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# 3D non-linear MHD simulation of the MHD response and density increase as a result of Shattered Pellet Injection

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**Abstract.** The MHD response and the penetration of a deuterium Shattered Pellet into a JET plasma is investigated via the non-linear reduced MHD code JOEAK with the Neutral Gas Shielding (NGS) ablation model. The dominant MHD destabilizing mechanism by the injection is identified as the local helical cooling at each rational surface, as opposed to the global current profile contraction. Thus the injected fragments destabilize each rational surface as they pass through them. The injection penetration is found to be much better compared to MGI, with the convective transport caused by core MHD instabilities (*e.g.* 1/1 kink) contributing significantly to the core penetration. Moreover, the injection with realistic JET SPI system configurations is simulated in order to provide some insights into future operations, and the impact on the total assimilation and penetration depth of varying injection parameters such as the injection velocity or fineness of shattering is assessed. Further, the effect of changing the target equilibrium temperature or  $q$  profile on the assimilation and penetration is also investigated. Such analysis will form the basis of further investigation into a desirable configuration for the future SPI system in ITER.

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## 1. Introduction

Shattered Pellet Injection (SPI) is the baseline concept for the ITER disruption mitigation system, the aim of which is to deplete the large thermal and magnetic energy stored within the plasma homogeneously by radiative losses so as to prevent localized energy deposition on the device (*e.g.* localized heat flux to plasma facing components by plasma deposition or runaway electrons beam strike). Specifically, the ITER SPI system has the capability to inject up to  $10^{25}$  atoms within several 10 milliseconds [1]. For the thermal quench (TQ) mitigation, the objective is to deplete the thermal energy by radiation as well as modify the conductive heat flux by dilution, thus mitigate the heat flux to the plasma facing components (PFC), and also to raise the core electron density to be high enough to prevent the hot-tail runaway seeds generation. For the current quench (CQ) mitigation, the objective is to reduce the heat flux through the plasma halo and to reduce electro-magnetic loads, both through appropriate levels of the radiated power. Should a runaway beam form during the current quench, despite the effort to suppress the seed generation, the ITER SPI system can also inject large quantities of argon to dissipate the runaway energy through collisions and line radiation [2]. TQ mitigation will largely determine also the CQ properties, namely the electron temperature and density, and thus the efficiency of CQ mitigation [3].

To achieve both objectives, it is desirable to deliver the material right into the plasma core, since this would result in both a more uniform radiative heat flux to the wall, and a higher core electron density to prevent runaway electron formation. This injection penetration in turn is related to the magneto-hydrodynamic (MHD) response from the plasma, as the MHD modes can play significant roles in the inward transport via the plasma convection and the destruction of flux surfaces. Thus it is of great interest to ensure high experimental availability of the ITER tokamak for experiments to acquire a better understanding about the MHD dynamics and associated injection penetration during SPI. This investigation will also form the basis of future self-consistent prediction of the heat flux onto the plasma facing components in a SPI-mitigated disruption.

In this paper, the aforementioned MHD perturbation and density increase are studied by modelling a deuterium SPI into JET target plasmas. JET will be equipped in 2018 with a SPI system that will serve as an important demonstration and extrapolation tool for the future ITER SPI design. The system of interest is described by the reduced non-linear MHD equations combined with a diffusive neutral species, solved by the 3D code JOREK [4, 5]. The ablation of the fragments is modelled by the Neutral Gas Shielding (NGS) model [6, 7]. To better demonstrate the principle of the MHD destabilization mechanism caused by the SPI, simulations with equatorial injection will be carried out first. The dominant mechanism will be shown to be the local helical cooling, thus helical current perturbation, at each low order rational resonant surface. Later on, the realistic JET SPI configuration with injection from an upper vertical port will be used to provide insight into the upcoming JET experiments, as well as to demonstrate the impact of various injection parameters on the assimilation and

penetration of the injection.

The rest of the paper will be arranged as follows. In Section 2, our system of interest will be described and the governing reduced MHD equations will be introduced, as well as the NGS model describing the ablation process. In Section 3, the MHD modes excited by the injection will be investigated. With the understanding gained regarding the MHD instabilities, we proceed to compare the difference between SPI and massive gas injection (MGI) behaviors in Section 4. We then explore the impact of varying injection parameters with the real JET SPI configuration in Section 5. Discussion and conclusion regarding the MHD behavior and the implications for future SPI operation will be presented in Section 6.

## 2. The system of interest

In this section, we introduce our governing equations and assumptions as well as the standard target equilibrium and the injection configurations.

### 2.1. The governing equations and the assumptions

We model the system by considering the reduced MHD equations combined with diffusive neutral species [4]. In the tokamak coordinates  $(R, Z, \phi)$ , the magnetic field and velocity field can be expressed as follows

$$\mathbf{B} = F_0 \nabla \phi + \nabla \psi \times \nabla \phi, \quad (1)$$

$$\mathbf{v} = v_{\parallel} \mathbf{B} - R^2 \nabla u \times \nabla \phi. \quad (2)$$

Here,  $F_0/R$  is the toroidal magnetic field and  $F_0$  is approximately seen as constant in our study, while  $\psi$  is the poloidal magnetic flux. Further,  $u$  is the flow potential for the  $\mathbf{E} \times \mathbf{B}$  flow,  $v_{\parallel}$  is the parallel velocity scaled by the magnetic field. The governing equations are then:

$$\frac{\partial \psi}{\partial t} = \eta (T_e) \Delta^* \psi - R \{u, \psi\} - F_0 \frac{\partial u}{\partial \phi}, \quad (3)$$

$$j = \Delta^* \psi, \quad j_{\phi} = -j/R, \quad (4)$$

$$\begin{aligned} R \nabla \cdot \left( R^2 \rho \nabla_{pol} \frac{\partial u}{\partial t} \right) &= \frac{1}{2} \{ R^2 |\nabla_{pol} u|^2, R^2 \rho \} + \{ R^4 \rho \omega, u \} + \{ \psi, j \} \\ &\quad - \frac{F_0}{R} \frac{\partial j}{\partial \phi} + \{ \rho T, R^2 \} + R \mu (T_e) \nabla^2 \omega \\ &\quad - \nabla \cdot \left[ \left( \rho \rho_n S_{ion} (T_e) - \rho^2 \alpha_{rec} (T_e) \right) R^2 \nabla_{pol} u \right], \end{aligned} \quad (5)$$

$$\omega = \frac{1}{R} \frac{\partial}{\partial R} \left( R \frac{\partial u}{\partial R} \right) + \frac{\partial^2 u}{\partial Z^2}, \quad (6)$$

$$\begin{aligned} \frac{\partial \rho}{\partial t} &= -\nabla \cdot (\rho \mathbf{v}) + \nabla \cdot \left( D_{\perp} \nabla_{\perp} \rho + D_{\parallel} \nabla_{\parallel} \rho \right) \\ &\quad + \rho \rho_n S_{ion} (T_e) - \rho^2 \alpha_{rec} (T_e), \end{aligned} \quad (7)$$

$$\begin{aligned} \frac{\partial(\rho T)}{\partial t} = & -\mathbf{v} \cdot \nabla(\rho T) - \gamma \rho T \nabla \cdot \mathbf{v} + \nabla \cdot (\kappa_{\perp} \nabla_{\perp} T + \kappa_{\parallel} \nabla_{\parallel} T) \\ & + \frac{2}{3R^2} \eta(T_e) j^2 - \xi_{ion} \rho \rho_n S_{ion}(T_e) - \rho \rho_n P_L(T_e) - \rho^2 P_B(T_e), \end{aligned} \quad (8)$$

$$\begin{aligned} \rho B^2 \frac{\partial v_{\parallel}}{\partial t} = & -\rho \frac{F_0}{2R^2} \frac{\partial}{\partial \phi} (B^2 v_{\parallel}^2) - \frac{\rho}{2R} \{B^2 v_{\parallel}^2, \psi\} - \frac{F_0}{R^2} \frac{\partial(\rho T)}{\partial \phi} + \frac{1}{R} \{\psi, \rho T\} \\ & + B^2 \mu_{\parallel}(T_e) \nabla_{pol}^2 v_{\parallel} + (\rho^2 \alpha_{rec}(T_e) - \rho \rho_n S_{ion}(T_e)) B^2 v_{\parallel}, \end{aligned} \quad (9)$$

$$\frac{\partial \rho_n}{\partial t} = \nabla \cdot (\mathbf{D}_n \cdot \nabla \rho_n) - \rho \rho_n S_{ion}(T_e) + \rho^2 \alpha_{rec}(T_e) + S_n. \quad (10)$$

In the above equations, Eq. (3) is the induction equation, Eq. (4) is the result of Ampère's law with the permeability absorbed into the current density, Eq. (5) is the vorticity equation, while Eq. (6) is the definition of vorticity. Moreover, Eq. (7) is the continuity equation, Eq. (8) is the pressure equation, Eq. (9) is the parallel momentum equation, and finally Eq. (10) is the diffusive neutral species density equation. Here, we have defined  $T \equiv T_e + T_i$  and  $T_e = T_i$ , we also used  $\Delta^* \psi \equiv R^2 \nabla \cdot (R^{-2} \nabla \psi)$ , and the dissipative coefficients  $\eta \propto T_e^{-3/2}$  and  $\mu \propto T_e^{-3/2}$  are the resistivity and viscosity. Here, due to numerical reasons, we have used an artificially large resistivity which is ten times larger than the Spitzer resistivity. Furthermore  $\kappa_{\perp}$  and  $\kappa_{\parallel} \propto T_e^{5/2}$  are the perpendicular heat conductivity and parallel Braginskii heat conductivity respectively [8]. The parallel and perpendicular plasma diffusion coefficients and the neutral diffusion coefficient are  $D_{\parallel}$ ,  $D_{\perp}$  and  $\mathbf{D}_n$  respectively, but we used  $D_{\parallel} = 0 m^2/s$  in our study, so that the parallel density relaxation is purely carried out by convective flows as the strong convective flow dominate over the diffusion process. As for the coefficients governing the interaction between the plasma and neutrals,  $S_{ion}(T)$  is the ionization rate and  $\alpha_{rec}(T)$  is the recombination coefficient, the detailed form of which are described in Ref. [5]. Further,  $\xi_{ion}$  is the normalized deuterium ionization energy,  $P_L(T_e)$  is the neutral line radiation coefficient and  $P_B(T_e)$  the bremsstrahlung radiation [5]. We assume that the newly ionized deuterium thermalize immediately so that the plasma always remains Maxwellian. The  $\{f, g\}$  in the above equations denotes the Poisson bracket with  $\{f, g\} \equiv R(\nabla f \times \nabla g) \cdot \nabla \phi$ . Thus Eq. (3)-Eq. (10) form our governing equations.

To close the equations, we still have to specify the neutral source term  $S_n$  caused by the ablation of fragments. We are not concerned with the shattering process itself and will treat the fragments as they are already generated. To this end, we consider the strongly shielded NGS model in a Maxwellian plasma [6, 7]. The principle of this model is simple, that is, we consider a given heat flux coming down along the field line towards the ablating fragment, the ablation rate of the fragment must be such that it maintains a certain line integrated neutral density along the field line to deplete the incoming heat flux so that the actual flux arriving at the fragment surface is negligible [7]. Hence for a given background electron temperature  $T_e$  and density  $n_e$ , the ablation rate for a spherical deuterium fragment with radius  $r_p$  is

$$\partial_t N [s^{-1}] = 4.12 \times 10^{16} r_p^{4/3} [m] n_e^{1/3} [m^{-3}] T_e^{1.64} [eV]. \quad (11)$$

Here,  $\partial_t N$  is the number of ablated atoms per second. In our model, we deposit those ablated neutrals around the fragment with the following gaussian shape

$$S_n \propto \exp\left(-\frac{(R - R_f)^2 + (Z - Z_f)^2}{\Delta r_{NG}^2}\right) \times \exp\left(-\left(\frac{\phi - \phi_f}{\Delta \phi_{NG}}\right)^2\right). \quad (12)$$

Here,  $R_f$ ,  $Z_f$  and  $\phi_f$  are the spatial position of fragments, while we choose the neutral cloud parameter  $\Delta r_{NG} = 2cm$  and  $\Delta \phi_{NG} = 0.6$ . The evolution of fragment size is then governed by the conservation of mass, that is

$$n_p 4\pi r_p^2 \frac{\partial}{\partial t} r_p = \frac{\partial}{\partial t} N, \quad (13)$$

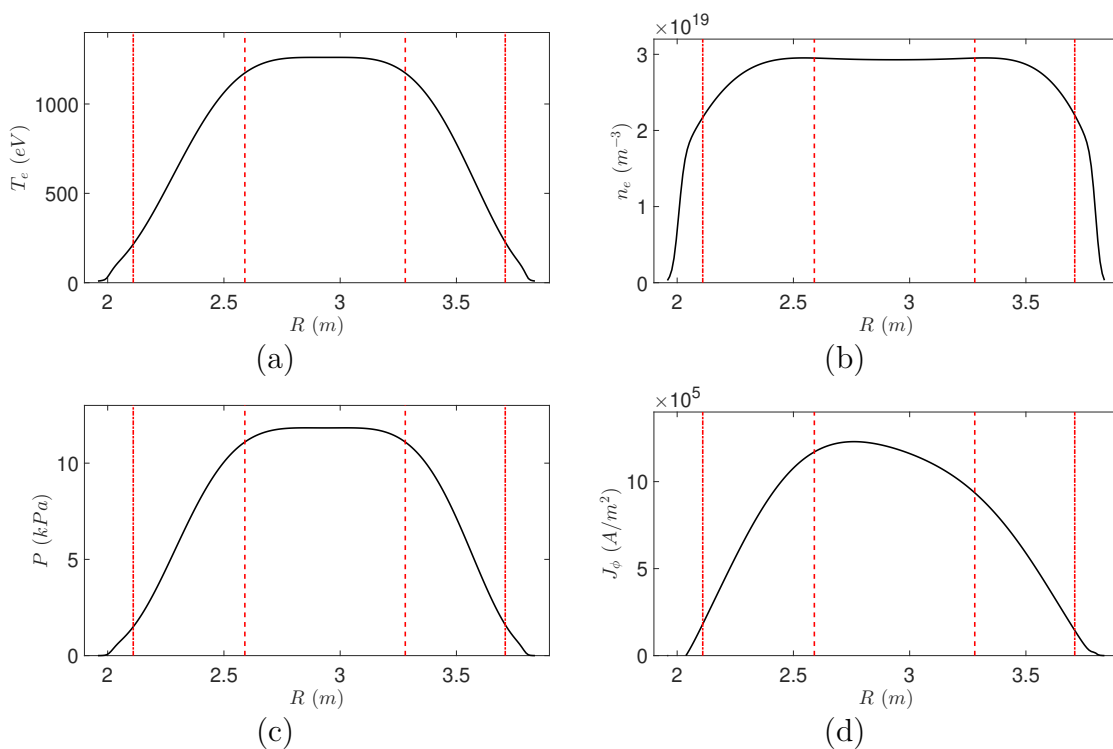
with  $n_p \simeq 5.958 \times 10^{28}/m^3$  being the atom density of the deuterium fragments. Furthermore, the fragments are treated as travelling through the plasma without drag. This is a reasonable assumption since the density difference between the fragments and the plasma is very large even after the injection:  $n_p/n_e > \mathcal{O}(10^6)$ , so that for the drag force to manifest itself within  $1ms$ , the fragment radius must be approximately smaller than  $10^{-6}m$ , by the time of which it becomes irrelevant to the plasma conditions and time evolution that we are considering here.

## 2.2. The target equilibrium and the injection configurations

We use the JET pulse No. 86887 as a template for the so called ‘‘standard equilibrium’’, with  $q_0 = 0.935$  and  $q_{95} = 2.9$ . The toroidal magnetic field  $B_t \simeq 2T$ , and the total plasma current is  $I_p \simeq 2MA$ . The plasma is in L-mode before injection, with central electron temperature  $T_e(0) \simeq 1.25keV$ , and central plasma density  $n_e(0) \simeq 2.9 \times 10^{19}/m^3$ . This set of L-mode plasma parameters is meant to represent a low thermal energy plasma before disruption after significant thermal energy loss has occurred [3]. No background impurity radiation is assumed. This particular equilibrium is stable to large scale tearing modes ( $m \geq 2$ ), and numerical observation indicates that the natural sawtooth period is long compared to the timescale we are concerned with here. Hence we consider the 1/1 resistive kink to be practically stable in the absence of injection in this study.

Midplane cuts of the electron temperature profile, the electron density profile, the pressure profile and the toroidal current density profile are shown in Fig. 1. The  $n_e$  and  $T_e$  profiles are generated by using the Thomson scattering data and the equilibrium is constructed by EFIT data as described in Ref. [5]. The chained and dashed red lines represent the major radius at the midplane for the  $q = 2$  and  $q = 1$  surfaces, respectively. As mentioned in Section 1, we consider an equatorial injection to demonstrate the principal mechanism of MHD destabilization by SPI, and a realistic JET SPI configuration to demonstrate the impact of the injection parameters on the injection penetration and assimilation. Sketches of the two injection configurations are shown in Fig. 2, with the red lines outlining the spread cone of the trajectories of the fragments.

As a further note, the grid size used in our investigation is 101 in  $\psi$  direction and 128 in  $\theta$  direction for the closed field line region, while the scrape-off layer has a grid

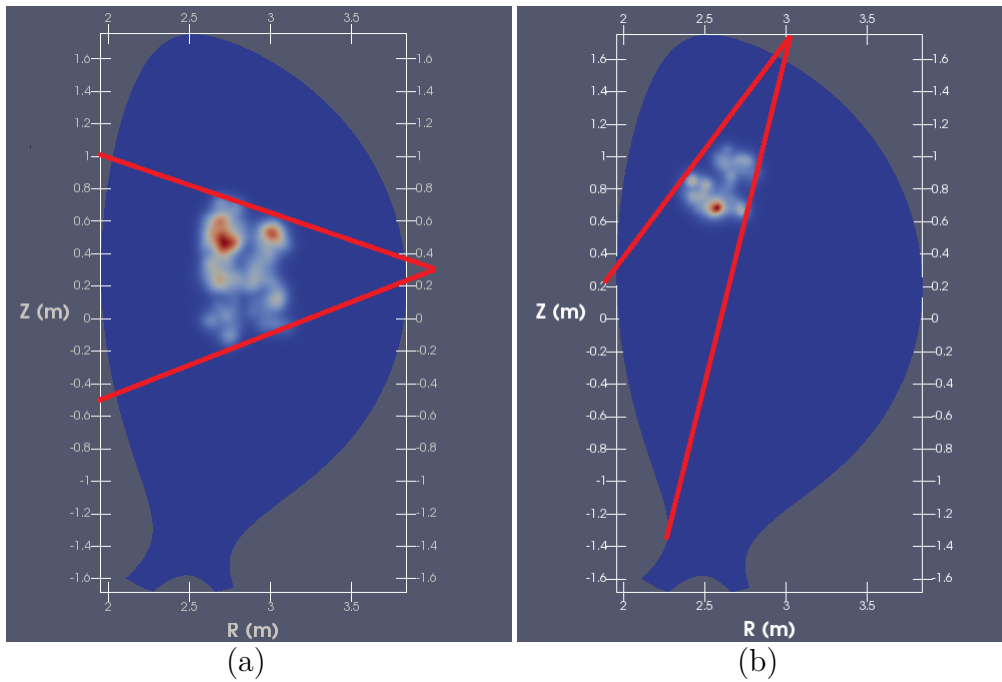


**Figure 1.** Profiles at the midplane of the standard target equilibrium for (a) the electron temperature profile, (b) the electron density profile, (c) the pressure profile and (d) the toroidal current density profile. The red chained and dashed lines denote the major radius of the  $q = 2$  and  $q = 1$  surfaces at the midplane, respectively.

size of 4 only. This small number of grid elements covering the open field line region is justified by our emphasis on the core MHD activity.

The reference parameters of the above two injection configurations are as follows. For the equatorial injection, the injection is carried out from the low field side (LFS) pointing purely along major radius as shown in Fig. 2(a). The total injection amount is  $5 \times 10^{22}$  particles, equally shattered into 100 fragments each with radius  $1.26\text{mm}$ . The injection speed is  $500 \pm 100\text{m/s}$  with a flat distribution function and a spread vertex angle 40 degrees. It should be noted that the spread angle is unrealistically large in this equatorial case, but numerical investigations with different spread angles and the speed spread shows that, for the values chosen, which are reasonable evaluations of those to be achieved in experiments, they make little difference to the MHD destabilization and consequentially the injection penetration. As for the JET-like injection, the injection is carried from upper LFS and pointing downwards [11] as shown in Fig. 2(b), and the reference injection direction (the axis of the velocity spreading cone) is within the  $(R, Z)$  plane. The total injection quantity is set to be  $3.6 \times 10^{22}$  deuterium atoms, corresponding to the medium sized injection as per the JET SPI design [11, 12]. Moreover, the injected quantity is shattered into 100 fragments with the following size distribution [13]

$$P(r_p) = \frac{r_p K_0(\kappa_p r_p)}{I}, \quad I \equiv \int_0^\infty r_p K_0(\kappa_p r_p) dr = \kappa_p^{-2}, \quad (14)$$



**Figure 2.** The SPI configurations for (a) the equatorial injection and (b) the JET-like injection.

where  $K_0$  is the modified Bessel function of the second kind, and  $\kappa_p$  is the inverse of the characteristic fragment size which is determined by requiring

$$n_p N_p \int_0^\infty P(r_p) \frac{4}{3} \pi r_p^3 dr_p = N, \quad (15)$$

with  $N_p$  being the total number of fragments,  $N$  being the total injected particles. Thus

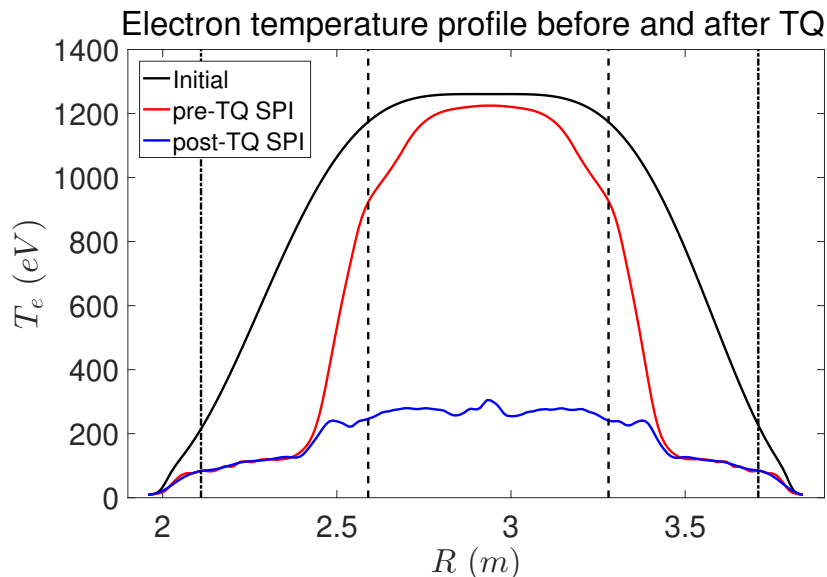
$$\kappa_p = \left( \frac{N}{6\pi^2 n_p N_p} \right)^{-1/3}. \quad (16)$$

Moreover, the fragment velocity is set to be  $200 \pm 40 m/s$ , with a vertex angle of 20 degrees. The reference speed and vertex angle are chosen according to JET SPI system design [11, 12], while the distribution of velocity is chosen ad hoc in want of deeper theoretical understanding or experimental observation at present.

Later on in Section 5, we will deviate from the reference injection parameters for the JET-like configuration to see the impact of varying injection quantities, injection speed and shattering fineness on the MHD activity and the injection penetration. Moreover, equilibria with a different  $q$  profile and different electron temperature are also investigated to see the impact on the assimilation rate.

### 3. MHD response caused by the SPI

Macroscopic current driven modes are the major players in the post-injection MHD response due to their global mode structure. Those large scale modes are destabilized by the current density displacement as a result of the drastic electron temperature

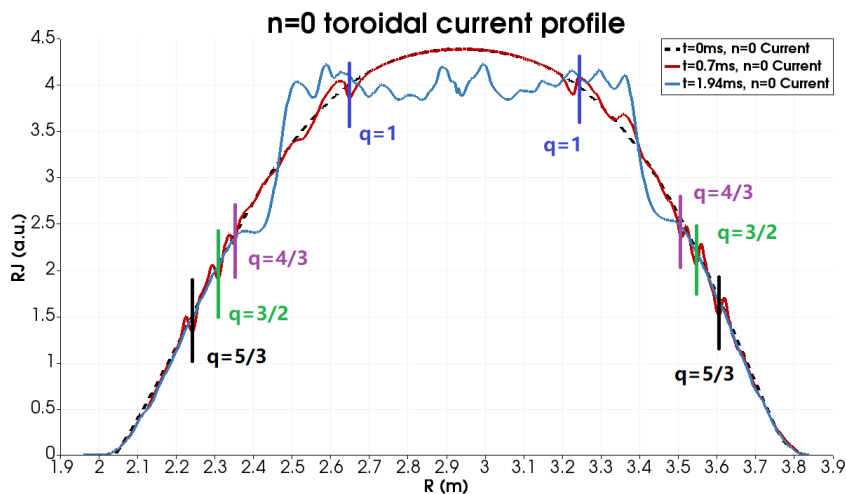


**Figure 3.** The mid-plane electron temperature profile at the beginning of the simulation, just before the thermal quench and just after the thermal quench. The black chained and dashed lines represent the  $q = 2$  and  $q = 1$  surfaces respectively.

change after SPI or MGI, which occurs on the local resistive timescale  $\tau_\eta \sim l^2/\eta$ , with  $l$  being the length scale of said displacement and  $\eta \propto T_e^{-3/2}$  is the Spitzer resistivity. The community has long established that the propagation of the cooling front along the minor radius, and consequentially the global current contraction contribute greatly to large scale MHD excitations [14, 15], though numerical investigation of deuterium MGI also pointed out the importance of local helical cooling to the growth of corresponding helical modes [4]. Indeed, as will be found in this paper, as long as the fragment travelling timescale is smaller or comparable with the current contraction timescale, the local current perturbation will dominate over the global current contraction as the main MHD destabilizing mechanism during SPI. This is due to the fact that the local perturbation length scale is much smaller than that of the global contraction.

Due to the mild radiation coefficient from hydrogen isotopes [16], the dominant cooling mechanism in our investigation is the plasma dilution caused by the fragments ablation. The post-injection electron temperature is still on the order of  $100\text{eV}$  during the pre-TQ phase, as can be seen in Fig.3 where the mid-plane cut of the electron temperature profile evolution for the equatorial SPI is shown. From this it can be estimated that the timescale for the global current profile to have a  $10\text{cm}$  contraction along the minor radius is about  $10\text{ms}$ . This is longer than the travelling timescale of the fragments assuming a velocity of  $200\text{m/s}$ , thus the global current contraction is unlikely to have a major contribution to the MHD excitation.

The above statement is supported by the  $n = 0$  current density profile evolution after the equatorial SPI, as is shown in Fig.4. It can be seen that the mean current density profile does not exhibit strong contraction even just before the onset of the thermal quench, and significant profile variation only occurs after the core current is

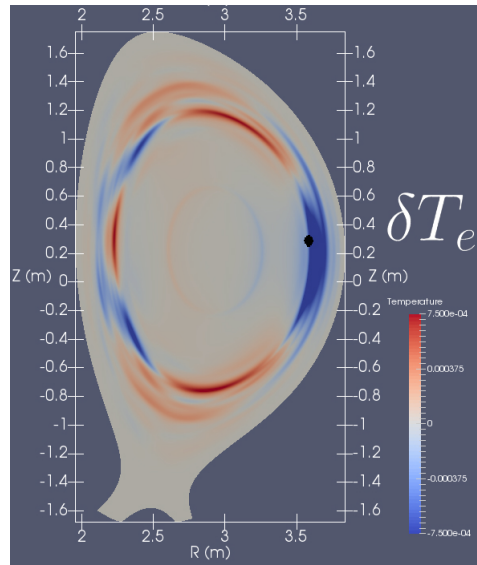


**Figure 4.** The mid-plane current density profile multiplied by the major radius  $R$  at the beginning of the simulation, just before the TQ ( $t = 0.7ms$ ) and about  $1ms$  after the TQ. The global current contraction is very limited before the TQ, while after the TQ the core current distribution is flattened by the hyper-resistivity.

flattened by non-linear  $\mathbf{v} \times \mathbf{B}$  induced hyper-resistivity after the thermal quench [9, 10]. It should be noted, however, that “jagged” features have developed along the minor radius which correspond to the positions of low order rational surfaces, as indicated on Fig. 4 by vertical lines. Those are the result of the local helical cooling as will be shown later in this section.

The aforementioned local helical cooling is essentially caused by the geometry of the magnetic field. As the fragments enter the plasma and begin to ablate, they induce a rapid cooling along field lines by parallel heat conduction. The typical timescale of such cooling can be estimated by considering the Braginskii heat conduction [8] of a plasma with  $300eV$  electron temperature and  $10^{20}/m^3$  density, and a connection length  $L_c \equiv 2\pi Rq \simeq 50m$ . The resulting parallel cooling time is then  $\tau_{\parallel} \sim 10^{-5}s$ . On irrational surfaces, this will ultimately result in a more or less uniform cooling of the whole flux surface, as the field lines will not connect with themselves. Near rational surfaces, however, field lines connect with themselves after several toroidal turns, and the fragments induce a helical cooling structure instead. This structure will decay on the perpendicular transport timescale, which is about hundreds of microseconds. The resulting  $n \neq 0$  temperature perturbation is shown in Fig. 5, where the dominant  $m = 2$ ,  $m = 5$  and  $m = 3$  components correspond to the  $2/1$ ,  $5/3$  and  $3/2$  helical harmonics, respectively. There is also a faint trace of  $m = 1$  component in the plasma core, which is caused by the  $1/1$  component of plasma displacement. This  $1/1$  component is likely to be the result of mode beating, such as the beating of the  $3/2$  and the  $2/1$  mode.

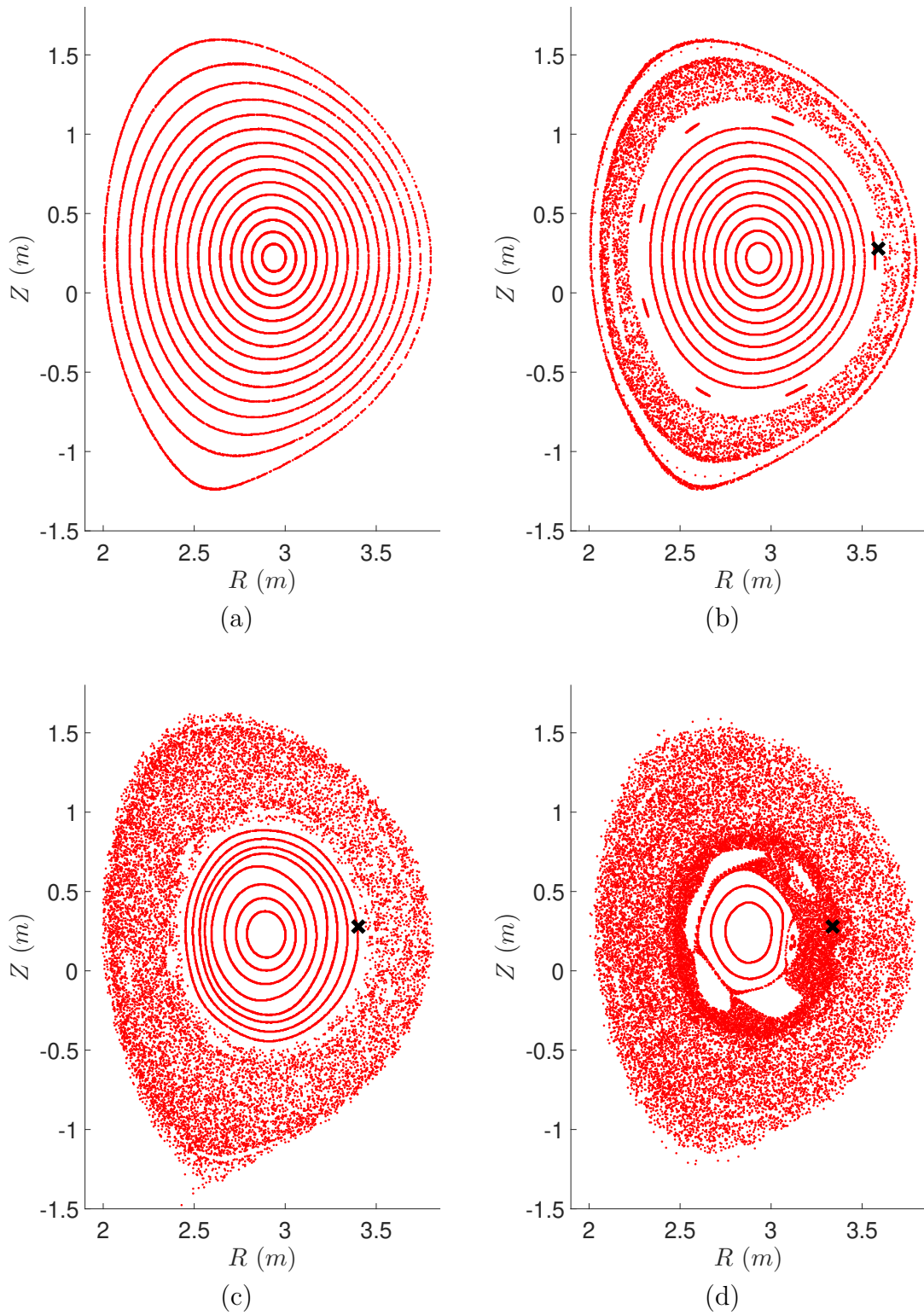
Such a helical cooling structure will induce a corresponding negative helical current perturbation, which is greatly destabilizing for resonant modes. Such a destabilizing effect is due to the above described helical perturbation modifying the local mode structure near the resonant surface in such a way that it lowers  $\psi'_s|_-/\psi_s$  and increases



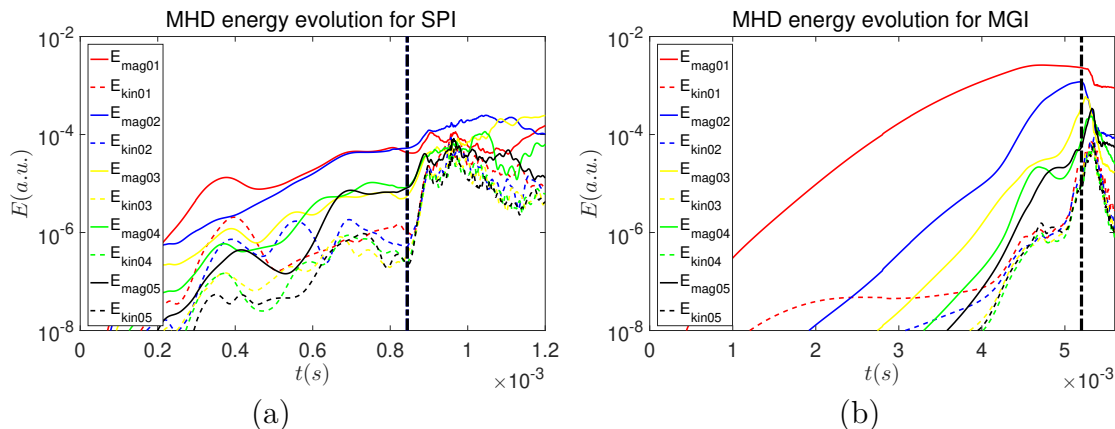
**Figure 5.** The  $n \neq 0$  relative electron temperature perturbation after the injection, the black point represents the position of the fragment cloud. Resonant helical cooling is evident at the corresponding rational surfaces, with the most dominant three components being the 2/1, 5/3 and 3/2, and a faint sign of 1/1 component at the core, which is caused by the plasma displacement.

$\psi'_s|_+/\psi_s$ , resulting in an increase of the stability criterion  $\Delta' \equiv \frac{\psi'_s|_+}{\psi_s|_-}$ , driving the tearing instability [17]. Here,  $\psi'_s|_-$ ,  $\psi'_s|_+$  and  $\psi_s$  are the radial gradient of the perturbed magnetic flux at both sides of the resonant surface and the value of the perturbed flux at the resonant surface, respectively. This mechanism is essentially the same as the cooling island mechanism proposed by White, Gates and others to explain the sudden growth of islands in density limit disruptions [18, 19], where the radiation within islands results in similar helical cooling structures, causing the destabilization of the islands. Furthermore, the additional bootstrap current profile modification within the island as a result of the pressure profile change can also play a role in the destabilization.

This local cooling mechanism implies that the fragments will destabilize successive rational surfaces as they travel across the plasma, generating a broad spectrum of magnetic perturbations. If those surfaces are packed densely enough, the resulting overlapping islands will cause large transport along field lines and thus a significant decrease of plasma confinement. This can be seen from the Poincaré plots of magnetic field lines shown in Fig. 6. The black cross in the figures represents the approximate location of the fragment cloud “vanguard”, although there exists some spread both within the poloidal plane and along the toroidal direction. Nonetheless, it can be seen that islands open up as the fragments pass by, and stochasticity follows the vanguard of the fragment cloud closely as they dive into the plasma core. A similar effect is previously reported for the fuelling/triggering pellet triggering of medium- $n$  modes, although there the triggering mechanism is attributed to the local pressure increase rather than the helical cooling [20]. This difference in mechanism is due to the fact that



**Figure 6.** The Poincaré plot of magnetic field lines at (a)  $t = 0\text{ms}$ , (b)  $t = 0.245\text{ms}$ , (c)  $t = 0.567\text{ms}$ , (d)  $t = 0.669\text{ms}$ . The black cross represents the approximate position of fragments cloud vanguards.

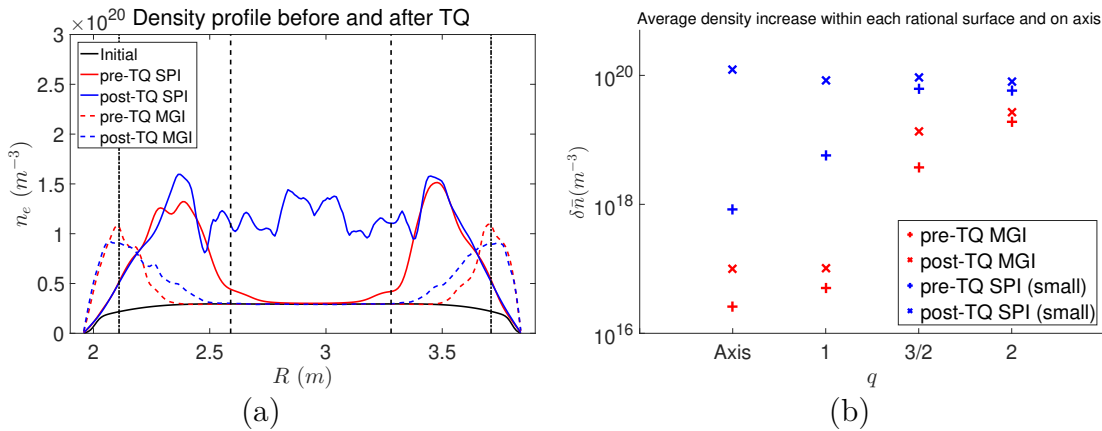


**Figure 7.** The magnetic and kinetic perturbation energy of  $n = 1 - 5$  modes for (a) the SPI and (b) the MGI. The black chained lines in both cases indicate the time of the onset of the TQ.

the fuelling/trigging pellet is simply too small to cool the plasma down drastically, thus the helical cooling effect is minimal. The thermal quench is ultimately triggered when the fragments enter the  $q = 1$  surface as shown in Fig. 6(d) and excite the  $1/1$  kink, which destroys the core confinement completely.

#### 4. MHD modes and injection penetration compared to MGI

With the above understanding of the MHD destabilization mechanism of SPI, we can proceed to investigate the MHD spectrum as a result of the injection. In this section, we compare the SPI result with that of a similar quantity MGI. The MGI configuration is the same as the one described in Ref. [4], with a small injection quantity of  $4.8 \times 10^{21}$  deuterium atoms. As for SPI, we use the equatorial injection configuration, as described in Section 2.2, but the total injection quantity is  $6.25 \times 10^{21}$  atoms. Note that there is still some difference between the total injection quantity of the MGI and SPI cases due to historical reasons, but the injected quantities are similar and thus we expect that the differences between SPI and MGI reported here are due to the different material injection schemes and not the slightly different amounts of material injected. Thus, this is unlikely to alter the following comparison significantly. The magnetic and kinetic perturbation energy of  $n = 1$  to  $n = 5$  harmonics for both cases are shown in Fig. 7. As a reference, the  $n = 0$  mean magnetic energy, which is not shown on the figure, has the order  $\mathcal{O}(1)$ . For the SPI case, the spectrum of MHD perturbations before the thermal quench is broad, as can be seen from Fig. 7(a) where there is hardly one order of magnitude difference between the magnetic perturbation energy of different harmonics. The thermal quench is triggered at  $t = 0.9 \text{ ms}$  when the fragments penetrate into the  $q = 1$  surface. This penetration time is somewhat longer ( $0.9 \text{ ms}$  compared to  $0.7 \text{ ms}$ ) than that is obtained in section III and this is due to the reduced amount considered in these simulations, since the “vanguard” fragments are burnt up before they can reach

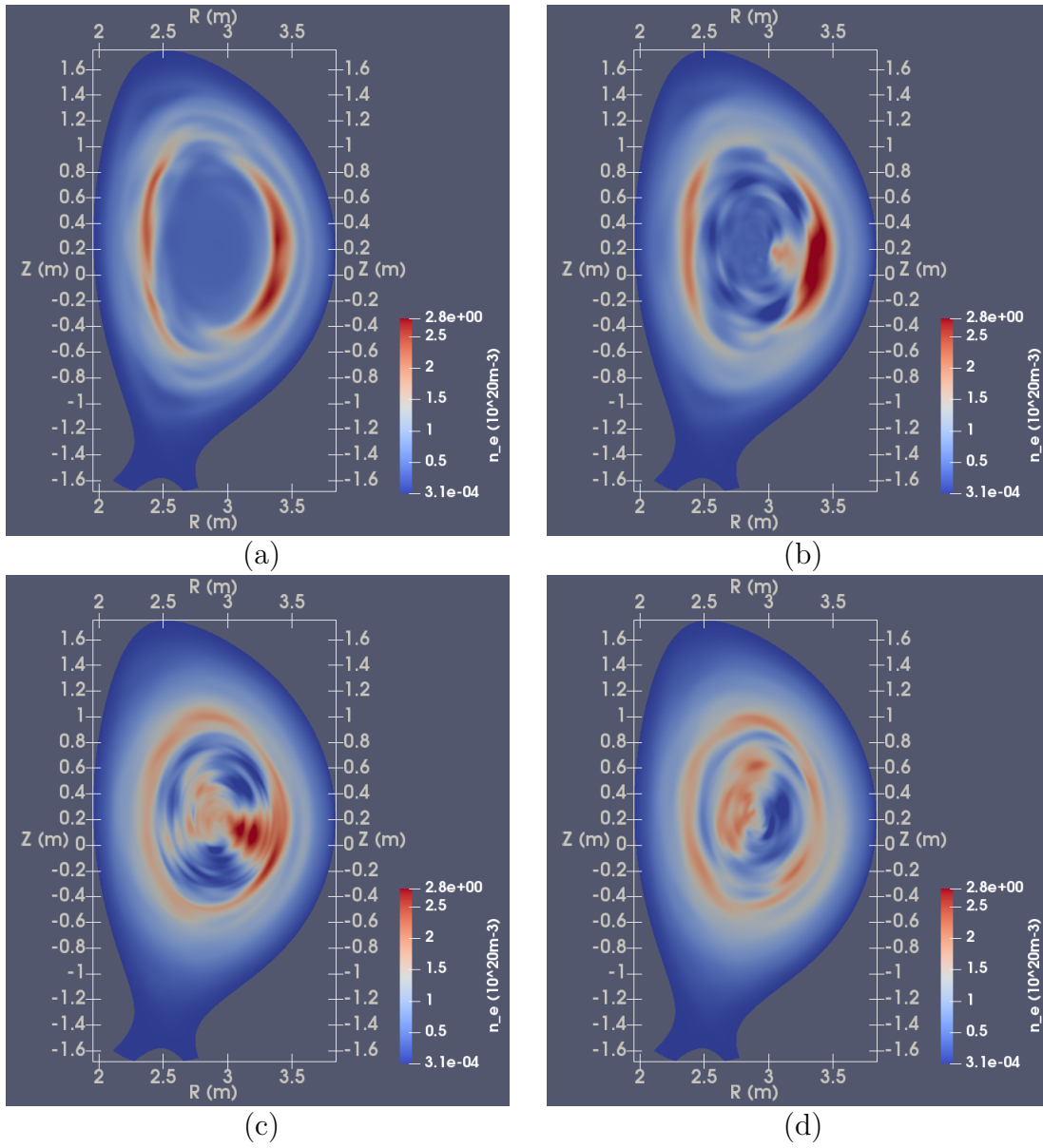


**Figure 8.** (a) The density profile before and after the TQ for both the MGI and the SPI case, the black chained and dashed lines represent the major radial position of the  $q = 2$  and  $q = 1$  surfaces, respectively; (b) The average density increase on the magnetic axis and within each flux surface just before and after the TQ for both the MGI and the SPI case.

the  $q = 1$  surface. The above behavior is in contrast to that of the MGI case as shown in Fig. 7(b) where the 2/1 mode is dominant, and the thermal quench is triggered at  $t = 5.2ms$  after the island growth is large enough to destabilize the 3/2 mode which then overlaps with it, destroying the flux surfaces [4].

This difference in the MHD response is due to the SPI penetrating much deeper before triggering the thermal quench compared to the MGI case, for which the penetration is limited to the  $q = 2$  surface. Thus, in the MGI case only the 2/1 mode is being destabilized, until it grows to a substantial amplitude to nonlinearly destabilize other modes. This difference in penetration can be readily seen by looking at the density profile evolution and the average density increase within each flux surface, as shown in Fig. 8, where the comparison of the density profile, as well as the average density increase before and after the thermal quench is presented for both cases. Here, the thermal quench occurs from  $t = 0.844ms$  to  $t = 1.26ms$  for the SPI case, while for the MGI case, it happens from  $t = 5.37ms$  to  $t = 6.07ms$ . From Fig. 8(a), it can be seen that SPI goes much deeper than MGI, with the latter accumulating at the  $q = 2$  surface. After the TQ, sufficient core mixing occurs for the SPI case, resulting in a substantial rise in core density. The MGI case shows, in contrast, only limited density spreading around the  $q = 2$  surface. From Fig. 8(b), it can be seen more directly that the density increase within the  $q = 3/2$  surface is negligible for the MGI case, while for the SPI case there are drastic increases on the magnetic axis and within the  $q = 1$  surface.

This significantly better core penetration of the SPI case is due to the excited 1/1 kink that convectively transports the density peak near the  $q = 1$  surface into the core region. This is shown in Fig. 9, where the kink motion is seen to drag “density cells” across the field lines into the  $q = 1$  surface in a  $\mathcal{O}(10^{-4}s)$  timescale. Such strong convection is due to the global mode structure of the 1/1 mode within the resonant surface. In the MGI case, however, the injected density accumulates far away from the



**Figure 9.** The density evolution during the thermal quench for the SPI, with the time corresponding to each figure being (a)  $t = 0.844ms$ , (b)  $t = 0.893ms$ , (c)  $t = 1.12ms$  and (d)  $t = 1.26ms$ . The kink motion “drags” the particles accumulated near the  $q = 1$  surface into the plasma core.

core modes at the  $q = 2$  surface, thus the core mixing is limited, and there is only some density spreading near the  $q = 2$  surface itself [4]. Therefore it can be concluded that the penetration of material before the thermal quench is triggered has very important consequences for the redistribution of the injected material in the core plasma due to MHD activity mixing. In this respect SPI offers optimization possibilities that the MGI cannot provide to increase the core plasma density during disruption mitigation thus possibly decreasing the probability of runaway formation following the thermal quench.

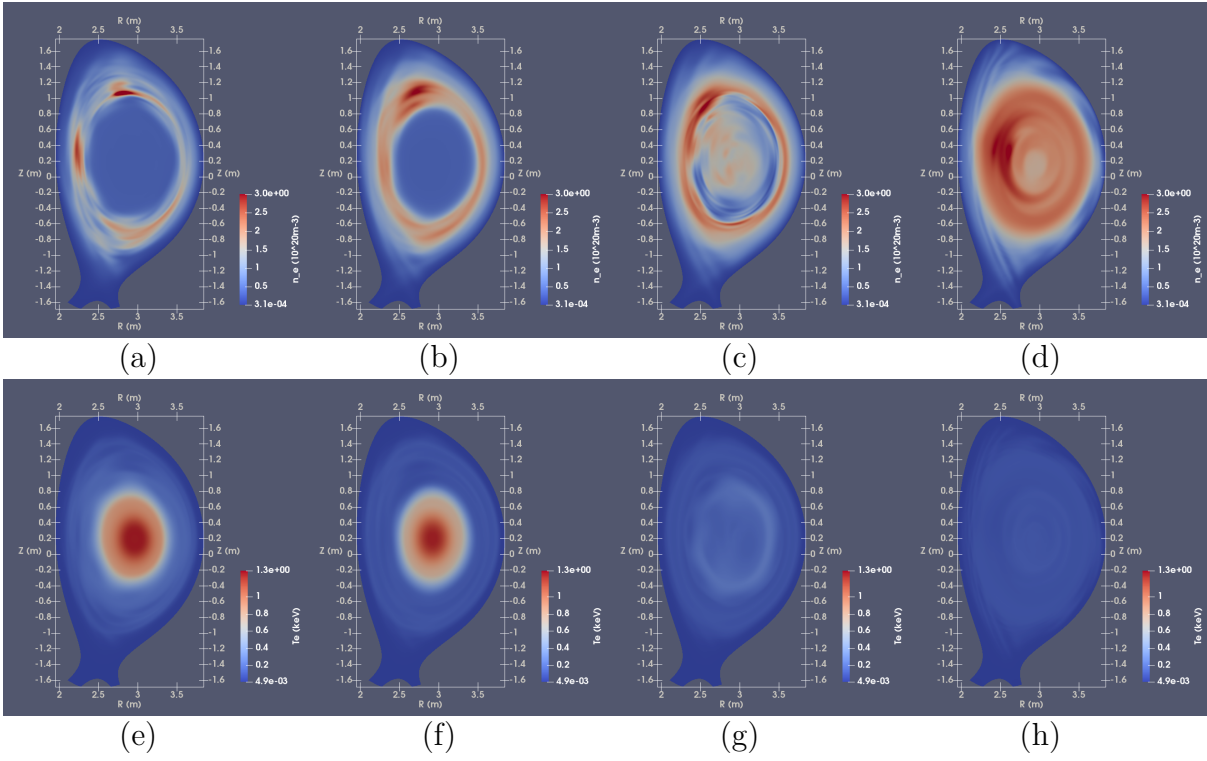
## 5. Injection penetration and assimilation for realistic JET SPI configurations

### 5.1. Density profile evolution for JET-like injection

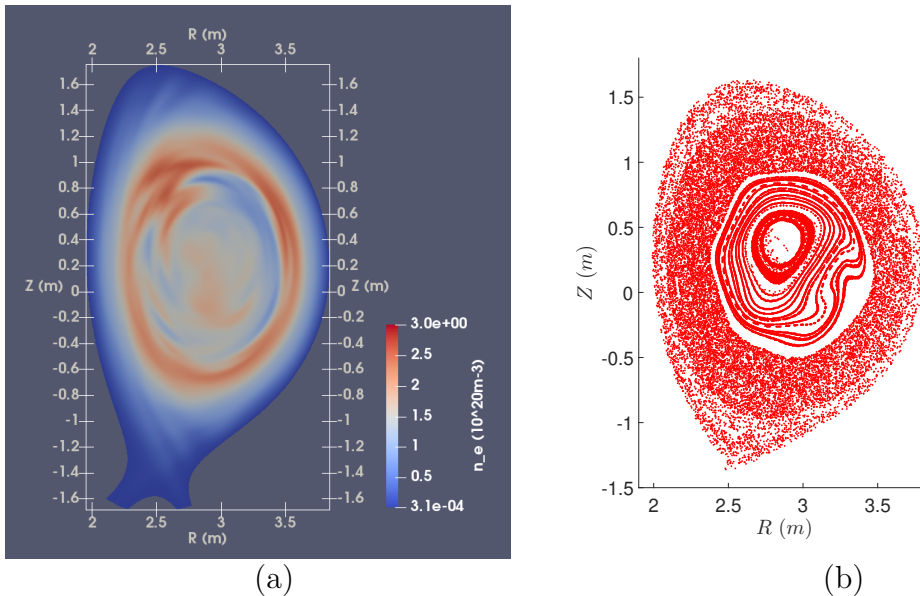
We now move on to investigate the JET-like injection as described in Section 2.2, where the line of injection does not pass through the magnetic axis, but rather only grazes on the  $q = 1$  surface as shown in Fig. 2(b) where the inner major radial position of the surface is at  $R \simeq 2.6m$  on the mid-plane, and the outer one is  $R = 3.28m$ .

The sequence of events for the JET-like injection is somewhat different from that described in the equatorial case, as the fragments are travelling at a much slower speed of  $200 \pm 40m/s$ , as compared with the  $500 \pm 100m/s$  for the equatorial case. This gives the outer modes such as 2/1 and 3/2 more time to grow and interact, the overlapping of which triggers a first stage thermal quench that flattens the electron temperature outside of the  $q = 1$  surface. The poloidal density profile cut at the end of this first stage thermal quench ( $t = 2.56ms$ ) is shown in Fig. 10(a). It can be seen that there is no core penetration at that time and the density profile exhibits an  $m = 3$  asymmetry, corresponding to the fragments' position near the  $q = 3/2$  surface. Later, as the fragments go deeper and reach the  $q = 1$  surface, a second stage thermal quench is triggered and the core temperature is completely flattened. The poloidal density profiles at the beginning ( $t = 3.28ms$ ) and the end ( $t = 3.98ms$ ) of this second thermal quench are shown in Fig. 10(b) and (c), respectively. Similar to the equatorial injection case discussed in Section 4, convective transport of “density cells” into the plasma core has been observed in Fig. 10(c), although it can be seen that the core mixing is not complete yet at the end of the thermal quench. Only milliseconds later at  $t = 6.14ms$  is the core mixing truly complete, as is shown in Fig. 10(d). The corresponding temperature profile of Fig. 10(a)-(d) are shown in Fig. 10(e)-(h).

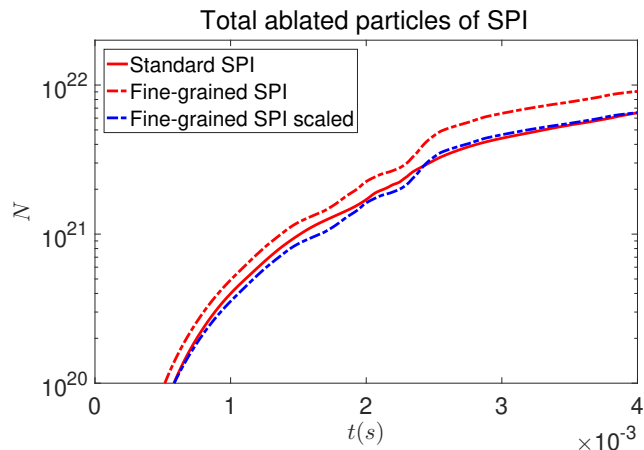
It can be seen that sufficient core density mixing can still occur via the 1/1 kink convection despite the fact that the SPI does not pass through the magnetic axis directly, although the timescale of complete mixing is longer than the timescale of the TQ duration. The fact that the temperature profile is flattened more rapidly than the density profile may have a strong impact on the runaway current formation, as the collision frequency of runaway electrons is proportional to the electron density [21] and localized runaway current filaments may be able to form before sufficient core mixing occurs. This is especially true if the magnetic surfaces recover from the stochastic state before the density mixing is complete, as is shown in Fig. 11, where the core mixing is still ongoing after the thermal quench while the core flux surfaces begin to reform, which is favorable for runaway electron formation. A more detailed study would investigate the Poincaré plot of the runaway trajectories to determine whether or not the seed runaways are well confined in those “hollow density” regions, but this is beyond the scope of this paper and is left for future studies.



**Figure 10.** The density evolution for the JET-like SPI at (a) the end of the first stage TQ ( $t = 2.56 \text{ ms}$ ), (b) the beginning of the second stage TQ ( $t = 3.28 \text{ ms}$ ), (c) the end of the second stage TQ ( $t = 3.98 \text{ ms}$ ) and (d) the completion of core mixing ( $t = 6.14 \text{ ms}$ ). The corresponding temperature profile of (a)-(d) are shown in (e)-(h).



**Figure 11.** (a) The density profile cut and (b) magnetic field Poincaré plot when the magnetic surfaces begin to recover at time  $t = 4.33 \text{ ms}$ , after the TQ.



**Figure 12.** The comparison of the total assimilated particles with different shattering fineness.

### 5.2. The impact of shattering fineness on the assimilation rate

The fineness of the shattering process is found to have an impact on the total assimilation rate of the SPI. That is, for a given total injection amount, the value of  $\kappa_p$  in Eq. (14), thus equivalently the total number of fragments, will influence the total assimilation rate. Here, we are considering the case when the fragments are not fully ablated (otherwise it is meaningless to compare the assimilation), and they are still coarse enough to be considered drag-less through the plasma.

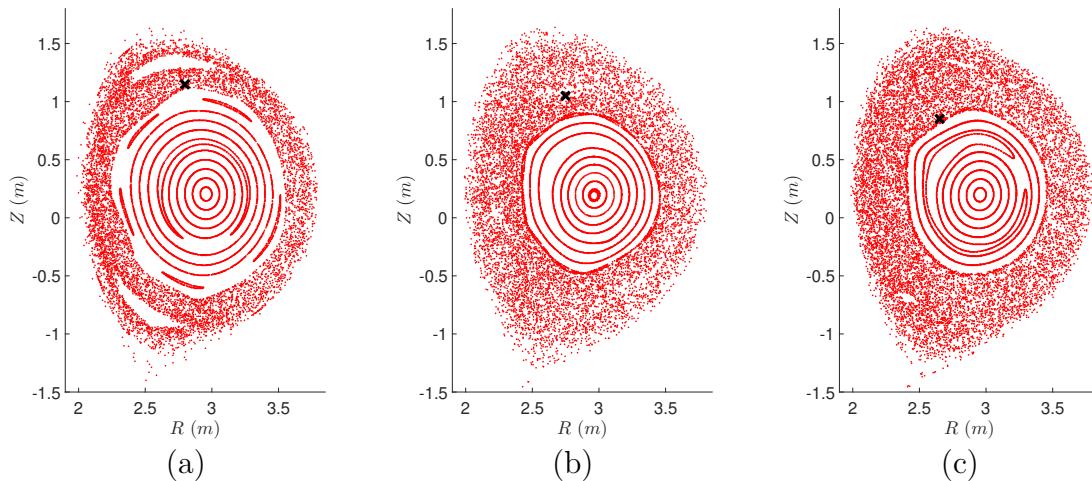
Naively, in a constant temperature plasma, the total ablation rate goes with  $\partial_t N_{tot} \propto \kappa_p^{-4/3} N_p \propto \kappa_p^{5/3}$ , according to Eq. (11), Eq. (14) and Eq. (16) with a given total injection of  $N$  atoms. However, since the plasma is being cooled down by the ablation, this simple power law is expected to be compromised since faster cooling will slow down further ablation.

For deuterium injection, one can model this reduced assimilation by realizing that the dilution of temperature dominates over the radiation [16], ionization (13.6eV) and sublimation loss (less than  $10^{-2}eV$ ) [22] as the main cooling mechanism for the scenarios we are concerned with here. As a consequence the total thermal energy loss due to ablation is negligible and  $n_e T_e \sim constant$ . To understand this, we have to consider the total ablation rate for two cases 1 and 2, with different shattering fineness characterized by  $\kappa_{p1}$  and  $\kappa_{p2}$ , but otherwise exactly the same injection parameters. We now assume *a priori* that there exists a power law with the form  $\partial_t N_{tot1} / \partial_t N_{tot2} = (\kappa_{p1} / \kappa_{p2})^\alpha$ , where  $\alpha$  is some power to be determined. We further realize that the electron density is strongly dominated by the injection, so that, for a given time  $t$ , we have  $n_{e1}(t) / n_{e2}(t) \simeq \partial_t N_{tot1} / \partial_t N_{tot2} = (\kappa_{p1} / \kappa_{p2})^\alpha$ . From the NGS model, we have

$$\partial_t N_{tot} \propto \kappa_p^{5/3} n_e^{1/3} T_e^{1.64} \simeq \kappa_p^{5/3} n_e^{-4/3} (n_e T_e)^{5/3}. \quad (17)$$

Then we naturally have

$$\kappa_{p1}^{5/3} n_{e1}^{-4/3} \simeq (\kappa_{p1} / \kappa_{p2})^\alpha \kappa_{p2}^{5/3} n_{e2}^{-4/3}. \quad (18)$$



**Figure 13.** The Poincaré plot of magnetic field lines at (a)  $t = 2.16ms$ , (b)  $t = 2.56ms$ , (c)  $t = 3.28ms$ . The black cross represents the approximate position of fragments cloud vanguards.

Which in turn indicates

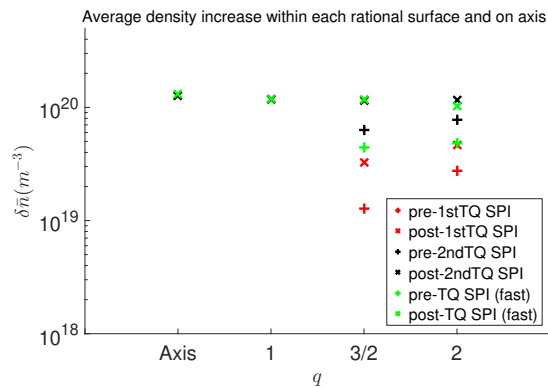
$$\left(\frac{\kappa_{p1}}{\kappa_{p2}}\right)^{5/3} \simeq \left(\frac{\kappa_{p1}}{\kappa_{p2}}\right)^{7\alpha/3}. \quad (19)$$

Hence we have  $\alpha \simeq 5/7$ . This suggests that the finer the shattering, the better the total assimilation of the SPI, as we recall that  $\kappa_p^{-1}$  is the characteristic fragment size.

This is in good agreement with the numerical observation, as is shown in Fig. 12, where the standard JET-like injection assimilation is compared with a “fine-grained” JET-like SPI. The latter is shattered into 400 fragments as opposed to the 100 fragments of the former, so that the standard case has a  $\kappa_p \simeq 2140m^{-1}$ , while the “fine-grained” case has  $\kappa_p \simeq 3397m^{-1}$ . In the figure, the red solid line and chained line represents the standard and “fine-grained” total assimilation, while the blue chained line represents the “fine-grained” scaled with the  $(\kappa_{p1}/\kappa_{p2})^{5/7}$  scaling. The good agreement confirms our modelling result of a  $5/7$  power law, indicating that the shattering fineness does have an impact on the total assimilation, though not as strong as one would expect from simply looking at the NGS model. In the very fine fragments limit, however, the drag of the plasma cannot be ignored anymore and the fine fragments will be stopped at the edge of the plasma, effectively reducing the assimilation. Also the injection penetration is reduced in the very fine fragments limit, and it would resemble that of a MGI as shown in Fig. 8(a). Hence, the fragment size has to be chosen to guarantee both high assimilation and sufficient rise in core density.

### 5.3. Impact of the injection velocity of SPI on the MHD response

The injection velocity also plays a significant role on the MHD response. Varying the injection velocity will effectively vary the timescale of fragments passing through each rational surface. Therefore, in general, a slower injection will mean more time for the

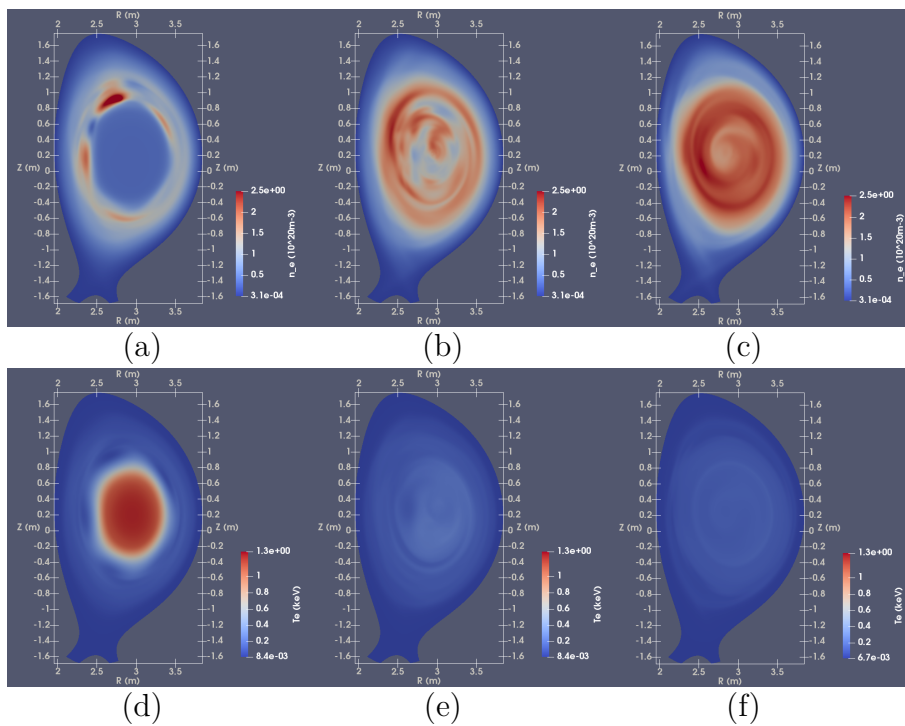


**Figure 14.** The comparison between injection penetration for the standard and fast JET-like SPI. The red signs correspond to before and after the 1st TQ of the standard case, the black ones represent those of the 2nd TQ, while the blue ones represent that of the fast case. There is little density rise in plasma core before the final TQ so the density rise within the  $q = 3/2$  surface is not shown for those cases.

modes to grow and interact with each other. This may cause the stochastic region due to island overlapping to propagate faster than the fragments if the mode nonlinear growth rate is larger than the inverse fragments passing time through the resonant surface, resulting in the two-staged thermal quench as described in Section 5.1. To see this, we look at the Poincaré plot of the magnetic field during a JET-like injection as shown in Fig. 13. Here, Fig. 13(a) corresponds to the time just before the first stage TQ at  $t = 2.16ms$ , (b) represents that just after the first stage TQ at  $t = 2.56ms$  and (c) shows that just before the second stage TQ at  $t = 3.28ms$ . The stochastic region is seen to propagate ahead of the fragments as a result of the larger amplitude of the outer modes such as the 2/1 and 3/2 during the first stage TQ, as opposed to the fast, equatorial injection shown in Fig. 6 where the fragments are always leading the stochastic region. The second stage TQ is still triggered by the fragments arriving at the  $q = 1$  surface.

This premature loss of thermal energy may be undesirable for TQ heat load mitigation, as the poloidal density relaxation is not yet complete as can be seen in Fig. 10(a), and the heat flux from the first stage quench may “leak” out, increasing the heat load on plasma facing components. It is also not likely to be desirable for runaway mitigation as the temperature in the central part of the plasma starts decreasing well before significant density increase. An artificially fast JET-like SPI with speed  $500 \pm 100m/s$ , but otherwise the same injection parameters with the standard case has been carried out for comparison, and it is numerically observed that only a single TQ occurs for the fast injection case. The final penetration at the end of the TQ shows no order of magnitude difference between those two cases, with the fast injection having slightly better penetration as can be seen in Fig. 14.

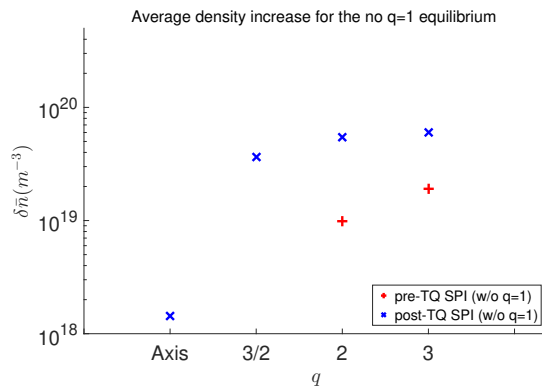
However, the difference in the injection speed does have an impact on the time difference between the core temperature relaxation and the core density relaxation. The density profile at the beginning of the TQ, at the end of the TQ and at the time of core density relaxation for the fast JET-like injection, as well as their corresponding



**Figure 15.** The density evolution for the fast JET-like SPI at (a) the onset of the TQ ( $t = 1.33 \text{ms}$ ), (b) the end of the TQ ( $t = 2.09 \text{ms}$ ) and (c) the completion of core density mixing ( $t = 2.87 \text{ms}$ ). The corresponding temperature profile of (a)-(d) are shown in (d)-(f).

temperature profiles, are shown in Fig. 15. Comparing with the standard JET-like injection in Section 5.1, the fast injection shows a smaller time difference (less than  $1 \text{ms}$ ) between the core density and temperature relaxation. This is potentially due to the step mode structure of the  $1/1$  kink, which means only injected atoms within and very close to the  $q = 1$  surface got transported into the plasma core. Hence more density are convectively transported at the early phase of the TQ for the fast injection than the standard injection simply due to more fragments arrive at the  $q = 1$  surface for the former case, facilitating the quicker relaxation. With this regard, it suggests that a faster injection velocity is beneficial for a more spontaneous core density and temperature relaxation, which might be desirable to reduce runaway electron generation.

It is also important to note that the two timescales that really matter here are the timescale of the mode growth and the timescale of the fragments passing. And since we have used an artificially large resistivity in our simulation, the linear and nonlinear growth rates are increased accordingly. Hence in a real plasma with realistic resistivity, the “threshold speed” of SPI for this two-stage thermal quench may be lower than it is shown in this section.



**Figure 16.** The injection penetration for the equilibrium with  $q(0) > 1$ . There is little change of density within the  $q = 2$  surface before the TQ so no point is shown within  $q = 2$  for the pre-TQ case.

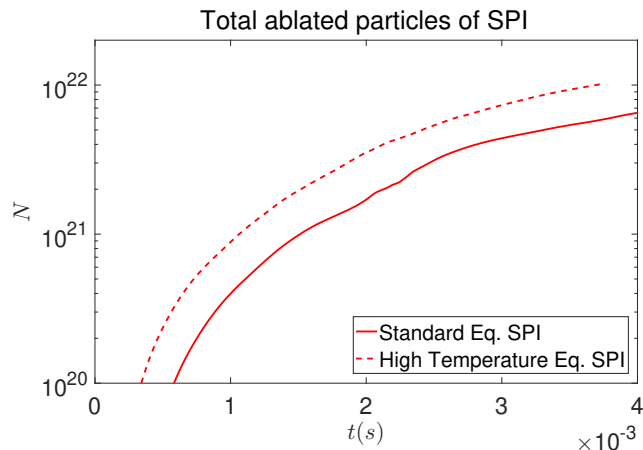
#### 5.4. JET-like injection into equilibria without $q = 1$ surface

The above investigated JET-like injections indicate that the 1/1 internal kink plays a major role in the MHD core mixing after the SPI, and consequentially is a major factor in the injection penetration since the trajectories of the fragments do not cross the magnetic axis. Hence, it is of interest to investigate the scenario where the  $q = 1$  surface, thus the 1/1 mode, is absent. For this, we consider an equilibrium with increased toroidal magnetic field but otherwise the same density, temperature and toroidal current density profiles compared to the one introduced in Section 2.2, so that the axis safety factor  $q(0) = 1.12$  and the boundary safety factor  $q_{95} = 3.39$ .

Due to the absence of the  $q = 1$  surface, the growth of the 2/1 mode dominates the pre-TQ dynamics of the plasma, and it triggers the other mode numbers when it grows large enough, and no strong perpendicular convection of density is observed during the TQ. Instead, the density spread during the TQ is rather localized around the  $q = 2$  surface, resulting in poor core penetration. The average density increase within the  $q = 3/2$ ,  $q = 2$  and  $q = 3$  surfaces for the equilibrium with  $q(0) > 1$  can be seen in Fig. 16. Compared to the penetration in the standard case, as shown in Fig. 14, the mixing is poor. This confirms the important role of the 1/1 mode for core mixing in the case of off-axis SPI.

#### 5.5. JET-like injection into a higher temperature equilibrium

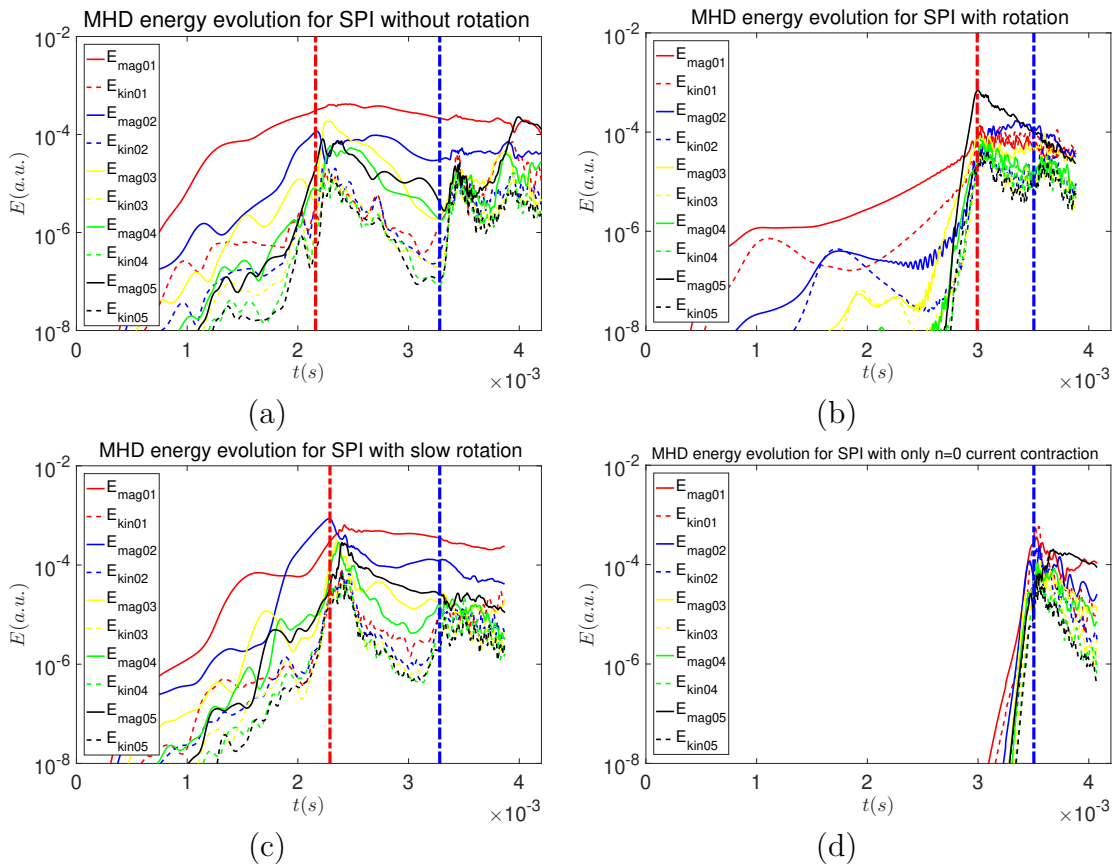
Another interesting issue to access is how the injection assimilation scales with the target equilibrium temperature, as this may provide some information on the extrapolation to higher performance plasmas. This extrapolation is limited, however, by the thermal equilibration time of the hot plasma electrons. In a 1keV plasma as we investigated above, the electron thermal equilibration time with a density of  $10^{20}/m^3$  in the cold, newly ionized electron cloud is about  $\tau_{eq} \simeq 4 \times 10^{-6}s$  [24]. This is much shorter than the timescale of fragments crossing flux tubes which is on the order of  $10^{-4}s$ . This



**Figure 17.** The comparison of the total assimilated particles for JET-like SPI into the standard and high  $T_e$  equilibria. The higher temperature is seen to be beneficial for the assimilation, though the power dependence of total assimilation on the initial temperature is less than one.

indicates that our previous assumption of Maxwellian electrons is valid. On the other hand,  $\tau_{eq}$  increases with  $T_e^{3/2}$ , so that this timescale becomes comparable with that of the fragments crossing as the electron temperature approaches  $10keV$ . This will result in a much higher ablation rate since the incoming heat flux “seen” by the fragments is carried by the still hot background electrons with longer mean free path, rather than the thermalized ones. This enhanced ablation is self-regulated, however, as  $\tau_{eq}$  decreases with  $n_e$ . So that the increased ablation will cause the electrons to equilibrate faster.

In this section, we look at an equilibrium with twice the electron temperature compared to the standard case we introduced in Section 2.2. The toroidal current profile and the density profile remain unchanged. The core temperature is  $T_e(0) = 2.5keV$ , the central safety factor  $q(0) = 0.943$  and the edge safety factor  $q_{95} = 2.697$ . Naively, if the plasma temperature was constant, which would correspond to the no thermalization limit, the ablation rate would have the simple scaling  $\partial_t N \propto T_e(t=0)^{1.64}$  according to Eq. (11). Once thermalization happens, however, this power dependence will not hold. Moreover, once the plasma is cooled down to  $T_e \sim \mathcal{O}(100eV)$ , the mean free path of the plasma electrons is reduced to  $\lambda_e \simeq \mathcal{O}(0.1m)$ , so that the core fragments in a fragments’ cloud are effectively shielded from the background plasma by the periphery fragments, reducing any further ablation. The comparison of total assimilation between the high temperature equilibrium and the standard equilibrium for the same SPI configuration is shown in Fig.17. It can be seen that the increased temperature is indeed increasing the total assimilation, although the power dependence of the latter on the former is less than unity.



**Figure 18.** The comparison of the perturbed magnetic and kinetic energy of  $n = 1 - 5$  modes for the JET-like SPI in (a) a stationary plasma, (b) a rigid body rotating plasma with rotation frequency  $10\text{kHz}$ , (c) a rigid body rotating plasma with rotation frequency  $1\text{kHz}$  and (d) a plasma where the  $n = 0$  current contraction is the only MHD destabilizing mechanism. The red and blue chained lines indicate the onset time of the first and the second TQ.

### 5.6. JET-like injection into a rotating plasma

Intuitively, the MHD destabilization mechanism detailed in Section 3 would be compromised in a fast toroidally rotating plasma, as the toroidal rotation will tend to average out the helical cooling effect and result in a more uniformly cooled plasma. In principle, such an effect would begin to manifest when the toroidal rotation frequency is comparable with the inverse timescale of fragments crossing the flux tubes, which can be approximated by  $\tau_c \simeq r_{ng}/v_p$ , with  $r_{ng}$  being the radius of the neutral cloud and  $v_p$  being the fragment velocity. For  $v_p \simeq 200\text{m/s}$  and  $r_{ng} \simeq 2\text{cm}$ , the crossing timescale would be  $\tau_c \simeq 10^{-4}\text{s}$ , corresponding to a  $10\text{kHz}$  rotation frequency.

On the other hand, however, once the islands opened up, ablation of the fragments will more easily increase the density within the island than outside of the island, thus the X-point cooling is small compared to the in-island cooling, causing a net helical cooling structure. This picture is similar to that of the unmodulated ECCD island stabilization, where the unmodulated current drive in the O-point dominates that in

the X-point, leading to a net stabilizing effect [23]. Thus, we would expect the toroidal rotation to suppress mode excitation before island formation, but to lose the stabilizing effect in the presence of a significant island. Furthermore, as the fragments ablate, the toroidal rotation will decrease significantly due to momentum conservation, thus removing this stabilization mechanism once substantial density increase occurs.

To show this, we investigate SPI cases for which a rigid body plasma rotation is mimicked by artificially moving the fragments along the toroidal direction with rotation frequency  $f = 10kHz$  and  $f = 1kHz$  respectively. As a comparison, we also look into a case where the resistivity and neutral source are artificially set to be toroidally symmetric so that the  $n = 0$  current contraction is the only MHD destabilizing mechanism. The other parameters are the same as for the standard JET-like injection, and the impact of toroidal rotation on the equilibrium has been neglected. The comparison of the MHD response with the non-rotating case is shown in Fig. 18. It can be seen that for the  $f = 10kHz$  rotating case, before the onset of the TQ, the rotational average has significant impact for the  $n > 1$  modes as their amplitude is significantly lower compare with the non-rotation case. Beginning from  $t = 2.4ms$  however, significant growth of  $n > 1$  modes occurs which ultimately leads to the first TQ at  $t = 3.0ms$ . This is then followed by the final collapse of the core temperature at  $t = 3.50ms$ . Since the two TQ are very close to each other, they are nearly indistinguishable on Fig. 18(b). For the slowly rotating case, the MHD spectrum looks more or less the same compared with the non-rotating one, and the time at which the first and second TQ occur is also similar, suggesting the toroidal rotation is not important once it's slow enough. For the  $n = 0$  current contraction only case, the TQ is triggered after a significant amount of current is contracted into the  $q = 1$  surface at  $t = 3.50ms$ .

Comparing Fig. 18(b) and (d), it is evident that the second TQ of the fast rotating case happens at the same time as the TQ in the  $n = 0$  current contraction case, suggesting the current contraction plays an important role in the final collapse of core temperature. Meanwhile, the helical effect is suppressed during the early part of the injection and only manifests itself very close to the final TQ, when the islands begin to form. Thus, it can be said that once the toroidal rotation is faster than the inverse time scale of fragments crossing the flux tube, then the rotation will indeed suppress the helical destabilizing effect, but as Fig. 18(c) has shown, once the rotation has slowed down, it is not important anymore. As a further note, here we are considering injection into an initially MHD free rotating plasma, injection into a rotating plasma with existing modes requires further investigation in the future.

## 6. Summary and Conclusion

The MHD instabilities and the density response during deuterium SPI into a JET L-mode plasma has been investigated in this work. The main focus has been on the macroscopic current driven modes that are responsible for triggering the thermal quench and for the convective density mixing. The evolution of the plasma following injection

of deuterium by SPI and MGI has been investigated using the 3-D non-linear reduced MHD code JOREK, combined with the strongly shielded NGS model to describe the fragments ablation.

It is found that the MHD destabilization by the deuterium SPI is dominated by the the local helical cooling and current perturbation instead of the  $n = 0$  current contraction. This helical effect is driven by the almost adiabatic local decrease of the electron temperature near rational surfaces due to the pellet ablation/ionization. Hence, the SPI fragments destabilize successive layers of rational surfaces as they fly into the plasma core, result in broad-spectrum MHD perturbation and widespread field line stochasticity in the wake of the fragments.

It is also found that the SPI shows superior penetration after the TQ compared with the MGI even when the fragment trajectories do not cross the magnetic axis, as the former deposit the injected material right into the  $q = 1$  surface before and during the onset of the TQ, where the 1/1 kink convection is most efficient. This results in sufficient core density mixing. As a consequence, the SPI enjoys a shorter time difference between the core temperature and density relaxation compared with its MGI counterpart, although some localized regions with cold, hollow density and good flux surfaces still exist, which may be vulnerable to runaway electron formation.

Further investigations reveals several important injection parameters which impact on the penetration and assimilation of injected deuterium. The shattering fineness of the SPI is found to increase the total assimilation of injected atoms, so long as the fragments can still be considered drag-less. Increased injection velocity can both improve the heat load mitigation by prevent premature triggering of the TQ, and shorten the time difference between the core density and temperature relaxation due to more efficient density convection by the 1/1 kink. Varying the spreading angle or the velocity dispersion has no significant impact.

Last, but not least, the property of the target equilibrium is shown to be crucial for achieving better injection efficiency. For instance, higher target plasma temperature results in increased assimilation. Meanwhile, off-axis deuterium SPI into an equilibrium with  $q(0) > 1$  shows no strong global density convection, thus poor core penetration. This, however, does not imply off-axis SPI would be ineffective for the ITER advanced scenario, as its  $q$  profile and temperature profile is different from the JET case we investigated here. The detailed study for both deuterium and impurity SPI into such a scenario is left for future work. Finally, fast toroidal plasma rotation shows suppression of helical cooling before islands formation, though such stabilizing effect becomes insignificant once the plasma has slowed down.

Although the above studies are carried out for a JET plasma, they also contribute to the future ITER DMS design, since the basic MHD processes are expected to remain the same. Disruption mitigation in ITER will rely on the injection of radiating impurities such as neon and argon. Inclusion of such impurity species in JOREK as well as SPI simulation into high temperature H-mode are therefore of high priority and will be done in the near future.

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