_1 &= \lambda_0 \mathcal{B}, \\ \tilde{Y}_2 &= \lambda_2 \mathcal{T}, & \tilde{U}_2 &= \lambda_2 \mathcal{U}, & \tilde{B}_2 &= \lambda_2 \mathcal{B}. \end{aligned} \quad (4.5)$$

The commutators of the expanded NR superalgebra are given by the bosonic EEB algebra (2.1) along with the following commutation relations:

$$\begin{aligned} [\tilde{J}, \tilde{Q}_\alpha^\pm] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{Q}_\beta^\pm, & [\tilde{J}, \tilde{R}_\alpha] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{H}, \tilde{Q}_\alpha^\pm] &= -\frac{1}{2\ell} (\gamma_0)_\alpha^\beta \tilde{Q}_\beta^\pm, \\ [\tilde{H}, \tilde{R}_\alpha] &= -\frac{1}{2\ell} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, & [\tilde{Z}, \tilde{Q}_\alpha^\pm] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{Q}_\beta^\pm, \\ [\tilde{Z}, \tilde{R}_\alpha] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{S}, \tilde{Q}_\alpha^+] &= -\frac{1}{2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, & [\tilde{M}, \tilde{Q}_\alpha^+] &= -\frac{1}{2\ell} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{T}, \tilde{Q}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_0)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{G}_a, \tilde{Q}_\alpha^+] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{Q}_\beta^-, & [\tilde{G}_a, \tilde{Q}_\alpha^-] &= -\frac{1}{2} (\gamma_a)_\alpha^\beta \tilde{R}_\beta, \\ [\tilde{P}_a, \tilde{Q}_\alpha^+] &= -\frac{1}{2\ell} (\gamma_a)_\alpha^\beta \tilde{Q}_\beta^-, \\ [\tilde{P}_a, \tilde{Q}_\alpha^-] &= -\frac{1}{2\ell} (\gamma_a)_\alpha^\beta \tilde{R}_\beta, & [\tilde{Z}_a, \tilde{Q}_\alpha^+] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{Q}_\beta^-, \\ [\tilde{Z}_a, \tilde{Q}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_a)_\alpha^\beta \tilde{R}_\beta, \end{aligned}$$

$$\begin{aligned} [\tilde{Y}_1, \tilde{Q}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^+, & [\tilde{Y}_1, \tilde{Q}_\alpha^-] &= -\frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^-, \\ [\tilde{Y}_1, \tilde{R}_\alpha] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, \\ [\tilde{Y}_2, \tilde{Q}_\alpha^+] &= \frac{1}{2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, & [\tilde{U}_1, \tilde{Q}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^+, \\ [\tilde{U}_1, \tilde{Q}_\alpha^-] &= -\frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{Q}_\beta^-, \\ [\tilde{U}_1, \tilde{R}_\alpha] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, & [\tilde{U}_2, \tilde{Q}_\alpha^+] &= \frac{1}{2\ell^2} (\gamma_0)_{\alpha\beta} \tilde{R}_\beta, \end{aligned} \quad (4.6)$$

while the fermionic generators satisfy the following anti-commutation relations:

$$\begin{aligned} \{\tilde{Q}_\alpha^+, \tilde{Q}_\beta^+\} &= -\frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \tilde{H} - (\gamma^0 C)_{\alpha\beta} \tilde{Z} \\ &\quad - (\gamma^0 C)_{\alpha\beta} \left(\tilde{U}_1 + \frac{1}{\ell} \tilde{B}_1 \right), \\ \{\tilde{Q}_\alpha^+, \tilde{Q}_\beta^-\} &= -\frac{1}{\ell} (\gamma^a C)_{\alpha\beta} \tilde{P}_a - (\gamma^a C)_{\alpha\beta} \tilde{Z}_a, \\ \{\tilde{Q}_\alpha^+, \tilde{R}_\beta\} &= -\frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \tilde{M} - (\gamma^0 C)_{\alpha\beta} \tilde{T} \\ &\quad - (\gamma^0 C)_{\alpha\beta} \left(\tilde{U}_2 + \frac{1}{\ell} \tilde{B}_2 \right), \\ \{\tilde{Q}_\alpha^-, \tilde{Q}_\beta^-\} &= -\frac{1}{\ell} (\gamma^0 C)_{\alpha\beta} \tilde{M} - (\gamma^0 C)_{\alpha\beta} \tilde{T} \\ &\quad + (\gamma^0 C)_{\alpha\beta} \left(\tilde{U}_2 + \frac{1}{\ell} \tilde{B}_2 \right). \end{aligned} \quad (4.7)$$

The expanded NR superalgebra corresponds to an alternative non-standard supersymmetric extension of the EEB algebra introduced in [36]. Unlike the previous EEB superalgebra obtained in this work, the NR centrally extended superalgebra (4.6) and (4.7) does not contain additional fermionic generators, reducing considerably the number of (anti-)commutators. One can notice that B_1 and B_2 are central charges, while $\{Y_1, Y_2\}$ and $\{U_1, U_2\}$ correspond to expansions of the relativistic R-symmetry generators \mathcal{T} and \mathcal{U} , respectively. Interestingly, both supersymmetric descriptions of the EEB algebra can be contracted through a vanishing cosmological constant limit $\ell \rightarrow \infty$. As we have shown, the flat limit applied to the EEB superalgebra given by (2.1), (3.12), (3.13), and (3.14) reproduces the MEB superalgebra [6] allowing to construct a consistent NR supergravity which generalizes the extended Bargmann supergravity [3]. On the other hand, although a flat limit can be done at the level of the non-standard EEB superalgebra (4.6) and (4.7), the resulting NR superalgebra does not guarantee the proper construction of a NR supergravity action. Similarly to the non-standard relativistic Maxwell superalgebra, the absence of the P_a generators in the anti-commutator after considering $\ell \rightarrow \infty$ would reproduce an exotic NR supersymmetric CS action. It would be interesting to avoid such difficulty by exploring a NR version of the non-standard Maxwell algebra using the Hietarinta basis [86].

One can notice that the EEB superalgebra (4.6) and (4.7) can be written alternatively as the direct sum of three copies of the Nappi–Witten algebra [79, 80] where one copy is supersymmetric. Indeed, one can consider the following redefinition of the generators,

$$\begin{aligned}\tilde{G}_a &= \hat{G}_a + G_a + G_a^*, \tilde{P}_a = \frac{1}{\ell} (G_a - G_a^*), \\ \tilde{Z}_a &= \frac{1}{\ell^2} (G_a + G_a^*), \\ \tilde{S} &= \hat{S} + S + S^*, \tilde{M} = \frac{1}{\ell} (S - S^*), \\ \tilde{T} &= \frac{1}{\ell^2} (S + S^*), \\ \tilde{J} &= \hat{J} + J + J^*, \tilde{H} = \frac{1}{\ell} (J - J^*), \\ \tilde{Z} &= \frac{1}{\ell^2} (J + J^*), \\ \tilde{Y}_1 &= \hat{T}_1 + T_1 + T_1^*, \tilde{B}_1 = \frac{1}{\ell} (T_1 - T_1^*), \\ \tilde{U}_1 &= \frac{1}{\ell^2} (T_1 + T_1^*), \\ \tilde{Y}_2 &= \hat{T}_2 + T_2 + T_2^*, \tilde{B}_2 = \frac{1}{\ell} (T_2 - T_2^*), \\ \tilde{U}_2 &= \frac{1}{\ell^2} (T_2 + T_2^*), \\ \tilde{Q}_\alpha^+ &= \frac{\sqrt{2}}{\ell} Q_\alpha^+, \tilde{Q}_\alpha^- = \frac{\sqrt{2}}{\ell} Q_\alpha^-, \\ \tilde{R}_\alpha &= \frac{\sqrt{2}}{\ell} R_\alpha,\end{aligned}\quad (4.8)$$

and see that both subsets spanned by $\{\hat{G}_a, \hat{S}, \hat{J}\}$ and $\{G_a^*, S^*, J^*\}$ define a Nappi–Witten algebra (3.16) coupled to $u(1)$ generators $\{\hat{T}_1, \hat{T}_2\}$ and $\{T_1^*, T_2^*\}$, respectively. On the other hand, the set of generators $\{G_a, S, J, T_1, T_2, Q_\alpha^+, Q_\alpha^-, R_\alpha\}$ satisfy a supersymmetric extension of the Nappi–Witten algebra whose (anti-)commutation relations are given by (3.17). It is interesting to point out that, although the non-standard EEB superalgebra seems to be quite different to the “standard” one studied previously, they actually differ just on the amount of supersymmetric copies of the Nappi–Witten algebra.

An alternative change of basis allows us to rewrite the EEB superalgebra (4.6) and (4.7) as the direct sum of the extended Newton–Hooke superalgebra and the Nappi–Witten algebra. Such structure can be obtained by considering the following redefinition of the generators:

$$\begin{aligned}\tilde{G}_a &= G_a^* + G_a, & \tilde{P}_a &= P_a, & \tilde{Z}_a &= \frac{1}{\ell^2} G_a, \\ \tilde{S} &= S^* + S, & \tilde{M} &= M, & \tilde{T} &= \frac{1}{\ell^2} S, \\ \tilde{J} &= J^* + J, & \tilde{H} &= H, & \tilde{Z} &= \frac{1}{\ell^2} J,\end{aligned}$$

$$\begin{aligned}\tilde{Y}_1 &= T_1^* + T_1, & \tilde{B}_1 &= B_1, & \tilde{U}_1 &= \frac{1}{\ell^2} T_1, \\ \tilde{Y}_2 &= T_2^* + T_2, & \tilde{B}_2 &= B_2, & \tilde{U}_2 &= \frac{1}{\ell^2} T_2, \\ \tilde{Q}_\alpha^+ &= \sqrt{\frac{1}{\ell}} Q_\alpha^+, & \tilde{Q}_\alpha^- &= \sqrt{\frac{1}{\ell}} Q_\alpha^-, & \tilde{R}_\alpha &= \sqrt{\frac{1}{\ell}} R_\alpha.\end{aligned}\quad (4.9)$$

The subset spanned by $\{G_a, P_a, S, M, J, H, T_1, T_2, Q_\alpha^+, Q_\alpha^-, R_\alpha\}$ corresponds to the extended Newton–Hooke superalgebra (3.19) (with, in this case, $U_1 = U_2 = 0$) now endowed with a central extension given by B_1 and B_2 . On the other hand, the set of generators $\{G_a^*, S^*, J^*, T_1^*, T_2^*\}$ satisfies the Nappi–Witten algebra (3.16) coupled to the $u(1)$ generators T_1^* and T_2^* . Here, we shall focus on the EEB superalgebra written as in (4.6) and (4.7) which, as we shall see, offers us an alternative way to introduce a cosmological constant in NR supergravity theory different from the one presented in the previous section and from the extended Newton–Hooke one [5].

4.2 Non-standard non-relativistic extended supergravity action

We can now construct a NR CS supergravity action based on the non-standard EEB superalgebra previously introduced. To this aim, let us start by writing the non-vanishing components of the invariant tensor for the non-standard EEB superalgebra, which can be determined by exploiting the S -expansion on (4.2) (see Appendix A). In particular, the non-standard EEB superalgebra admits the bosonic invariant tensor (2.4) along with the following non-vanishing components of the invariant tensor

$$\begin{aligned}\langle \tilde{Y}_1 \tilde{Y}_2 \rangle &= \tilde{\alpha}_0, & \langle \tilde{Y}_1 \tilde{B}_2 \rangle &= \tilde{\alpha}_1, & \langle \tilde{Y}_2 \tilde{B}_1 \rangle &= \tilde{\alpha}_1, \\ \langle \tilde{Y}_1 \tilde{U}_2 \rangle &= \tilde{\alpha}_2, & \langle \tilde{Y}_2 \tilde{U}_1 \rangle &= \tilde{\alpha}_2, & \langle \tilde{U}_1 \tilde{B}_2 \rangle &= \frac{\tilde{\alpha}_1}{\ell^2}, \\ \langle \tilde{U}_2 \tilde{B}_1 \rangle &= \frac{\tilde{\alpha}_1}{\ell^2}, & \langle \tilde{B}_1 \tilde{B}_2 \rangle &= \frac{\tilde{\alpha}_1}{\ell^2}, & \langle \tilde{U}_1 \tilde{U}_2 \rangle &= \frac{\tilde{\alpha}_2}{\ell^2}, \\ \langle \tilde{Q}_\alpha^- \tilde{Q}_\beta^- \rangle &= 2 \left(\frac{\tilde{\alpha}_1}{\ell} + \tilde{\alpha}_2 \right) C_{\alpha\beta}, \\ \langle \tilde{Q}_\alpha^+ \tilde{R}_\beta \rangle &= 2 \left(\frac{\tilde{\alpha}_1}{\ell} + \tilde{\alpha}_2 \right) C_{\alpha\beta},\end{aligned}\quad (4.10)$$

where $\tilde{\alpha}_0$, $\tilde{\alpha}_1$ and $\tilde{\alpha}_2$ are arbitrary constants which are related to the relativistic ones through (3.22).

The gauge connection one-form A for the non-standard EEB superalgebra is

$$\begin{aligned}A &= \omega \tilde{J} + \omega^a \tilde{G}_a + \tau \tilde{H} + e^a \tilde{P}_a + k \tilde{Z} + k^a \tilde{Z}_a + m \tilde{M} \\ &\quad + s \tilde{S} + t \tilde{T} + y_1 \tilde{Y}_1 + y_2 \tilde{Y}_2 \\ &\quad + b_1 \tilde{B}_1 + b_2 \tilde{B}_2 + u_1 \tilde{U}_1 + u_2 \tilde{U}_2 + \bar{\psi}^+ \tilde{Q}^+ \\ &\quad + \bar{\psi}^- \tilde{Q}^- + \bar{\rho} \tilde{R},\end{aligned}\quad (4.11)$$

and the corresponding curvature two-form is given by

$$\begin{aligned}
 F = & F(\omega)\tilde{J} + F^a(\omega^b)\tilde{G}_a + F(\tau)\tilde{H} + F^a(e^b)\tilde{P}_a \\
 & + F(k)\tilde{Z} + F^a(k^b)\tilde{Z}_a + F(m)\tilde{M} \\
 & + F(s)\tilde{S} + F(t)\tilde{T} + F(y_1)\tilde{Y}_1 + F(y_2)\tilde{Y}_2 \\
 & + F(b_1)\tilde{B}_1 + F(b_2)\tilde{B}_2 + F(u_1)\tilde{U}_1 \\
 & + F(u_2)\tilde{U}_2 + \nabla\bar{\psi}^+\tilde{Q}^+ + \nabla\bar{\psi}^-\tilde{Q}^- + \nabla\bar{\rho}\tilde{R}, \quad (4.12)
 \end{aligned}$$

where

$$\begin{aligned}
 F(\omega) &= R(\omega), \\
 F^a(\omega^b) &= R^a(\omega^b), \\
 F(\tau) &= R(\tau) + \frac{1}{2\ell}\bar{\psi}^+\gamma^0\psi^+, \\
 F^a(e^b) &= R^a(e^b) + \frac{1}{\ell}\bar{\psi}^+\gamma^a\psi^-, \\
 F(k) &= R(k) + \frac{1}{2}\bar{\psi}^+\gamma^0\psi^+, \\
 F^a(k^b) &= R^a(k^b) + \bar{\psi}^+\gamma^a\psi^-, \\
 F(m) &= R(m) + \frac{1}{\ell}\bar{\psi}^+\gamma^0\rho + \frac{1}{2\ell}\bar{\psi}^-\gamma^0\psi^-, \\
 F(s) &= R(s), \\
 F(t) &= R(t) + \bar{\psi}^+\gamma^0\rho + \frac{1}{2}\bar{\psi}^-\gamma^0\psi^-, \quad (4.13)
 \end{aligned}$$

being $R(\omega)$, $R^a(\omega^b)$, $R(\tau)$, $R^a(e^b)$, $R(k)$, $R^a(k^b)$, $R(m)$, $R(s)$, and $R(t)$ the bosonic EEB curvatures defined in (2.9), together with

$$\begin{aligned}
 F(y_1) &= dy_1, \\
 \hat{F}(y_2) &= dy_2, \\
 F(b_1) &= db_1 + \frac{1}{2\ell}\bar{\psi}^+\gamma^0\psi^+, \\
 F(b_2) &= db_2 + \frac{1}{\ell}\bar{\psi}^+\gamma^0\rho - \frac{1}{2\ell}\bar{\psi}^-\gamma^0\psi^-, \\
 F(u_1) &= du_1 + \frac{1}{2}\bar{\psi}^+\gamma^0\psi^+, \\
 F(u_2) &= du_2 + \bar{\psi}^+\gamma^0\rho - \frac{1}{2}\bar{\psi}^-\gamma^0\psi^-. \quad (4.14)
 \end{aligned}$$

On the other hand, the covariant derivatives of the spinor 1-form fields read

$$\begin{aligned}
 \nabla\psi^+ &= d\psi^+ + \frac{1}{2}\omega\gamma_0\psi^+ - \frac{1}{2}y_1\gamma_0\psi^+ + \frac{1}{2\ell}\tau\gamma_0\psi^+ \\
 &\quad + \frac{1}{2\ell^2}k\gamma_0\psi^+ - \frac{1}{2\ell^2}u_1\gamma_0\psi^+, \\
 \nabla\psi^- &= d\psi^- + \frac{1}{2}\omega\gamma_0\psi^- + \frac{1}{2}\omega^a\gamma_a\psi^+ + \frac{1}{2}y_1\gamma_0\psi^- + \frac{1}{2\ell}\tau\gamma_0\psi^- \\
 &\quad + \frac{1}{2\ell}e^a\gamma_a\psi^+ + \frac{1}{2\ell^2}k^a\gamma_a\psi^+ \\
 &\quad + \frac{1}{2\ell^2}k\gamma_0\psi^- + \frac{1}{2\ell^2}u_1\gamma_0\psi^-,
 \end{aligned}$$

$$\begin{aligned}
 \nabla\rho &= d\rho + \frac{1}{2}\omega\gamma_0\rho + \frac{1}{2}\omega^a\gamma_a\psi^- + \frac{1}{2}s\gamma_0\psi^+ \\
 &\quad - \frac{1}{2}y_2\gamma_0\psi^+ - \frac{1}{2}y_1\gamma_0\rho + \frac{1}{2\ell}e^a\gamma_a\psi^- \\
 &\quad + \frac{1}{2\ell^2}k^a\gamma_a\psi^- + \frac{1}{2\ell}m\gamma_0\psi^+ \\
 &\quad + \frac{1}{2\ell}\tau\gamma_0\rho + \frac{1}{2\ell^2}t\gamma_0\psi^+ + \frac{1}{2\ell^2}k\gamma_0\rho \\
 &\quad - \frac{1}{2\ell^2}u_1\gamma_0\rho - \frac{1}{2\ell^2}u_2\gamma_0\psi^+. \quad (4.15)
 \end{aligned}$$

A CS supergravity action based on the non-standard (NS) EEB superalgebra (4.6) and (4.7) can be constructed by combining the non-zero components of the invariant tensor (2.4) and (4.10) with the gauge connection one-form A (4.11). One can see that the explicit CS supergravity action can be split in three independent sectors:

$$I_{\text{NR}}^{\text{NS}} = \tilde{\alpha}_0\tilde{I}_0 + \tilde{\alpha}_1\tilde{I}_1 + \tilde{\alpha}_2\tilde{I}_2, \quad (4.16)$$

where

$$\begin{aligned}
 \tilde{I}_0 &= \int \omega_a R^a(\omega^b) - 2sR(\omega) + 2y_1dy_2, \\
 \tilde{I}_1 &= \int e_a R^a(\omega^b) + \omega_a R^a(e^b) - 2mR(\omega) - 2\tau ds \\
 &\quad - \frac{2}{\ell^2}mR(k) - \frac{2}{\ell^2}\tau dt + \frac{1}{\ell^2}e_a R^a(k^b) \\
 &\quad + \frac{1}{\ell^2}k_a R^a(e^b) + 2b_1dy_1 + 2b_1dy_2 \\
 &\quad + \frac{2}{\ell^2}u_1db_2 + \frac{2}{\ell^2}u_2db_1 - \frac{2}{\ell}b_1db_2 - \frac{2}{\ell}\bar{\psi}^-\nabla\psi^- \\
 &\quad - \frac{2}{\ell}\bar{\psi}^+\nabla\rho - \frac{2}{\ell}\bar{\rho}\nabla\psi^+, \\
 \tilde{I}_2 &= \int \omega_a R^a(k^b) + k_a R^a(\omega^b) + e_a R^a(e^b) \\
 &\quad - 2sR(k) - 2tR(\omega) - 2mR(\tau) + \frac{1}{\ell^2}k_a R^a(k^b) \\
 &\quad - \frac{2}{\ell^2}tR(k) + 2u_2dy_1 + 2u_1dy_2 + \frac{2}{\ell^2}u_1du_2 \\
 &\quad - 2\bar{\psi}^-\hat{\nabla}\psi^- - 2\bar{\psi}^+\hat{\nabla}\rho - 2\bar{\rho}\hat{\nabla}\psi^+. \quad (4.17)
 \end{aligned}$$

The CS action (4.17) obtained here describes a non-standard Enlarged Extended Bargmann supergravity model whose field equations are given by the vanishing of the curvature 2-form (4.12). Although it contains the bosonic EEB gravity theory, the non-standard version of the EEB superalgebra leads to a CS action diverse from the “standard” one (3.28) discussed in the previous section. Both supersymmetric descriptions contain the exotic NR gravity action $\tilde{I}_0 = I_0$ as particular subcase. However, \tilde{I}_1 and \tilde{I}_2 are quite different from the I_1 and I_2 contributions appearing in the standard case. Indeed, the fact that the non-standard EEB superalgebra contains only three different spinor generators reduces considerably the CS expression. Furthermore, the vanishing

cosmological constant limit $\ell \rightarrow \infty$ in the non-standard case reproduces an exotic supersymmetric theory. In particular, the flat limit applied to the \tilde{I}_1 contribution does not reproduce a supergravity theory anymore, but leads us to the extended Bargmann gravity [3] coupled to extra generators $\{y_1, y_2, b_1, b_2\}$. Such peculiar behavior is inherited from the relativistic counterpart and is responsible of the label “non-standard”. However, a different interpretation of the gauge field could overcome such difficulty to reproduce a supergravity theory in the flat limit by using the Hietarinta–Maxwell interpretation [85]. It would be interesting to explore the Physical implications of considering the Hietarinta basis at both the relativistic and the NR levels.

On the other hand, as in the “standard” case, the CS supergravity action presented here is restricted exclusively to the CS terms related to the λ_2 element of the semigroup $S_E^{(2)}$. The exotic-like terms produced due to the λ_0 element has been omitted intentionally. Indeed, such terms would modify the CS bosonic sector, meaning that one would have a supersymmetric extension of a diverse CS gravity theory.

An alternative procedure to recover the non-standard EEB supergravity action (4.17) can be performed by considering an $S_E^{(2)}$ -expansion of the relativistic $\mathcal{N} = 2$ non-standard AdS- \mathcal{L} CS supergravity action,⁴

$$\begin{aligned} I_R^{\text{NS}} = & \int \alpha_0 \left[\omega_A d\omega^A + \frac{1}{3} \epsilon_{ABC} \omega^A \omega^B \omega^C + t dt \right] \\ & + \alpha_1 \left[2e_A R^A + \frac{1}{3\ell^2} \epsilon_{ABC} e^A e^B e^C + \frac{2}{\ell^2} e_A F^A \right. \\ & \left. - \frac{2}{\ell} \tilde{\psi}^i \nabla \hat{\psi}^i + t db + \frac{1}{\ell^2} u db + \frac{1}{\ell} b du \right] \\ & + \alpha_2 \left[2\sigma_A R^A + e_A T^A + \frac{2}{\ell^2} e_A F^A + \frac{1}{\ell^2} \epsilon_{ABC} e^A \sigma^B e^C \right. \\ & \left. - 2\tilde{\psi}^i \nabla \hat{\psi}^i + t du + \frac{1}{\ell^2} u du \right], \end{aligned} \quad (4.18)$$

where R^A , T^A , and F^A are defined as in (3.31), while the covariant derivative of the fermionic gauge field is given by

$$\begin{aligned} \nabla \hat{\psi}^i = & d\hat{\psi}^i + \frac{1}{2} \omega^A \gamma_A \hat{\psi}^i + \frac{1}{2\ell} e^A \gamma_A \hat{\psi}^i + \frac{1}{2\ell^2} \sigma^A \gamma_A \hat{\psi}^i \\ & + t \epsilon^{ij} \hat{\psi}^j + \frac{1}{\ell^2} u \epsilon^{ij} \hat{\psi}^j. \end{aligned} \quad (4.19)$$

In fact, the NR gauge fields can be written in terms of the relativistic ones and the semigroup elements as

$$\begin{aligned} \omega &= \lambda_0 \omega_0, & s &= \lambda_2 \omega_0, & \omega_a &= \lambda_1 \omega_a, \\ \tau &= \lambda_0 e_0, & m &= \lambda_2 e_0, & e_a &= \lambda_1 e_a, \\ k &= \lambda_0 \sigma_0, & t &= \lambda_2 \sigma_0, & k_a &= \lambda_1 \sigma_a, \\ \psi_\alpha^+ &= \lambda_0 \hat{\psi}_\alpha^+, & \rho_\alpha &= \lambda_2 \hat{\psi}_\alpha^+, & \psi_\alpha^- &= \lambda_1 \hat{\psi}_\alpha^-, \end{aligned}$$

⁴ The relativistic CS action appears considering the non-vanishing components of the invariant tensor for the $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra (4.2).

$$\begin{aligned} y_1 &= \lambda_0 \mathbf{t}, & u_1 &= \lambda_0 \mathbf{u}, & b_1 &= \lambda_0 \mathbf{b}, \\ y_2 &= \lambda_2 \mathbf{t}, & u_2 &= \lambda_2 \mathbf{u}, & b_2 &= \lambda_2 \mathbf{b}, \end{aligned} \quad (4.20)$$

where we have performed the split $A = \{0, a\}$ with $a = 1, 2$ and where $\hat{\psi}_\alpha^\pm$ is defined as in (3.33). Thus, the non-standard EEB supergravity action can be obtained by considering the expanded gauge fields (4.20), the expanded parameters (3.22), the multiplication law (3.5), together with the 0_s -reduction property.

5 Discussion

In this work we have presented two diverse supersymmetric extensions of the enlarged extended Bargmann gravity in three spacetime dimensions introduced in [36]. The new NR superalgebras we have introduced differ on the number of fermionic generators and have been obtained from two different relativistic $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebras considering the semigroup expansion procedure [64]. We have shown that both EEB superalgebras allows us to introduce a cosmological constant to a NR supergravity different from the extended Newton–Hooke one [5]. Nevertheless, only the “standard” description reproduces a consistent vanishing cosmological constant limit $\ell \rightarrow \infty$ leading to the Maxwellian version of the extended Bargmann supergravity presented in [6]. Indeed, the flat limit of the non-standard version of the EEB supergravity leads us to an exotic supersymmetric theory since $\{Q, Q\} \sim Z_a$. One way to overcome such difficulty could be performed by interchanging the role of the P_a and Z_a generators as in the relativistic Hietarinta–Maxwell gravity theory [85]. Motivated by the fact that both topological and minimal massive gravity theories [87, 88] appear as particular cases of a generalized minimal massive gravity arising from a spontaneous breaking of the Hietarinta–Maxwell in a CS theory, it would be worth it to explore the Physical implications of a possible NR version of the three-dimensional Hietarinta–Maxwell CS (super)gravity.

On the other hand, both EEB superalgebras can be written in terms of three copies of the Nappi–Witten algebra where one or two copies are augmented by supersymmetry depending on the case. An alternative redefinition of the generators reveals that the (non-)standard EEB superalgebra can also be written as the direct sum of the extended Newton–Hooke superalgebra [5] and the Nappi–Witten (super)algebra. Such structures are inherited from their relativistic counterparts, in which case the AdS- \mathcal{L} superalgebras are isomorphic to three copies of the Lorentz (super)algebra and to the direct sum of the AdS superalgebra and the (super) Lorentz one. It would be interesting to explore if the supersymmetric extension of the Nappi–Witten algebra obtained here can be used to obtain novel superalgebras through the S -expansion procedure [89]. In particular, one could extend the results obtained in [36, 67, 73] to the supersymmetric case.

The S -expansion method used here is a powerful tool to obtain new NR superalgebras and to construct the respective NR CS supergravity theories. Consequently, the procedure performed here could serve as a starting point for diverse further studies. In particular, we have shown that the S -expansion of a relativistic $\mathcal{N} = 2$ AdS- \mathcal{L} superalgebra using $S_E^{(2)}$ as the relevant semigroup reproduces its NR counterpart. It would be interesting to first check if known NR superalgebras appears from $\mathcal{N} = 2$ superalgebras considering the same approach. Then, one could go further and study which kind of superalgebras can be obtained by considering $S_E^{(N)}$ -expansions for $N > 2$ to diverse relativistic superalgebras. On the other hand, one could apply our method to four and higher spacetime dimensions and construct NR supergravity models.

Another aspect that deserves further developments is the study of the asymptotic symmetry of the EEB (super)gravity theory. At the relativistic level, a semi-simple enlargement of the \mathfrak{bms}_3 algebra results to describe the boundary dynamics of the three-dimensional AdS- \mathcal{L} gravity [53]. Interestingly, such asymptotic structure can be written as three copies of the Virasoro algebra or as the direct sum of the conformal symmetry and the Virasoro one [90, 91]. A future development could consist on the study of consistent boundary conditions for the EEB gauge field to unveil the asymptotic symmetry of the theory. One could expect to obtain a canonical realization of a new infinite-dimensional symmetry which could be written as three copies of the NR version of the Virasoro algebra.

Besides, since these NR supersymmetry algebras share the presence of extra fermionic generators, it could be intriguing to study higher-dimensional cases and, in particular, to carry on an analysis on their possible hidden gauge structure in higher dimensions, on the same lines of [92–94]. This could give new insights in the NR regime of supergravity theories. In correspondence with the present results, it would be worth to analyze the NR counterpart of [94] in the presence of a cosmological constant.

Concerning other possible future developments, it would also be interesting to extend the study to the ultra-relativistic regime, where some results in the context of supergravity have been recently presented in [95, 96] [work in progress].

Acknowledgements This work was funded by the National Agency for Research and Development ANID (ex-CONICYT)-PAI Grant No. 77190078 (P.C.) and the FONDECYT Project N° 3170438 (E.R.). P.C. would like to thank to the Dirección de Investigación and Vice-rectoría de Investigación of the Universidad Católica de la Santísima Concepción, Chile, for their constant support. L.R. would like to thank A. Gamba and F. Dolcini for financial support.

Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors' comment: The results obtained and presented in the manuscript did not require any data.]

Open Access This article is licensed under a Creative Commons Attribution 4.0 International License, which permits use, sharing, adaptation, distribution and reproduction in any medium or format, as long as you give appropriate credit to the original author(s) and the source, provide a link to the Creative Commons licence, and indicate if changes were made. The images or other third party material in this article are included in the article's Creative Commons licence, unless indicated otherwise in a credit line to the material. If material is not included in the article's Creative Commons licence and your intended use is not permitted by statutory regulation or exceeds the permitted use, you will need to obtain permission directly from the copyright holder. To view a copy of this licence, visit <http://creativecommons.org/licenses/by/4.0/>. Funded by SCOAP³.

Appendix A: Invariant tensor and semigroup expansion method

The semigroup expansion method, introduced in [64] and latter studied in [97–99], consists in combining the elements of a semigroup S with the structure constant of a Lie (super)algebra \mathfrak{g} . The new Lie (super)algebra $\mathfrak{G} = S \times \mathfrak{g}$ is said to be an S -expanded (super)algebra whose structure constants are related to the structure constant of the original Lie (super)algebra \mathfrak{g} as

$$C_{(A,\alpha)(B,\beta)}^{(C,\gamma)} = K_{\alpha\beta}^{\gamma} C_{AB}^C, \quad (\text{A.1})$$

where $K_{\alpha\beta}^{\gamma}$ is the 2-selector which encodes the information from the multiplication law of the semigroup S and satisfy

$$K_{\alpha\beta}^{\gamma} = \begin{cases} 1 & \text{when } \gamma = \rho(\alpha, \beta), \\ 0 & \text{otherwise.} \end{cases} \quad (\text{A.2})$$

Interestingly, the S -expansion procedure allows us to obtain the invariant tensor of the expanded Lie (super)algebra \mathfrak{G} . The invariant tensor is a crucial ingredient for the construction of a CS action based on the expanded Lie (super)algebra.

According to [64], let us consider \mathfrak{g} being a Lie (super)algebra spanned by $\{T_A\}$ and let S be an abelian semigroup. If the original Lie (super)algebra \mathfrak{g} admits $\langle T_{A_1} \cdots T_{A_n} \rangle$ as the invariant tensor, then the expanded Lie (super)algebra $\mathfrak{G} = S \times \mathfrak{g}$ admits the following invariant tensor,

$$\langle T_{A_1, \alpha_1} \cdots T_{A_n, \alpha_n} \rangle = \alpha_{\gamma} K_{\alpha_1 \cdots \alpha_n}^{\gamma} \langle T_{A_1} \cdots T_{A_n} \rangle, \quad (\text{A.3})$$

where the α_{γ} 's are arbitrary constants. Here $K_{\alpha_1 \cdots \alpha_n}^{\gamma}$ is the n -selector for the semigroup S which is defined as

$$K_{\alpha_1 \cdots \alpha_n}^{\gamma} = \begin{cases} 1 & \text{when } \gamma = \rho(\alpha_1, \dots, \alpha_n) \\ 0 & \text{otherwise.} \end{cases} \quad (\text{A.4})$$

The possibility to find the invariant tensor for an expanded Lie (super)algebra is an additional advantage of the S -expansion method. Furthermore, the S -expansion procedure

allows us to obtain new expanded Lie (super)algebras which cannot be obtained through other Lie (super)algebra expansion methods. Indeed, an S -expansion performed by using the particular semigroup $S_E^{(N)}$ reproduces the expanded Lie (super)algebras derived through the Maurer-Cartan forms power series expansion [62, 63]. A different choice of the semigroup would lead to other expanded Lie (super)algebras which cannot be obtained through the Maurer-Cartan formalism.

Appendix B: Gauge transformations

The supersymmetry transformation laws of the EEB superalgebra read

$$\begin{aligned}
 \delta\omega &= 0, & \delta\omega^a &= 0, & \delta\tau &= -\bar{\epsilon}^+\gamma^0\psi^+ \\
 & & & & & -\frac{1}{\ell^2}\bar{\varphi}^+\gamma^0\xi^+, \\
 \delta e^a &= -\bar{\epsilon}^+\gamma^a\psi^- - \bar{\epsilon}^-\gamma^a\psi^+ - \frac{1}{\ell^2}\bar{\varphi}^+\gamma^a\xi^- - \frac{1}{\ell^2}\bar{\varphi}^-\gamma^a\xi^+, \\
 \delta k &= -\bar{\epsilon}^+\gamma^0\xi^+ - \bar{\varphi}^+\gamma^0\psi^+, \\
 \delta k^a &= -\bar{\epsilon}^+\gamma^a\xi^- - \bar{\epsilon}^-\gamma^a\xi^+ - \bar{\varphi}^+\gamma^a\psi^- - \bar{\varphi}^-\gamma^a\psi^+, \\
 \delta m &= -\bar{\epsilon}^-\gamma^0\psi^- - \bar{\epsilon}^+\gamma^0\rho - \bar{\eta}\gamma^0\psi^+ - \frac{1}{\ell^2}\bar{\varphi}^-\gamma^0\xi^- \\
 & & & & & -\frac{1}{\ell^2}\bar{\varphi}^+\gamma^0\chi - \frac{1}{\ell^2}\bar{\xi}\gamma^0\xi^+, \\
 \delta s &= 0, \\
 \delta t &= -\bar{\epsilon}^-\gamma^0\xi^- - \bar{\varphi}^-\gamma^0\psi^- - \bar{\epsilon}^+\gamma^0\chi \\
 & & & & & -\bar{\xi}\gamma^0\psi^+ - \bar{\varphi}^+\gamma^0\rho - \bar{\eta}\gamma^0\xi^+, \\
 \delta y_1 &= 0, \\
 \delta y_2 &= 0, \\
 \delta b_1 &= -\bar{\epsilon}^+\gamma^0\xi^+ - \bar{\varphi}^+\gamma^0\psi^+, \\
 \delta b_2 &= \bar{\epsilon}^-\gamma^0\xi^- + \bar{\varphi}^-\gamma^0\psi^- - \bar{\epsilon}^+\gamma^0\chi - \bar{\xi}\gamma^0\psi^+ \\
 & & & & & -\bar{\varphi}^+\gamma^0\rho - \bar{\eta}\gamma^0\xi^+, \\
 \delta u_1 &= -\bar{\epsilon}^+\gamma^0\psi^+ - \frac{1}{\ell^2}\bar{\varphi}^+\gamma^0\xi^+, \\
 \delta u_2 &= \bar{\epsilon}^-\gamma^0\psi^- - \bar{\epsilon}^+\gamma^0\rho - \bar{\eta}\gamma^0\psi^+ + \frac{1}{\ell^2}\bar{\varphi}^-\gamma^0\xi^- \\
 & & & & & -\frac{1}{\ell^2}\bar{\varphi}^+\gamma^0\chi - \frac{1}{\ell^2}\bar{\xi}\gamma^0\xi^+, \\
 \delta\psi^+ &= d\epsilon^+ + \frac{1}{2}\omega\gamma_0\epsilon^+ - \frac{1}{2}y_1\gamma_0\epsilon^+ + \frac{1}{2\ell^2}\tau\gamma_0\epsilon^+ + \frac{1}{2\ell^2}k\gamma_0\epsilon^+ \\
 & & & & & -\frac{1}{2\ell^2}u_1\gamma_0\epsilon^+ - \frac{1}{2\ell^2}b_1\gamma_0\epsilon^+, \\
 \delta\psi^- &= d\epsilon^- + \frac{1}{2}\omega\gamma_0\epsilon^- + \frac{1}{2}\omega^a\gamma_a\epsilon^+ + \frac{1}{2}y_1\gamma_0\epsilon^- + \frac{1}{2\ell^2}\tau\gamma_0\epsilon^- \\
 & & & & & + \frac{1}{2\ell^2}e^a\gamma_a\epsilon^+ + \frac{1}{2\ell^2}k^a\gamma_a\epsilon^+ \\
 & & & & & + \frac{1}{2\ell^2}k\gamma_0\epsilon^- + \frac{1}{2\ell^2}u_1\gamma_0\epsilon^- + \frac{1}{2\ell^2}b_1\gamma_0\epsilon^-, \\
 \delta\xi^+ &= d\varphi^+ + \frac{1}{2}\omega\gamma_0\varphi^+ + \frac{1}{2}\tau\gamma_0\epsilon^+ - \frac{1}{2}y_1\gamma_0\varphi^+ \\
 & & & & & -\frac{1}{2}u_1\gamma_0\epsilon^+ + \frac{1}{2\ell^2}k\gamma_0\varphi^+ - \frac{1}{2\ell^2}b_1\gamma_0\varphi^+,
 \end{aligned}$$

$$\begin{aligned}
 \delta\xi^- &= d\varphi^- + \frac{1}{2}\omega\gamma_0\varphi^- + \frac{1}{2}\tau\gamma_0\epsilon^- + \frac{1}{2}e^a\gamma_a\epsilon^+ \\
 & & & & & + \frac{1}{2}\omega^a\gamma_a\varphi^+ + \frac{1}{2}y_1\gamma_0\varphi^- + \frac{1}{2}u_1\gamma_0\epsilon^- \\
 & & & & & + \frac{1}{2\ell^2}k^a\gamma_a\varphi^+ + \frac{1}{2\ell^2}k\gamma_0\varphi^- + \frac{1}{2\ell^2}b_1\gamma_0\varphi^-, \\
 \delta\rho &= d\eta + \frac{1}{2}\omega\gamma_0\eta + \frac{1}{2}\omega^a\gamma_a\epsilon^- + \frac{1}{2}s\gamma_0\epsilon^+ \\
 & & & & & -\frac{1}{2}y_2\gamma_0\epsilon^+ - \frac{1}{2}y_1\gamma_0\eta + \frac{1}{2\ell^2}e^a\gamma_a\varphi^- \\
 & & & & & + \frac{1}{2\ell^2}k^a\gamma_a\epsilon^- + \frac{1}{2\ell^2}m\gamma_0\varphi^+ + \frac{1}{2\ell^2}\tau\gamma_0\zeta \\
 & & & & & + \frac{1}{2\ell^2}t\gamma_0\epsilon^+ + \frac{1}{2\ell^2}k\gamma_0\eta \\
 & & & & & -\frac{1}{2\ell^2}u_1\gamma_0\zeta - \frac{1}{2\ell^2}u_2\gamma_0\varphi^+ - \frac{1}{2\ell^2}b_1\gamma_0\eta - \frac{1}{2\ell^2}b_2\gamma_0\epsilon^+, \\
 \delta\chi &= d\zeta + \frac{1}{2}\omega\gamma_0\zeta + \frac{1}{2}\omega^a\gamma_a\varphi^- + \frac{1}{2}e^a\gamma_a\epsilon^- + \frac{1}{2}\tau\gamma_0\eta \\
 & & & & & + \frac{1}{2}s\gamma_0\varphi^+ + \frac{1}{2}m\gamma_0\epsilon^+ \\
 & & & & & -\frac{1}{2}y_2\gamma_0\varphi^+ - \frac{1}{2}y_1\gamma_0\zeta - \frac{1}{2}u_2\gamma_0\epsilon^+ - \frac{1}{2}u_1\gamma_0\eta \\
 & & & & & + \frac{1}{2\ell^2}k^a\gamma_a\varphi^- + \frac{1}{2\ell^2}t\gamma_0\varphi^+ + \frac{1}{2\ell^2}k\gamma_0\zeta \\
 & & & & & -\frac{1}{2\ell^2}b_1\gamma_0\zeta - \frac{1}{2\ell^2}b_2\gamma_0\varphi^+,
 \end{aligned} \tag{B.1}$$

where ϵ^\pm , φ^\pm , η , and ζ are the fermionic gauge parameters related to \hat{Q}^\pm , $\hat{\Sigma}^\pm$, \hat{R} , and \hat{W} , respectively. Let us observe that by taking the flat limit $\ell \rightarrow \infty$ of one recovers the supersymmetry transformations under which the curvatures associated with the MEB superalgebra transform in a covariant way (see [6]).

References

1. R. Andringa, E.A. Bergshoeff, J. Rosseel, E. Sezgin, 3D Newton–Cartan supergravity. *Class. Quantum Gravity* **30**, 205005 (2013). [arXiv:1305.6737](#)
2. E. Bergshoeff, J. Rosseel, T. Zojer, Newton–Cartan supergravity with torsion and Schrödinger supergravity. *JHEP* **11**, 180 (2015). [arXiv:1509.04527](#)
3. E.A. Bergshoeff, J. Rosseel, Three-dimensional extended Bargmann supergravity. *Phys. Rev. Lett.* **116**(25), 251601 (2016). [arXiv:1604.08042](#)
4. N. Ozdemir, M. Ozkan, O. Tunca, U. Zorba, Three-dimensional extended Newtonian (super)gravity. *JHEP* **05**, 130 (2019). [arXiv:1903.09377](#)
5. N. Ozdemir, M. Ozkan, U. Zorba, Three-dimensional extended Lifshitz, Schrödinger and Newton–Hooke supergravity. *JHEP* **11**, 052 (2019). [arXiv:1909.10745](#)
6. P. Concha, L. Ravera, E. Rodríguez, Three-dimensional Maxwellian extended Bargmann supergravity. *JHEP* **04**, 051 (2020). [arXiv:1912.09477](#)
7. D. Son, Toward an AdS/cold atoms correspondence: a geometric realization of the Schrödinger symmetry. *Phys. Rev. D* **78**, 046003 (2008). [arXiv:0804.3972](#)
8. K. Balasubramanian, J. McGreevy, Gravity duals for non-relativistic CFTs. *Phys. Rev. Lett.* **101**, 061601 (2008). [arXiv:0804.4053](#)

9. S. Kachru, X. Liu, M. Mulligan, Gravity duals of Lifshitz-like fixed points. *Phys. Rev. D* **78**, 106005 (2008). [arXiv:0808.1725](#)
10. A. Bagchi, R. Gopakumar, Galilean conformal algebras and AdS/CFT. *JHEP* **07**, 037 (2009). [arXiv:0902.1385](#)
11. A. Bagchi, R. Gopakumar, I. Mandal, A. Miwa, GCA in 2d. *JHEP* **08**, 004 (2010). [arXiv:0912.1090](#)
12. M.H. Christensen, J. Hartong, N.A. Obers, B. Rollier, Torsional Newton–Cartan geometry and Lifshitz holography. *Phys. Rev. D* **89**, 061901 (2014). [arXiv:1311.4794](#)
13. M.H. Christensen, J. Hartong, N.A. Obers, B. Rollier, Boundary stress-energy tensor and Newton–Cartan geometry in Lifshitz holography. *JHEP* **01**, 057 (2014). [arXiv:1311.6471](#)
14. J. Hartong, E. Kiritsis, N.A. Obers, Lifshitz space-times for Schrödinger holography. *Phys. Lett. B* **746**, 318–324 (2015). [arXiv:1409.1519](#)
15. J. Hartong, E. Kiritsis, N.A. Obers, Schrödinger invariance from Lifshitz isometries in holography and field theory. *Phys. Rev. D* **92**, 066003 (2015). [arXiv:1409.1522](#)
16. J. Hartong, E. Kiritsis, N.A. Obers, Field theory on Newton–Cartan backgrounds and symmetries of the Lifshitz vacuum. *JHEP* **08**, 006 (2015). [arXiv:1502.00228](#)
17. M. Taylor, Lifshitz holography. *Class. Quantum Gravity* **33**(3), 033001 (2016). [arXiv:1512.03554](#)
18. C. Hoyos, D.T. Son, Hall viscosity and electromagnetic response. *Phys. Rev. Lett.* **108**, 066805 (2012). [arXiv:1109.2651](#)
19. D.T. Son, Newton–Cartan geometry and the quantum hall effect. [arXiv:1306.0638](#)
20. A.G. Abanov, A. Gromov, Electromagnetic and gravitational responses of two-dimensional noninteracting electrons in a background magnetic field. *Phys. Rev. B* **90**(1), 014435 (2014). [arXiv:1401.3703](#)
21. M. Geracie, K. Prabhu, M.M. Roberts, Curved non-relativistic spacetimes, Newtonian gravitation and massive matter. *J. Math. Phys.* **56**(10), 103505 (2015). [arXiv:1503.02682](#)
22. A. Gromov, K. Jensen, A.G. Abanov, Boundary effective action for quantum Hall states. *Phys. Rev. Lett.* **116**(12), 126802 (2016). [arXiv:1506.07171](#)
23. G. Festuccia, N. Seiberg, Rigid supersymmetric theories in curved superspace. *JHEP* **06**, 114 (2011). [arXiv:1105.0689](#)
24. V. Pestun, Localization of gauge theory on a four-sphere and supersymmetric Wilson loops. *Commun. Math. Phys.* **313**, 71–129 (2012). [arXiv:0712.2824](#)
25. M. Mariño, Lectures on non-perturbative effects in large N gauge theories, matrix models and strings. *Fortsch. Phys.* **62**, 455–540 (2014). [arXiv:1206.6272](#)
26. E. Witten, (2+1)-dimensional gravity as an exactly soluble system. *Nucl. Phys. B* **311**, 46 (1988)
27. A. Achúcarro, P. Townsend, A Chern–Simons action for three-dimensional anti-De Sitter supergravity theories. *Phys. Lett. B* **180**, 89 (1986)
28. J. Zanelli, Lecture notes on Chern–Simons (super-)gravities, Second edition (February 2008), in *7th Mexican Workshop on Particles and Fields* vol. 2 (2005). [arXiv:hep-th/0502193](#)
29. J. Lévy-Leblond, Group theory and its applications, vol. I
30. D. Grigore, The projective unitary irreducible representations of the Galilei group in (1+2)-dimensions. *J. Math. Phys.* **37**, 460–473 (1996). [arXiv:hep-th/9312048](#)
31. S. Bose, The Galilean group in (2+1) space-times and its central extension. *Commun. Math. Phys.* **169**, 385–396 (1995)
32. C. Duval, P. Horvathy, The ‘Peierls substitution’ and the exotic Galilei group. *Phys. Lett. B* **479**, 284–290 (2000). [arXiv:hep-th/0002233](#)
33. R. Jackiw, V. Nair, Anyon spin and the exotic central extension of the planar Galilei group. *Phys. Lett. B* **480**, 237–238 (2000). [arXiv:hep-th/0003130](#)
34. P.A. Horvathy, M.S. Plyushchay, Non-relativistic anyons, exotic Galilean symmetry and noncommutative plane. *JHEP* **06**, 033 (2002). [arXiv:hep-th/0201228](#)
35. G. Papageorgiou, B.J. Schroers, A Chern–Simons approach to Galilean quantum gravity in 2+1 dimensions. *JHEP* **11**, 009 (2009). [arXiv:0907.2880](#)
36. P. Concha, E. Rodríguez, Non-relativistic gravity theory based on an enlargement of the extended Bargmann algebra. *JHEP* **07**, 085 (2019). [arXiv:1906.00086](#)
37. R. Aldrovandi, A. Barbosa, L. Crispino, J. Pereira, Non-relativistic spacetimes with cosmological constant. *Class. Quantum Gravity* **16**, 495–506 (1999). [arXiv:gr-qc/9801100](#)
38. G. Gibbons, C. Patricot, Newton–Hooke space-times, Hpp waves and the cosmological constant. *Class. Quantum Gravity* **20**, 5225 (2003). [arXiv:hep-th/0308200](#)
39. J. Bruges, J. Gomis, K. Kamimura, Newton–Hooke algebras, non-relativistic branes and generalized pp-wave metrics. *Phys. Rev. D* **73**, 085011 (2006). [arXiv:hep-th/0603023](#)
40. P.D. Alvarez, J. Gomis, K. Kamimura, M.S. Plyushchay, (2+1)D exotic Newton–Hooke symmetry, duality and projective phase. *Ann. Phys.* **322**, 1556–1586 (2007). [arXiv:hep-th/0702014](#)
41. C. Duval, P. Horvathy, Conformal Galilei groups, veronese curves, and Newton–Hooke spacetimes. *J. Phys. A* **44**, 335203 (2011). [arXiv:1104.1502](#)
42. G. Papageorgiou, B.J. Schroers, Galilean quantum gravity with cosmological constant and the extended q -Heisenberg algebra. *JHEP* **11**, 020 (2010). [arXiv:1008.0279](#)
43. J. Hartong, Y. Lei, N.A. Obers, Nonrelativistic Chern–Simons theories and three-dimensional Hořava–Lifshitz gravity. *Phys. Rev. D* **94**(6), 065027 (2016). [arXiv:1604.08054](#)
44. C. Duval, G. Gibbons, P. Horvathy, Conformal and projective symmetries in Newtonian cosmology. *J. Geom. Phys.* **112**, 197–209 (2017). [arXiv:1605.00231](#)
45. P. Concha, L. Ravera, E. Rodríguez, Three-dimensional exotic Newtonian gravity with cosmological constant. *Phys. Lett. B* **804**, 135392 (2020). [arXiv:1912.02836](#)
46. L. Avilés, E. Frodden, J. Gomis, D. Hidalgo, J. Zanelli, Non-relativistic Maxwell Chern–Simons gravity. *JHEP* **05**, 047 (2018). [arXiv:1802.08453](#)
47. D.V. Soroka, V.A. Soroka, Tensor extension of the Poincaré algebra. *Phys. Lett. B* **607**, 302–305 (2005). [arXiv:hep-th/0410012](#)
48. D.V. Soroka, V.A. Soroka, Semi-simple extension of the (super)Poincaré algebra. *Adv. High Energy Phys.* **2009**, 234147 (2009). [arXiv:hep-th/0605251](#)
49. J. Gomis, K. Kamimura, J. Lukierski, Deformations of Maxwell algebra and their dynamical realizations. *JHEP* **08**, 039 (2009). [arXiv:0906.4464](#)
50. P. Concha, R. Durka, C. Inostroza, N. Merino, E. Rodríguez, Pure lovelock gravity and Chern–Simons theory. *Phys. Rev. D* **94**(2), 024055 (2016). [arXiv:1603.09424](#)
51. P. Concha, N. Merino, E. Rodríguez, Lovelock gravities from Born–Infeld gravity theory. *Phys. Lett. B* **765**, 395–401 (2017). [arXiv:1606.07083](#)
52. P. Concha, E. Rodríguez, Generalized pure lovelock gravity. *Phys. Lett. B* **774**, 616–622 (2017). [arXiv:1708.08827](#)
53. P. Concha, N. Merino, E. Rodríguez, P. Salgado-Rebolledo, O. Valdivia, Semi-simple enlargement of the \mathfrak{bms}_3 algebra from a $\mathfrak{so}(2, 2) \oplus \mathfrak{so}(2, 1)$ Chern–Simons theory. *JHEP* **02**, 002 (2019). [arXiv:1810.12256](#)
54. P. Concha, E. Rodríguez, P. Salgado, Generalized supersymmetric cosmological term in $N = 1$ supergravity. *JHEP* **08**, 009 (2015). [arXiv:1504.01898](#)
55. M. Ipinza, P. Concha, L. Ravera, E. Rodríguez, On the supersymmetric extension of Gauss–Bonnet like gravity. *JHEP* **09**, 007 (2016). [arXiv:1607.00373](#)

56. A. Banaudi, L. Ravera, Generalized AdS-Lorentz deformed supergravity on a manifold with boundary. *Eur. Phys. J. Plus* **133**(12), 514 (2018). [arXiv:1803.08738](#)
57. D.M. Peñañiel, L. Ravera, Generalized cosmological term in $D = 4$ supergravity from a new AdS-Lorentz superalgebra. *Eur. Phys. J. C* **78**(11), 945 (2018). [arXiv:1807.07673](#)
58. P. Concha, L. Ravera, E. Rodríguez, On the supersymmetry invariance of flat supergravity with boundary. *JHEP* **01**, 192 (2019). [arXiv:1809.07871](#)
59. P. Concha, D.M. Peñañiel, E. Rodríguez, On the Maxwell supergravity and flat limit in $2 + 1$ dimensions. *Phys. Lett. B* **785**, 247–253 (2018). [arXiv:1807.00194](#)
60. P. Concha, \mathcal{N} -extended Maxwell supergravities as Chern–Simons theories in three spacetime dimensions. *Phys. Lett. B* **792**, 290–297 (2019). [arXiv:1903.03081](#)
61. M. Hatsuda, M. Sakaguchi, Wess–Zumino term for the AdS superstring and generalized Inonu–Wigner contraction. *Prog. Theor. Phys.* **109**, 853–867 (2003). [arXiv:hep-th/0106114](#)
62. J.A. de Azcarraga, J.M. Izquierdo, M. Picon, O. Varela, Generating Lie and gauge free differential (super)algebras by expanding Maurer–Cartan forms and Chern–Simons supergravity. *Nucl. Phys. B* **662**, 185–219 (2003). [arXiv:hep-th/0212347](#)
63. J. de Azcarraga, J. Izquierdo, M. Picon, O. Varela, Expansions of algebras and superalgebras and some applications. *Int. J. Theor. Phys.* **46**, 2738–2752 (2007). [arXiv:hep-th/0703017](#)
64. F. Izaurieta, E. Rodríguez, P. Salgado, Expanding lie (super)algebras through Abelian semigroups. *J. Math. Phys.* **47**, 123512 (2006). [arXiv:hep-th/0606215](#)
65. E. Bergshoeff, J.M. Izquierdo, T. Ortín, L. Romano, Lie algebra expansions and actions for non-relativistic gravity. *JHEP* **08**, 048 (2019). [arXiv:1904.08304](#)
66. J.A. de Azcarraga, D. Gútiez, J.M. Izquierdo, Extended $D = 3$ Bargmann supergravity from a Lie algebra expansion. [arXiv:1904.12786](#)
67. D.M. Peñañiel, P. Salgado-Rebolledo, Non-relativistic symmetries in three space-time dimensions and the Nappi–Witten algebra. *Phys. Lett. B* **798**, 135005 (2019). [arXiv:1906.02161](#)
68. L. Romano, Non-relativistic four dimensional p-Brane supersymmetric theories and lie algebra expansion. [arXiv:1906.08220](#)
69. T. Harmark, J. Hartong, L. Menculini, N.A. Obers, G. Oling, Relating non-relativistic string theories. *JHEP* **11**, 071 (2019). [arXiv:1907.01663](#)
70. J. Gomis, A. Kleinschmidt, J. Palmkvist, P. Salgado-Rebolledo, Newton–Hooke/Carrollian expansions of (A)dS and Chern–Simons gravity. *JHEP* **02**, 009 (2020). [arXiv:1912.07564](#)
71. E. Bergshoeff, J. Gomis, P. Salgado-Rebolledo, Non-relativistic limits and three-dimensional coadjoint Poincare gravity. [arXiv:2001.11790](#)
72. O. Kasikci, N. Ozdemir, M. Ozkan, U. Zorba, Three-dimensional higher-order Schrödinger algebras and Lie algebra expansions. *JHEP* **04**, 067 (2020). [arXiv:2002.03558](#)
73. P. Concha, M. Ipinza, E. Rodríguez, Generalized Maxwellian exotic Bargmann gravity theory in three spacetime dimensions. *Phys. Lett. B* **807**, 135593 (2020). [arXiv:2004.01203](#)
74. A. Fontanella, L. Romano, Lie algebra expansion and integrability in superstring sigma-models. [arXiv:2005.01736](#)
75. P. Concha, L. Ravera, E. Rodríguez, G. Rubio, Three-dimensional Maxwellian extended Newtonian gravity and flat limit. [arXiv:2006.13128](#)
76. M. Ipinza, F. Lingua, D. Peñañiel, L. Ravera, An analytic method for S -expansion involving resonance and reduction. *Fortsch. Phys.* **64**(11–12), 854–880 (2016). [arXiv:1609.05042](#)
77. D. Peñañiel, L. Ravera, Infinite S -expansion with ideal subtraction and some applications. *J. Math. Phys.* **58**(8), 081701 (2017). [arXiv:1611.05812](#)
78. E. Inonu, E.P. Wigner, On the contraction of groups and their representations. *Proc. Natl. Acad. Sci.* **39**, 510–524 (1953)
79. C.R. Nappi, E. Witten, A WZW model based on a non-semisimple group. *Phys. Rev. Lett.* **71**, 3751–3753 (1993). [arXiv:hep-th/9310112](#)
80. J.M. Figueroa-O’Farrill, S. Stanciu, More D-branes in the Nappi–Witten background. *JHEP* **01**, 024 (2000). [arXiv:hep-th/9909164](#)
81. P.S. Howe, J. Izquierdo, G. Papadopoulos, P. Townsend, New supergravities with central charges and Killing spinors in $(2+1)$ -dimensions. *Nucl. Phys. B* **467**, 183–214 (1996). [arXiv:hep-th/9505032](#)
82. P. Concha, R. Durka, E. Rodríguez, Resonant superalgebras and $\mathcal{N} = 1$ supergravity theories in three spacetime dimensions. *Phys. Lett. B* **808**, 135659 (2020). [arXiv:2005.11803](#)
83. J. Lukierski, Generalized Wigner–Inönü contractions and Maxwell (super)algebras. *Proc. Steklov Inst. Math.* **272**(1), 183–190 (2011). [arXiv:1007.3405](#)
84. S. Bansal, D. Sorokin, Can Chern–Simons or Rarita–Schwinger be a Volkov–Akulov goldstone? *JHEP* **07**, 106 (2018). [arXiv:1806.05945](#)
85. D. Chernyavsky, N.S. Deger, D. Sorokin, Spontaneously broken $3d$ Hietarinta/Maxwell Chern–Simons theory and minimal massive gravity. *Eur. Phys. J. C* **80**(6), 556 (2020). [arXiv:2002.07592](#)
86. J. Hietarinta, Supersymmetry generators of arbitrary spin. *Phys. Rev. D* **13**, 838 (1976)
87. S. Deser, R. Jackiw, S. Templeton, Topologically massive gauge theories. *Ann. Phys.* **140**, 372–411 (1982). (Erratum: *Ann. Phys.* **185**, 406 (1988); *Ann. Phys.* **281**, 409 (2000))
88. E. Bergshoeff, O. Hohm, W. Merbis, A.J. Routh, P.K. Townsend, Minimal massive 3D gravity. *Class. Quantum Gravity* **31**, 145008 (2014). [arXiv:1404.2867](#)
89. P. Concha, M. Ipinza, L. Ravera, E. Rodríguez (**In preparation**)
90. R. Caroca, P. Concha, E. Rodríguez, P. Salgado-Rebolledo, Generalizing the \mathfrak{bms}_3 and 2D-conformal algebras by expanding the Virasoro algebra. *Eur. Phys. J. C* **78**(3), 262 (2018). [arXiv:1707.07209](#)
91. P. Concha, H. Safari, On stabilization of Maxwell–BMS algebra. *JHEP* **04**, 073 (2020). [arXiv:1909.12827](#)
92. L. Andrianopoli, R. D’Auria, L. Ravera, Hidden gauge structure of supersymmetric free differential algebras. *JHEP* **08**, 095 (2016). [arXiv:1606.07328](#)
93. L. Andrianopoli, R. D’Auria, L. Ravera, More on the hidden symmetries of 11D supergravity. *Phys. Lett. B* **772**, 578–585 (2017). [arXiv:1705.06251](#)
94. L. Ravera, Hidden role of Maxwell superalgebras in the free differential algebras of $D = 4$ and $D = 11$ supergravity. *Eur. Phys. J. C* **78**(3), 211 (2018). [arXiv:1801.08860](#)
95. L. Ravera, AdS Carroll Chern–Simons supergravity in $2 + 1$ dimensions and its flat limit. *Phys. Lett. B* **795**, 331–338 (2019). [arXiv:1905.00766](#)
96. F. Ali, L. Ravera, \mathcal{N} -extended Chern–Simons Carrollian supergravities in $2 + 1$ spacetime dimensions. *JHEP* **02**, 128 (2020). [arXiv:1912.04172](#)
97. R. Caroca, I. Kondrashuk, N. Merino, F. Nadal, Bianchi spaces and their three-dimensional isometries as S -expansions of two-dimensional isometries. *J. Phys. A* **46**, 225201 (2013). [arXiv:1104.3541](#)
98. L. Andrianopoli, N. Merino, F. Nadal, M. Trigiante, General properties of the expansion methods of Lie algebras. *J. Phys. A* **46**, 365204 (2013). [arXiv:1308.4832](#)
99. M. Artebani, R. Caroca, M. Ipinza, D. Peñañiel, P. Salgado, Geometrical aspects of the Lie algebra S -expansion procedure. *J. Math. Phys.* **57**(2), 023516 (2016). [arXiv:1602.04525](#)